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1 **Validation of spaceborne precipitation radar**
2 **data by rain gauges and disdrometers over**
3 **the complex topography of the northeastern**
4 **Indian subcontinent**

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Abstract

39

40 Near-surface rain rate datasets derived from the Tropical Rainfall Mea-
41 suring Mission Precipitation Radar (TRMM PR) and Global Precipitation
42 Measurement Dual-frequency Precipitation Radar (GPM DPR) and near-
43 surface raindrop size distribution (DSD) parameters derived from the GPM
44 DPR were validated using 43 tipping-bucket rain gauges installed over the
45 northeastern Indian subcontinent and two Parsivel² disdrometers installed
46 on the Meghalaya Plateau, India. Both TRMM PR version 7 and version 8
47 products significantly underestimated the rainfall over the Indian subconti-
48 nent during the monsoon season (June–September). The GPM DPR version
49 06A product also significantly underestimated the rainfall at stations on the
50 Meghalaya Plateau, India. The heavy rainfall area (HRA) of the Megha-
51 laya Plateau in the TRMM PR climatology showed lighter rainfall on the
52 plateau, whereas heavier rainfall was detected in adjacent valleys. Intense
53 surface rainfall over the HRA may be detectable, because such intense rain-
54 falls tended to occur from deeper convections, that were less affected by
55 the ground clutter interferences. A comparison of the statistical features of
56 the DSD parameters between the disdrometers and GPM DPR retrievals
57 around the Meghalaya Plateau confirmed that an adequate assumption of
58 the adjustment factor ϵ is important for improving the DSD parameters in
59 GPM DPR retrievals.

60 **Keywords** validation of spaceborne precipitation radar; orographic rainfall;
61 Indian monsoon; raindrop size distribution

62 1. Introduction

63 The Tropical Rainfall Measuring Mission (TRMM) and its successor, the
64 Global Precipitation Measurement (GPM), orbit the Earth between 35°S
65 and 35°N and between 65°S and 65°N, respectively. The onboard precipita-
66 tion radar (TRMM PR) (Kummerow et al. 1998; Kozu et al. 2001; Iguchi
67 et al. 2000, 2009) and dual-frequency precipitation radar (GPM DPR) (Ko-
68 jima et al. 2012; Hou et al. 2014; Skofronick-Jackson et al. 2017) have
69 provided information on three-dimensional rainfall distributions and aided
70 advanced precipitation research on a global-scale since their launches in 1997
71 and 2014, respectively (e.g., Houze et al. 2015). These radars have also con-
72 tributed to global rain rate distribution datasets, such as the Global Satel-
73 lite Mapping of Precipitation (GSMaP) (Kubota et al. 2020) and Integrated
74 Multi-satellite Retrievals for GPM (IMERG) (Huffman et al. 2020), which
75 use spaceborne microwave sensors with enhanced temporal resolutions. The
76 GPM DPR enables estimations of the mass-weighted mean drop diame-
77 ter (D_{mass}) of the precipitation drop size distribution (DSD) (Skofronick-
78 Jackson et al. 2017). The normalized gamma DSD (Willis 1984; Testud
79 et al. 2001) and the relationship between rain rate (R) and the D_{mass} were

80 adopted in the GPM DPR algorithm (Seto et al. 2013b). The assumed
81 normalized gamma DSD has two additional parameters: the normalized
82 intercept parameter, N_w , and the shape parameter, μ , which has a value
83 equal to 3. The accuracy of the DPR retrieved D_{mass} has been proved
84 to some extent (e.g., D’Adderio et al. 2018; Chase et al. 2020; Gatlin
85 et al. 2020) by a validation with dual-polarization radar data and ground-
86 based disdrometers. Recently, Liao and Meneghini (2022) reported that the
87 range-independent assumption of the adjustable parameter in the R - D_{mass}
88 relation degraded the accuracy of the R and D_{mass} estimation. The spatial
89 distribution of GPM DPR-retrieved D_{mass} values have been described by
90 Radhakrishna et al. (2020) and Yamaji et al. (2020).

91 Many efforts have been made to validate spaceborne radars that pass
92 over specific locations using data from ground-based rain gauges (e.g., Ami-
93 tai et al. 2012; Seto et al. 2013b) and ground-based radars (e.g., Wolff
94 et al. 2005; Wolff and Fisher 2008; Amitai et al. 2009; Petracca et al.
95 2018; Watters et al. 2018; Petersen et al. 2020). Additionally, many meth-
96 ods have been proposed to improve spaceborne radar estimates (e.g., Ma
97 et al. 2020; Arulraj and Barros 2019, 2021; Hirose et al. 2021). Rain gauges
98 are thought to provide the most reliable measurements (Kidd et al. 2017).
99 However, comparative studies of instantaneous TRMM PR near-surface rain
100 rate (NSR) data with in-situ rain gauge networks are limited. One of the

101 reasons for this is the difficulty to obtain rain gauge data with high time res-
102 olution; generally, the time resolution of operational rainfall data is coarser
103 than 10 min. In developing countries, this situation is even more severe. For
104 example, the Bangladesh Meteorological Department operationally observes
105 rainfall every 3 h. Therefore, rainfall datasets with high time resolutions
106 are valuable for the validation of spaceborne radar rain retrievals.

107 The TRMM PR has revealed large spatial gradients of precipitation
108 over complex terrains worldwide (e.g., Nesbitt and Anders 2009; Hirose
109 et al. 2017). However, the estimation of precipitation over complex topog-
110 raphy using spaceborne radars contains various errors. For example, ground
111 clutter over the complex terrain raises the clutter free bottom level, mak-
112 ing it difficult to detect shallow rain and increasing reflectivity toward the
113 ground, which tends to occur in orographic seeder-feeder clouds (e.g., Prat
114 and Barros 2010; Speirs et al. 2017; Arulraj and Barros 2021; Shimizu et al.
115 2023). The influence of ground clutter contamination into path-integrated
116 attenuation deteriorates the estimation of near-surface rainfall. In addition,
117 the non-uniform beam filling caused by complex topography results in poor
118 path-integrated attenuation estimates, which are due to the degraded qual-
119 ity of the reference dataset of normalized radar cross section used in the
120 surface reference technique (Meneghini et al. 2015). A better understand-
121 ing of the characteristics of precipitation over complex terrain is needed to

122 improve the reliability of space-borne radar retrievals.

123 This study aimed to validate surface rainfall with TRMM PR V7, V8,
124 and GPM DPR V06A products using a tipping-bucket rain gauge network
125 encompassing the northeastern Indian subcontinent based on the method
126 proposed by Terao et al. (2017). The difference in product versions, as
127 well as the difference between instruments used in TRMM PR and GPM
128 Ku-band PR, regulate the performance of precipitation estimation. Further-
129 more, we attempted to validate the GPM DPR-retrieved DSD parameters
130 using in situ disdrometers located on the Meghalaya Plateau. The study
131 area has a complex topography including the rainfall station "Cherrapunji",
132 which is located on the southern slope of the Meghalaya Plateau and is
133 reported to have the heaviest rainfall in the world (Jennings 1950; Mu-
134 rata et al. 2017). The characteristics are different between premonsoon
135 and monsoon seasons, while diurnal and intraseasonal variations are domi-
136 nant in this region. Such large spatiotemporal variation in rainfall over the
137 northeastern Indian subcontinent represents a unique testbed to validate
138 rainfall products. The remaining sections of this paper are organized as
139 follows. Section 2 provides details on the datasets and methodology used.
140 The validation results are presented in Section 3, and Section 4 presents
141 the properties of rainfall over the heavy rainfall area (HRA) in Meghalaya.
142 Section 5 discusses the validation of DSD parameters and the contrast in

143 rainfall between the plateau and the valleys over the HRA. Finally, Section
144 6 provides a summary.

145 **2. Data and analysis method**

146 *2.1 Rain gauges*

Fig. 1

147 We installed 43 tipping-bucket type rain gauges manufactured by Ikeda
148 Keiki (Shizuoka, Japan) and Dynalab Weathertech (Maharashtra, India)
149 in Bangladesh and the Assam and Meghalaya Indian states, respectively
150 (Fig. 1). The rain gauges had a resolution of 0.5 mm. Most stations were
151 installed between 2004 and 2006. In 2014, all stations in India were re-
152 placed with the same 0.5-mm tipping-bucket rain gauges manufactured by
153 the Komatsu Factory Co., Ltd (Tokyo, Japan). In 2016, we installed four
154 additional rain gauges in Meghalaya. Two were installed in the grid where
155 heavier rainfall was observed by the TRMM PR (see Fig. 2). The analysis
156 period varies depending on the rain gauge site (details are provided in Sup-
157 plement 1–4). For the data logger, we utilized the HOBO Pendant Event
158 Logger (UA-003-64, Onset Computer Corporation, Bourne, MA, USA). All
159 loggers recorded the timing of tipping with a second resolution. On average,
160 we visited each station once or twice a year to download the accumulated
161 data and adjust the data loggers' clock. However, a maximum clock devia-

162 tion of several minutes may have still occurred due to logger clock drift.

163 The rain gauges of our network were matched with spaceborne radar
164 beams whose centers were within the matching radius D according to the
165 procedure adopted by Terao et al. (2017). When multiple beams were ob-
166 tained, each beam was treated as an independent event, assuming that the
167 information from each satellite footprint was considered as an independent
168 sample for comparison. In the study by Terao et al. (2017) the average of
169 multiple beams was calculated and counted as one event. In the present
170 study, we modified this portion of the method used by Terao et al. (2017)
171 to increase the number of samples. The rain rate from rain gauge data
172 was calculated from the number of tipping within the time window between
173 $t + \tau - \Delta t$ and $t + \tau + \Delta t$. Here, t is the spaceborne radar scanned timing,
174 Δt is the half length of the time window to count tipping, and τ is the esti-
175 mated time lag during which spaceborne radar-observed precipitation falls
176 and reaches the rain gauge on the ground (Amitai et al. 2012; Seto et al.
177 2013b; Terao et al. 2017). Terao et al. (2017) examined several sets of these
178 parameters by calculating the correlation coefficient with the spaceborne
179 radar matchups, confirming the best robustness and representativeness for
180 $\tau = 300$ s, $D = 3.5$ km, and $\Delta t = 150$ s.

181 We applied the percentile method of the bootstrap test (Efron 1979) with
182 a Monte Carlo algorithm to calculate the confidence interval for ensemble

183 averages for the ΔR error. This is the difference between the spaceborne
184 radar NSR of the matched pixel (SAT) and the rain rate of rain gauges dur-
185 ing the time window of $2\Delta t$ (RG), expressed as $\Delta R = SAT - RG$. In total,
186 10,000 resampling averages were calculated for the original observations of
187 ΔR to obtain the 2.5 and 97.5 percentiles, which defined the 95% confi-
188 dence intervals. This test was performed only when more than 20 non-zero
189 samples were available.

Fig. 2

190 Figure 2 shows the TRMM PR V7 climatological rainfall map over the
191 southern slope of the Meghalaya Plateau based on TRMM PR data from
192 1998–2013 (Hirose and Okada 2018); elevation contours are also included.
193 The comparison between precipitation distribution and elevation contours
194 revealed that the HRA was distributed over a narrow west-east elongated
195 area, which corresponds to a steep slope area between 500- and 1500-m
196 contours in the southern Meghalaya Plateau. Figure 2 also shows that the
197 rainfall in the HRA is heavier in the valley and lighter on the plateau.
198 However, the rainfall stations renowned for heavy rainfall, Cherrapunji and
199 Mawsynram, and a comparable rainfall station, Pynursla, are located on
200 the plateau.

Fig. 3

201 In May 2017 we installed a new rain gauge at Cherrapunji beside the
202 disdrometer in the India Meteorological Department (IMD) observatory to
203 validate the disdrometer. Figure 3a shows a comparison between the daily

204 rainfall measured by the rain gauge in the IMD observatory and that mea-
205 sured by our rain gauge network in the Cherrapunji station. The IMD
206 observatory is located approximately 1 km east of our Cherrapunji station,
207 closer to the edge of the plateau. Figure 3b compares the daily rainfall
208 measured by the rain gauge at the IMD observatory and that measured by
209 the rain gauge station at Sohkhme, 6 km southeast of the IMD site. The
210 Sohkhme village is located in a valley within the heavier rainfall grid of the
211 TRMM climatological rain map (Fig. 2). We confirmed a high correlation
212 between the neighboring gauges in Figs. 3a and b. The obtained regression
213 coefficients show that rainfall at the IMD site was higher than that measured
214 at Cherrapunji with our rain gauge network and that at Sohkhme. The av-
215 erage rainfall during the simultaneous observation period was higher at the
216 rain gauge located on the plateau than that in the valley. The bootstrap
217 test, which examined the difference in the averages of artificially resampled
218 data, showed that the rainfall deficit between the data collected at Sohkhme
219 and the IMD observatory was nearly statistically significant with a 90% con-
220 fidence interval. The causes for the observed differences in rainfall between
221 the plateau and valley are discussed in Sections 4.3 and 5.

222 2.2 Disdrometers

223 We utilized a second-generation laser optical OTT PARTicle SIze and
224 VELOCITY (Parsivel²) disdrometer (Tokay et al. 2014). This device simul-
225 taneously measures the fall speed and size of precipitation particles. The
226 smallest observable diameter was 0.312 mm. We first conducted a quality
227 check of the disdrometer data as follows. First, data with a bad sensor status
228 > 1 were excluded. The periods within and after heavy rainfall sometimes
229 result in bad sensor status or missing data. Kalina et al. (2014) considered
230 three sources of error: strong wind effects, raindrops falling within the mar-
231 gin of the observation area, and splash. Raindrops with a fall speed 60%
232 faster or slower than the empirical fall speed–diameter relationship (Gunn
233 and Kinzer 1949; Atlas et al. 1973) were eliminated to avoid these errors,
234 though the number of eliminated data was small. Finally, DSDs with more
235 than 100 raindrops were used to avoid the distortion associated with the
236 estimation of DSD shapes (Smith and Liu 1993; Smith 2016).

237 The modeled DSD in normalized gamma form (Willis 1984; Testud et al.
238 2001; Bringi et al. 2002) has three parameters, namely, N_w , D_{mass} , and μ .
239 N_w is the normalized intercept parameter and represents the intercept of an
240 equivalent exponential DSD with the same liquid water content and mass-
241 weighted mean diameter D_{mass} as the gamma DSD (Testud et al. 2001); μ is
242 a shape parameter. Throughout this paper, the main unit of N_w is decibels,

243 which equals $10\log_{10}N_w$, while the original unit of N_w is $\text{m}^{-3}\text{mm}^{-1}$. A
244 comparison with a spaceborne radar was conducted using the same rain-rate
245 validation method (Section 2.1). Parameters D , Δt , and τ were assigned
246 values of 3.5 km, 180 s, and 300 s, respectively. The selection of Δt changed
247 from 150 to 180 s because the time resolution of disdrometers was 1 min.
248 Thus, a time window spanning 2–8 min after the passage of the satellite
249 was utilized for the comparison.

250 The 1 min rain samples were observed using two disdrometers installed
251 in the Cherrapunji IMD (91.734°E , 25.269°N) and on the rooftop of the
252 building of the Department of Geography, North–Eastern Hill University
253 (91.896°E , 25.610°N) in Shillong (Fig. 2) from May 2017 to March 2020. A
254 tipping-bucket rain gauge was installed at each station, and the number of
255 tippings was recorded every 1 min to validate the disdrometer-derived rain-
256 fall. We found that the disdrometer systematically underestimated hourly
257 rainfall at Cherrapunji, by approximately 30% (Murata et al. 2020).

258 2.3 *Spaceborne radars*

259 This study primarily employed the TRMM PR V8 data from June 2004
260 to March 2014 and the GPM DPR V6A from March 2014 to March 2020.
261 Both datasets use the same retrieval algorithms. TRMM PR V7 was used
262 to confirm the effect of a longer analysis period compared with the results

263 of Terao et al. (2017). The dataset for the dual-frequency algorithm (DPR
 264 Level-2 product, DPRL2 hereafter) was applied using measurements from ei-
 265 ther the Ku-band, Ka-band, or both when available (Seto et al. 2013a; Seto
 266 and Iguchi 2015; Seto et al. 2021). The DPR algorithm assumes that the
 267 DSD follows a normalized gamma form, as described in Section 2.2, where
 268 μ is set to 3 and D_{mass} and N_w are obtained from GPM DPR observations.
 269 The DPRL2 algorithm uses the relationship between rain rate R (mm h^{-1})
 270 and D_{mass} (mm), as presented in the following equations for stratiform and
 271 convective rain types, respectively:

$$R = 0.392\epsilon^{4.815} D_{\text{mass}}^{6.131} \quad \text{for stratiform rain}$$

272

$$R = 1.348\epsilon^{4.373} D_{\text{mass}}^{5.418} \quad \text{for convective rain}$$

273 where ϵ is the adjustment factor. Note that the R - D_{mass} relation used
 274 in the GPM DPR V6 is different from both the GPM DPR V5 and V7.
 275 The precipitation classification method for spaceborne radars is described
 276 by Awaka et al. (2021). Stratiform rain is mainly defined by the detection of
 277 bright bands, while convective rain includes not only precipitation with large
 278 radar reflectivity but also shallow convections. In the DPRL2 algorithm,
 279 the dual-frequency surface reference technique (Meneghini et al. 2015) and
 280 radar reflectivity factor of the Ka-band precipitation radar were used to
 281 adjust ϵ (Seto et al. 2021). We validated D_{mass} and N_w at the clutter-free

282 bottom (CFB) level using ground-based rain gauges and disdrometers. The
283 estimation of the CFB level was different between TRMM PR V7 and V8,
284 and the CFB level in V8 was further raised up when the contamination with
285 the sidelobe clutter occurred.

286 3. Validation results

287 3.1 *Rainfall*

288 We compared the rainfall matchups between rain gauge data and the
289 NSR of the TRMM PR V7, TRMM PR V8, and GPM DPR products over
290 Meghalaya, Meghalaya/new, Assam, Sylhet+Barak, and Bengal Plain areas
291 during the monsoon (June–September) (Table 1) and premonsoon (March–
292 May) (Table 2) seasons. The area classifications of the rain gauge station
293 are shown in Fig. 1.

294 As presented in Table 1, both the TRMM PR V7 and V8 datasets
295 significantly underestimated rainfall with 99% confidence intervals for all
296 four areas during the monsoon season. However, the latter showed rela-
297 tive improvement over Meghalaya and the Sylhet+Barak areas, which are
298 influenced by orographic rainfall. In contrast, the degree of underestima-
299 tion increased in the plain areas of Assam and Bengal. GPM DPR V6A
300 showed a significant underestimation for only the Meghalaya area; notably,

Table 1

Table 2

301 the new stations installed in Meghalaya (Meghalaya/new) showed overes-
302 timated rainfall, although the difference was not statistically significant.
303 Table 2 indicated that TRMM PR V7 significantly overestimated rainfall
304 during the premonsoon season in the Assam and Bengal plains, whereas
305 TRMM PR V8 did not have this issue. Seto (2022) compared the precip-
306 itation rate estimates between the TRMM PR V8 and GPM Ku-band PR
307 (KuPR) Version 6, confirming that the precipitation rate estimate of the
308 TRMM PR exceeded that of the GPM KuPR counterpart. The authors
309 attributed this to a larger value of the adjustment factor ϵ related to the
310 adjustment of the attenuation correction with GPM KuPR. Furthermore,
311 there were no significant differences in rainfall among the areas during the
312 premonsoon season for the GPM DPR V6A product (Table 2c).

313 *3.2 DSD parameters*

Table 3

314 Table 3 shows contingency tables for rainfall detection between disdrom-
315 eters at Cherrapunji (Tables 3a–c) and Shillong (Tables 3d–f), and the GPM
316 DPR NSR matchups during all periods (Tables 3a and d), monsoon season
317 (Tables 3b and e), and premonsoon season (Tables 3c and f). The per-
318 centage of GPM DPR misdetection was high at Cherrapunji throughout
319 the year, and the probability of detection (POD) (e.g., Kidd et al. 2012)
320 was 53% (Table 3a). Meanwhile, misdetection was high at Shillong, only

321 during the monsoon season, and the POD value was 53% (Table 3e). The
 322 misdetection may also be caused by the difference between the representa-
 323 tive spatiotemporal scales of the two measurements. This is because the
 324 disdrometer continuously measures at the same position, while the GPM
 325 DPR measures the instantaneous return signal from the beam coverage
 326 area, which is a circle with a 5.2-km diameter at the nadir. The maximum,
 327 mean, and median values of the rain rate observed by disdrometers for the
 328 DPR misdetection events were 3.93, 0.74, and 0.16 mm h⁻¹ at Cherrapunji,
 329 and 0.65, 0.15, and 0.08 mm h⁻¹ at Shillong, respectively. The result shows
 330 the GPM DPR misses light rains.

Fig. 4

331 Figure 4 shows scatter plot comparisons of the DSD parameters between
 332 the disdrometers and GPM DPR retrievals for both rainy samples, distin-
 333 guished as stratiform and convective types according to the GPM DPR
 334 algorithm (Awaka et al. 2021). Most samples were stratiform in Shillong,
 335 whereas half were convective at Cherrapunji. Colored marks indicate the
 336 rain rates of the samples corresponding to the spaceborne radar NSR. Al-
 337 though the relationship between D_{mass} and rain rate is used in the GPM
 338 DPR algorithm, it is unclear in Figs. 4a and b implying a contribution of
 339 the adjustment factor ϵ in the algorithm. The D_{mass} of convective rain has
 340 both small and large values because both deep convections and shallow rains
 341 are classified as convective rain. D_{mass} showed better correspondence be-

342 tween disdrometers and GPM DPR retrievals, and the mean absolute error
343 was <0.5 mm; however, several outliers were included, which deteriorate
344 the correlation coefficient. The correlation coefficient of the N_w at Cherra-
345 punji (Fig. 4d) was higher than that at Shillong (Fig. 4c), while the GPM
346 DPR retrievals corresponded rather well with the disdrometer counterpart
347 at Cherrapunji. The N_w of the GPM DPR retrievals tended to concentrate
348 in the 30–40 dB range (Figs. 4c and d).

Fig. 5

349 Several studies have identified distinct characteristics in the geographic
350 distribution of D_{mass} and its seasonal variation (Yamaji et al. 2020; Rad-
351 hakrishna et al. 2020). Figure 5 shows the spatial distribution of average
352 D_{mass} and average N_w at the CFB level over the Meghalaya Plateau and ad-
353 jacent areas of the Bengal Plain during premonsoon and monsoon seasons.
354 The fluctuation of D_{mass} values during the premonsoon season was more
355 significant than that during the monsoon season, reflecting its small rainy
356 samples and a higher percentage of convective rain (Hirose and Nakamura
357 2002; Islam and Uyeda 2008). During the monsoon season, the value of
358 D_{mass} tended to be small on the Meghalaya Plateau and large in the Bengal
359 Plain south of the Meghalaya Plateau (Fig. 5c). Meanwhile, the N_w values
360 in the plateau area were larger than those in the plain area, with a high N_w
361 distributed over the southern and western slopes of the plateau, including
362 the HRA (Fig. 5d).

Fig. 6

363 The statistical characteristics of the DSD parameters are sometimes
364 represented by $D_{\text{mass}}-N_w$ diagrams (e.g., Bringi et al. 2009; Dolan et al.
365 2018; Arulraj and Barros 2019). The $D_{\text{mass}}-N_w$ diagrams shown in Figs. 6a
366 and d represent GPM DPR retrievals using the sample bins inside the area
367 drawn in Fig. 2a ($90^\circ-93^\circ\text{E}$, $24.5^\circ-26^\circ\text{N}$), while Figs. 6b and e correspond
368 to the disdrometers at Cherrapunji. Figs. 6c and f represent the data from
369 Shillong. The GPM DPR retrievals show the concentration of samples with
370 a D_{mass} of 1.0–1.5 mm and N_w of 30–35 dB (Fig. 6a), with a low quantity
371 of small drops ($D_{\text{mass}} < 1$ mm) are much less. In contrast, the disdrome-
372 ters show the concentration of samples with $D_{\text{mass}} < 1$ mm and $N_w \geq 45$
373 dB at Cherrapunji (Fig. 6b) and $D_{\text{mass}} \approx 1.0$ mm and $N_w \approx 35-40$ dB at
374 Shillong (Figs. 6c). The $D_{\text{mass}}-N_w$ diagram is distinguished by six rain rate
375 categories in Figs. 6d–f. The minimum rain rate was set as 0.2 mm h^{-1} ,
376 which approximately corresponds to the minimum detectable rain rate of
377 the DPR (Skofronick-Jackson et al. 2017). There are differences between
378 the GPM DPR retrievals (Fig. 6d) and the disdrometer results (Figs. 6e and
379 f). For example, the GPM DPR (Fig. 6d), with many large drops ($D_{\text{mass}} \approx$
380 2–3 mm) retrieved even for the light rain rate category with less than 5
381 mm h^{-1} . Moreover, the GPM DPR retrieved large $N_w (> 45 \text{ dB})$ for the
382 heavy rain rate category with more than 50 mm h^{-1} . These features were
383 not observed in the disdrometers (Figs. 6e and f).

384 Liao et al. (2020) found that the gamma DSD model fits the power law
 385 equation $R = a \times N_w \times D_{\text{mass}}^b$, where $a = 1.588 \times 10^{-4}$ and $b = 4.706$,
 386 independent of the shape factor μ . Figures 6g-i show that the GPM DPR
 387 retrievals and the disdrometer data at Cherrapunji and Shillong fit well with
 388 the equation, except at both ends of the line. The accuracy of both small
 389 and large D_{mass} ends may be difficult to discuss because both small and large
 390 drops are susceptible to errors in the Parsivel disdrometer observation (e.g.,
 391 Tokay et al. 2013). However, differences were still observed between re-
 392 trievals and disdrometers. The smallest limit of R/N_w in each rain rate
 393 category increased with D_{mass} with color gradation clearly observed in the
 394 disdrometers (Figs. 6h and i), but it was unclear in the retrievals (Fig. 6g).
 395 This feature is related to a distinct reduction in the upper limit of N_w and
 396 increase in the lower limit of D_{mass} with an increase in the rain rate category
 397 (Figs. 6e and f).

398 4. Properties of rainfall over the HRA

399 4.1 General features

Fig. 7

400 We estimated the contoured frequency by altitude diagram (CFAD) of
 401 stratiform and convective radar reflectivity (Fig. 7a and b) from spaceborne
 402 radars over the HRA and compared them with the profiles in other moun-

tainous areas around the globe (Anders and Nesbitt 2015). The convective profiles showed deep convections where the 0.05% frequency contour crossed 40 dBZ at an altitude of approximately 9 km. Conversely, the stratiform profiles showed higher reflectivity below the melting level at approximately 4.5 km of altitude with the 0.05% frequency contour crossing 40 dBZ at around 5 km. The composite over the HRA was generally very similar to the “tropical regime” such as the Himalayas, New Guinea, and the Peruvian Andes, whereas the composites of the convective profile included many shallow convections similar to those of the “wet monsoon regime”, such as the Western Ghats and Myanmar coast.

Fig. 8

The NSR rainfall distribution of a rare heavy rain case during the TRMM PR overpass, with a rain rate of around 150 mm h^{-1} was simultaneously observed by the rain gauges at Mawsynram and Cherrapunji, while approximately 80 mm h^{-1} was observed at Pynurla (Fig. 8). The HRA was positioned near the edge of the TRMM PR pass. Although the rain intensity of TRMM PR did not match the in-situ rain gauges, the distribution included a very intense rainfall area with rain rates exceeding 100 mm h^{-1} . The intense rainfall area was located over a windward steep slope with a narrow west-east elongated shape, similar to the climatological rainfall distribution (Fig. 2).

423 4.2 *Angle-bin dependence*

Fig. 9

424 Hirose et al. (2021) and Seto et al. (2021) showed that precipitation
425 statistics from the spaceborne radars strongly depend on the scanning an-
426 gle. Figures 9a–c show the average NSR over the HRA using all-angle
427 bins (Fig. 9a), near-nadir bins (Fig. 9b), and off-nadir bins (Fig. 9c) of the
428 TRMM PR V8. Here, near-nadir (off-nadir) data were defined as the angle
429 bin number of 22–28 (1–21 and 29–49), which corresponds to a local zenith
430 angle of $< 2.5^\circ$ ($> 2.5^\circ$). The NSR corresponds to the rain rate at the CFB,
431 so Fig. 9a using TRMM PR V8 may be different from Fig. 2 using TRMM
432 PR V7 because the estimation method of the CFB level has been changed.
433 Nonetheless, the rain rate was more intense in the valley and less intense on
434 the plateau, as also observed in Fig. 2. The contrast became sharp and the
435 rain rate in the valley was strongest at near-nadir bins (Fig. 9b), consistent
436 with the findings of Hirose et al. (2021).

Table 4

437 Table 4 shows the contingency tables for surface rainfall detection be-
438 tween the near-nadir and off-nadir data for spaceborne radar matchups with
439 rain gauges over the HRA. The POD was 86% (66%) and the false alarm
440 ratio (FAR) was 39% (53%) for the near-nadir (off-nadir) data, respectively,
441 which confirms the higher accuracy of near-nadir data. The FAR tends to
442 be larger than the errors observed for the disdrometer data (Table 3), which
443 is because the minimum rain gauge resolution is 0.5 mm. The average rate

444 of rainfall detected by the rain gauges (the spaceborne radars) was 17.6 mm
445 h^{-1} (11.8 mm h^{-1}) for near-nadir data, and 19.5 mm h^{-1} (8.2 mm h^{-1}) for
446 off-nadir data, respectively, and implying that near-nadir data represents
447 the rainfall amount better than off-nadir data.

448 The CFB level itself has angle bin dependence (Hirose et al. 2021).
449 Figures 9d–f show the horizontal distributions of CFB thickness, which cor-
450 responds to the distance between the ground and CFB level, for all-angle
451 bins (Fig. 9d), near-nadir bins (Fig. 9e), and off-nadir bins (Fig. 9f). The
452 CFB thickness is large over the steep slope area, with a maximum average
453 value of approximately 1.7 km. The CFB thickness generally decreased in
454 near-nadir bins (Fig. 9e) and increased in off-nadir bins (Fig. 9f), although
455 the degree of change in CFB thickness was rather small in the steep slope
456 area. The CFB thickness of near-nadir bins decreased to less than 1.0 km
457 over the Bengal Plain and on the Meghalaya Plateau; however, it was 1.5
458 km over the steep slope area.

459 4.3 *Difference in rainfall between plateau and valley*

Fig. 10

460 The $0.05^\circ \times 0.05^\circ$ grid on the plateau where Cherrapunji is located was
461 labeled grid-A, while that in the valley where Sohkhme is located was la-
462 beled grid-B (Fig. 9a). Figures 10a and b show the rain rate profiles for
463 near-nadir bins in grid-A and grid-B, respectively. Only near-nadir data

464 was used because the performance of the retrievals was better than that for
465 the off-nadir data (Table 4). The red-colored portion represents the profiles
466 between the ground and CFB level. The rain rate below the CFB level are
467 blind owing to ground clutter, so they were retrieved by regarding radar
468 reflectivity as the same as that in the CFB level. The slight decrease in
469 rainfall intensity was a result of considering the denser air near the ground
470 and the slow fall speed rate of raindrops. The ground level and CFB thick-
471 ness in grid-A are rather uniform, while various ground levels from near sea
472 level to plateau level and CFB thickness of more than 1.5 km are observed
473 in grid-B. A greater number of heavy NSRs with more than 10 mm h^{-1} was
474 observed in grid-B, but most of the NSRs in both grid-A and grid-B were
475 less than 10 mm h^{-1} . Interestingly, heavy NSRs tended to have higher rain
476 rates up to higher altitudes far above the maximum altitude of the plateau.
477 Notably, some heavy NSRs in grid-B were more intense and rapidly in-
478 creased downward toward the ground. Figure 10c shows the average rain
479 rate profiles for grid-A (black line) and B (blue line). The average rain rate
480 of grid-B was larger than that of grid-A below an altitude of 6 km. The
481 profile of grid-A was nearly constant below 5 km, whereas that of grid-B
482 increases downward, and the rain rate doubled at the 2 km level.

483 Hamada and Takayabu (2014) reported the presence of suspicious ex-
484 treme rainfall in the TRMM PR V7 product, mostly over the land. They

485 showed that most suspicious extremes have a significant monotonic increase
486 in radar reflectivity toward the echo bottom and isolated extreme NSR
487 with large differences from the surrounding pixels. Some profiles in grid-B
488 (Fig. 10b) show similar characteristics to the suspicious extreme rainfall.
489 However, they were not removed using the filter proposed by Hamada and
490 Takayabu (2014). Only one of the profiles had the ratio of the NSR to the
491 average NSRs in the four surrounding pixels exceeding 300, but the vertical
492 gradient of the two lowest bins was smaller than 20 dB km^{-1} .

Fig. 11

493 Figure 11 shows 76 vertical profiles of radar reflectivity within the HRA
494 at pixels that matched rain gauges observed more than 30 mm h^{-1} . If the Z -
495 R relationship ($Z [\text{mm}^6 \text{ m}^{-3}] = 124 \times R [\text{mm h}^{-1}]^{1.50}$) is adopted based on the
496 disdrometer observation at Cherrapunji during May–October 2017 (Murata
497 et al. 2020), then 30 mm h^{-1} corresponds to 43 dBZ. Most profiles increased
498 toward the ground below the melting layer, suggesting the dominance of the
499 collisional growth of rain drops or the seeder-feeder process. In addition,
500 the storm-top height (STH) was below 10 km in altitude for the 86% cases,
501 confirming that heavy rain does not necessarily have a tall STH (Hamada
502 et al. 2015). However, intense rain rate cases of more than 80 mm h^{-1}
503 have a comparably higher STH ($\geq 8 \text{ km}$) and stronger radar reflectivity
504 throughout the profile ($\geq 45 \text{ dBZ}$ below the melting layer), confirming that
505 the heavy NSR tends to have higher rain rates extending to higher altitudes

506 (Fig. 10).

507 **5. Discussion**

508 Here, we discuss the validation of GPM DPR-retrieved DSD parame-
509 ters with two disdrometers in the Meghalaya Plateau. We also discuss the
510 distinct contrast in TRMM PR climatology over the HRA, which features
511 heavier rainfall in the valley and lighter rainfall on the ridge.

512 The average values of the spaceborne radar during the monsoon season
513 (Fig. 5) tended to have a relatively smaller D_{mass} and larger N_w over the
514 Meghalaya Plateau than those over the plain area in the southern plateau.
515 This is reasonable because the disdrometers exhibited many samples with
516 small D_{mass} and large N_w (Fig. 6). This result is also consistent with
517 the characteristics of DSD in orographic rains (e.g., Rosenfeld and Ulbrich
518 2003). However, samples with small $D_{\text{mass}} < 1$ mm and large $N_w > 45$
519 dB were rare in the GPM DPR retrievals (Figs. 6a and b). The satellite-
520 retrieved D_{mass} and N_w were concentrated in the 1.0–1.5 mm and 30–40 dB
521 ranges, respectively. Gatlin et al. (2020) also reported a severely limited
522 range of N_w estimates at approximately 35 dB.

523 Both GPM DPR retrievals and disdrometers fit well along the line pro-
524 posed by Liao et al. (2020), implying strong constraints among D_{mass} , N_w ,
525 and rain rate. Although the relationship between D_{mass} and rain rate is

526 utilized as the basis of the DSD parameter retrievals (Seto et al. 2021),
527 the correlation between D_{mass} and rain rate was weak in Fig. 5, possibly
528 because an adjustment factor ϵ substantially decides the D_{mass} value. Some
529 outliers in Figs. 5a and b show large D_{mass} (≈ 2 mm) for weak rain rate
530 (< 5 mm h⁻¹), while the disdrometer observations (Figs. 6b–c, e–f) feature
531 these outlier values substantially less.

532 The disdrometer results (Figs. 6h and i) showed that the lower limits of
533 D_{mass} and R/N_w increase with R , where R is the rain rate. This indicates
534 that the minimum value of D_{mass} (the maximum value of N_w increasing
535 (decreasing) with the rain rate is a principal characteristics of the disdrom-
536 eter results (Figs. 6e and f), which coincides with other observation results
537 (e.g., Fig. 5 of Tokay et al. 2020). The color gradation in the GPM DPR
538 retrievals was unclear (Figs. 6d and g) and corresponded to the upper-right
539 portion of Fig. 6d, which shows where no data was found in the disdrom-
540 eter results (Figs. 6e and f). Updating the DSD database in GPM DPR
541 V7 and changing the algorithm from a range-independent ϵ assumption in
542 this validated GPM DPR V6 dataset to a range-variable ϵ model in GPM
543 DPR V7 (Liao and Meneghini 2022) may improve the DSD parameters of
544 the GPM DPR retrievals.

545 A comparison of the rain rate profiles of near-nadir data between the
546 grid in the valley and that on the plateau (Fig. 10) revealed an intense

547 surface rain rate more frequently in the valley (grid-B). Some intense rain
548 rate profiles and the average profiles in grid-B still imply a possibility of
549 ground clutter contamination in the valley profiles. The high CFB in the
550 valley becomes an obstacle in detecting shallow precipitation in the blind
551 zone below the CFB level (Shimizu et al. 2023). However, the intense
552 NSR profiles tended to have different properties from those of weak profiles
553 in that they have stronger reflectivity up to far above the ground level
554 (Figs. 10 and 11). This suggests that intense NSRs have less influence on
555 the blind zones, implying that the rainfall distribution over the HRA with
556 heavier rainfall in the valley may not be artificial. However, there are still
557 other factors that cause errors in precipitation retrievals from spaceborne
558 radars. For example, the surface backscattering cross-section of spaceborne
559 radars over land also increases in the presence of precipitation, degrading
560 precipitation retrievals by estimating the path-integrated attenuation in the
561 surface reference technique procedure (Seto et al. 2022). Figure 3a showed
562 the daily rainfall at the IMD observatory in Cherrapunji was systematically
563 heavier than that at our Cherrapunji station, which was located further
564 apart from the valley. It also implies that heavier rainfalls are produced
565 in valleys. The newly installed rain gauges in 2016 in the valley (Fig. 3b
566 and Meghalaya/new in Tables 1 and 2) showed rather lower rainfall in the
567 valley. This suggests that there may be differences in rainfall in the valley

568 between rain at the CFB level and that on the ground owing to the effects of
569 environmental fields, such as wind. Further research is necessary to derive
570 stronger conclusions.

571 The TRMM LIS analysis showed that the frequency of thunder is very
572 severe over the southern Meghalaya Plateau (Dewan et al. 2018), which
573 supports the frequent occurrence of deep convection over the area. Ahmed
574 et al. (2022) conducted numerical modeling to simulate a heavy rainfall
575 case at Mawsynram and produced deep convection that strengthened over
576 the upslope region of the Meghalaya Plateau. The increased horizontal
577 resolution in the simulation led to steeper slopes, which resulted in heavier
578 precipitation in the upslope region. Medina et al. (2003) analyzed intensive
579 observations over the Southern Alps and showed the formation of graupels
580 over the steep slope of the Alps.

581 **6. Summary**

582 In this study, we attempted to validate the rain rate retrieved from
583 TRMM PR V7 and V8 and GPM DPR V6A, and DSD parameters retrieved
584 from GPM DPR V6A with tipping-bucket rain gauges over the northeastern
585 Indian subcontinent and disdrometers installed in the Meghalaya Plateau.
586 We also discussed the features in TRMM PR climatological rainfall distri-
587 bution that show lighter rainfall on the plateau and heavier rainfall in the

588 adjacent valleys.

589 The extension of the analysis period of validation based on Terao et
590 al. (2017) supported the underestimation of monsoon precipitation over the
591 northeastern Indian subcontinent and a significant overestimation of pre-
592 monsoon precipitation over the Assam and Bengal plains by TRMM PR
593 V7. A significant underestimation of monsoon precipitation was also ob-
594 served in the TRMM PR V8; however, a significant underestimation of
595 monsoon precipitation was observed only over Meghalaya in the GPM DPR
596 V6A data.

597 The statistical features of D_{mass} and N_{w} derived from the disdrometers
598 at Cherrapunji and Shillong were compared with those of the GPM DPR-
599 retrieved D_{mass} and N_{w} around the Meghalaya Plateau. Both disdrometers
600 showed a dominance of rainfall with a large N_{w} and small D_{mass} , which is
601 a feature of orographic rainfall, while the GPM DPR retrieved D_{mass} and
602 N_{w} showed a limited range of variation in comparison. The disdrometer
603 observation fitted well the line proposed by Liao et al. (2020), which is
604 the same as the GPM DPR-retrievals, implying a strong constraint among
605 D_{mass} , N_{w} , and rain rate. The disdrometer results showed that the minimum
606 value of D_{mass} (the maximum value of N_{w}) increases (decreases) with rain
607 rate. As the relationship between D_{mass} and rain rate is used in the retrieval
608 algorithm, the adequate range of adjustment factor ϵ in the relationship is

609 important for improving DSD parameter retrievals. Better assumptions
610 of DSD parameters in the GPM DPR algorithm will greatly develop the
611 understanding of precipitation over the world, because the characteristics
612 of DSD reflect differences in precipitation mechanisms.

613 TRMM PR climatological rainfall distributions in the monsoon season
614 showed a distinct dependence on topography over the HRA in Meghalaya,
615 with higher rainfall in the valley and lower rainfall on the ridge. Rainfall
616 over complex terrains has various factors that deteriorate the quality of
617 spaceborne radar rain retrieval. This study suggests that the heavy rain
618 over the HRA tended to occur owing to deeper convections and may be
619 less affected by ground clutter blind zones. Such heavy rains were more
620 frequent in the valley than on the plateau. Further observations (e.g., via
621 in-situ weather radars and more detailed analyses) are required to gener-
622 ate conclusions regarding the mechanisms underlying heavy rainfalls over
623 the Meghalaya Plateau. The improved schemes and new parameters in the
624 DPR Version 7 algorithm are also expected to contribute in elucidating a
625 more accurate rainfall distribution. In addition, the achieved enhanced un-
626 derstanding of precipitation characteristics in various meteorological and
627 geographical conditions will be useful to improve satellite-borne precipita-
628 tion radar retrievals.

629

630 **Data Availability Statements**

631 The datasets generated and/or analyzed in this study are available from the
632 corresponding author on reasonable request, subject to all authors' permis-
633 sion.

634 **Supplement**

635 Supplements 1–4 show details of rain gauge stations used for the validation
636 with TRMM PR in Meghalaya and Assam areas, and in Sylhet+Barak and
637 Bengal Plain, respectively. Supplements 3 and 4 show details of rain gauge
638 stations used for the validation with GPM DPR in Meghalaya and Assam
639 areas, and in Sylhet+barak and Bengal Plain, respectively.

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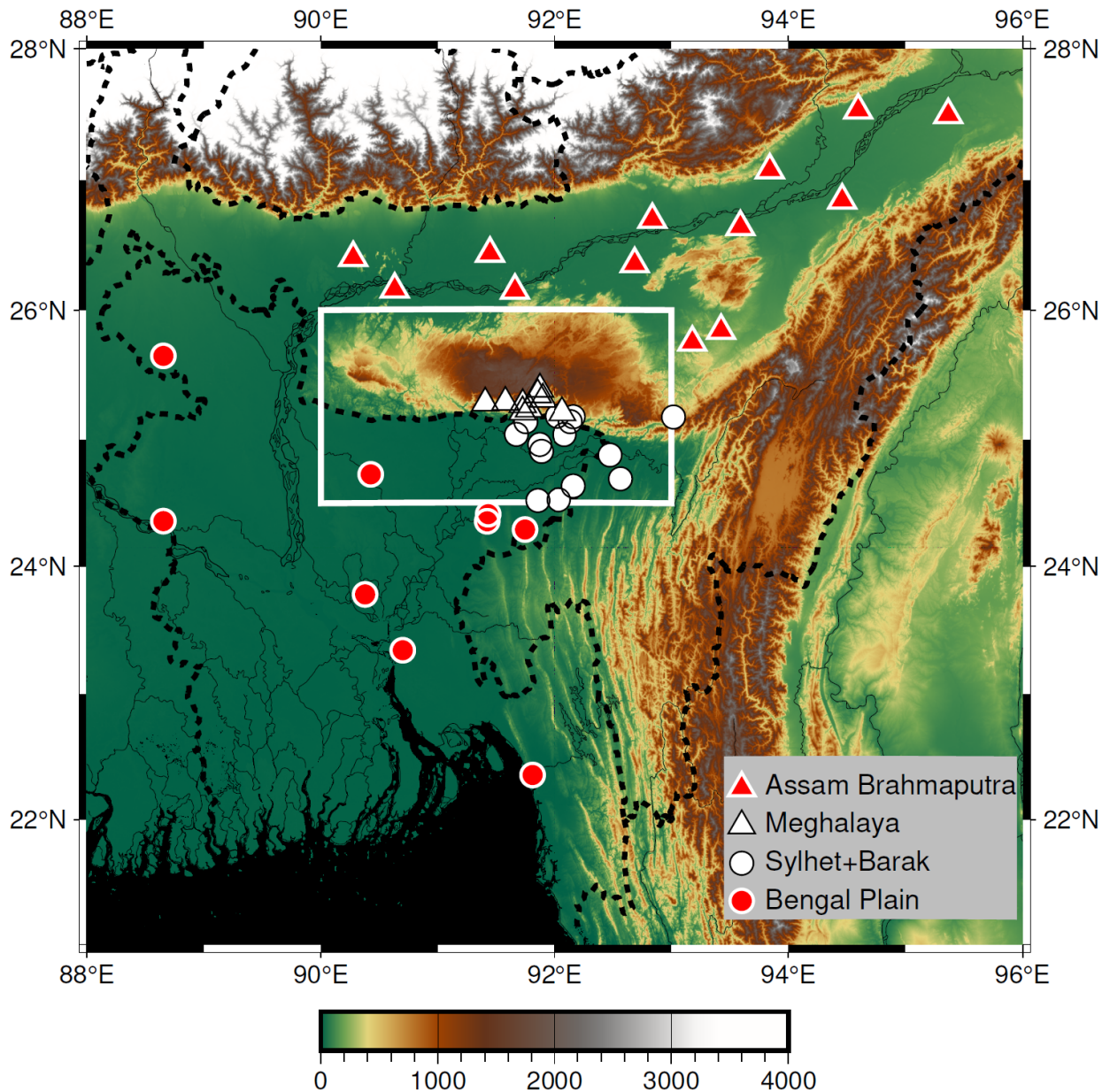


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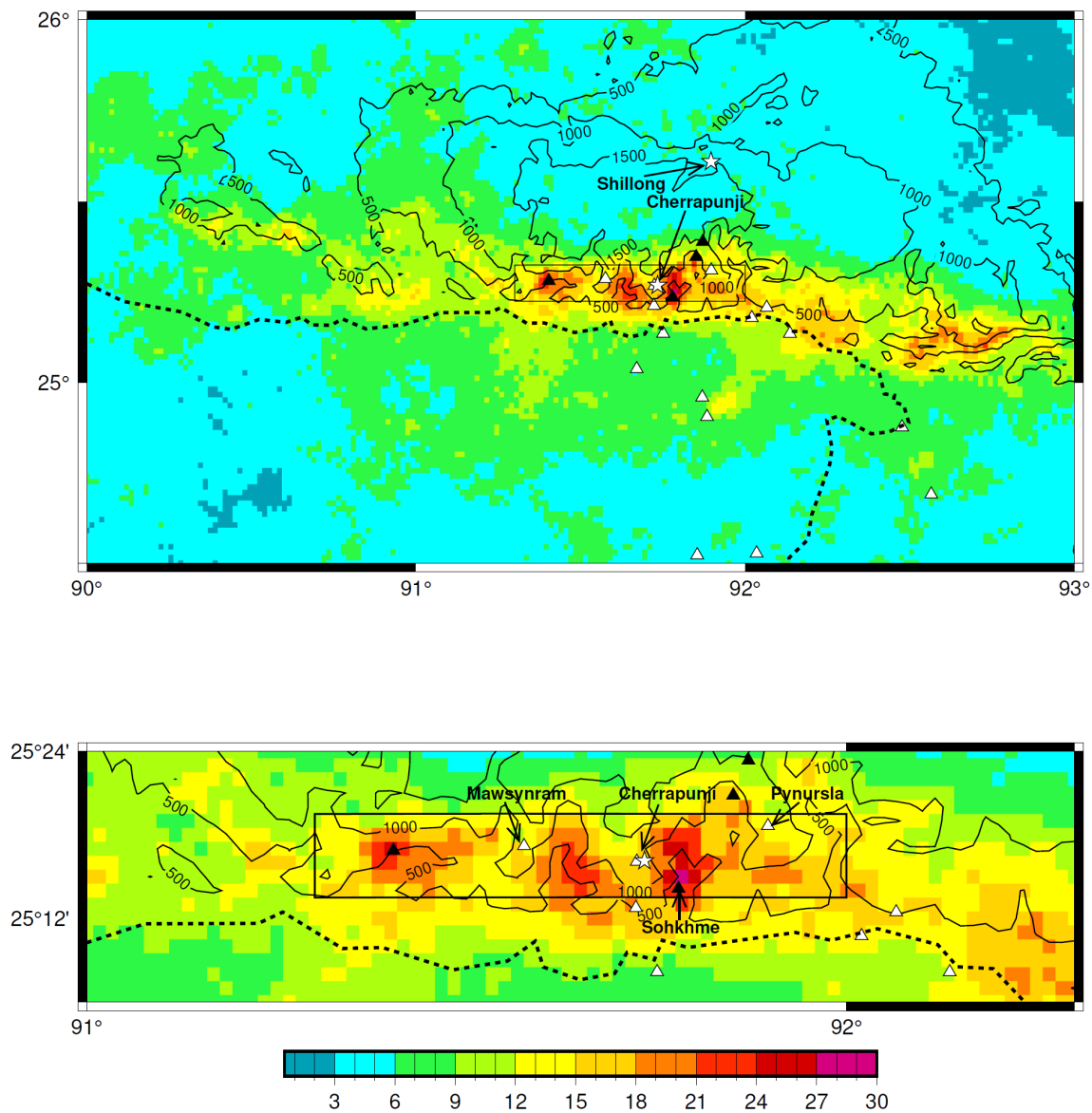


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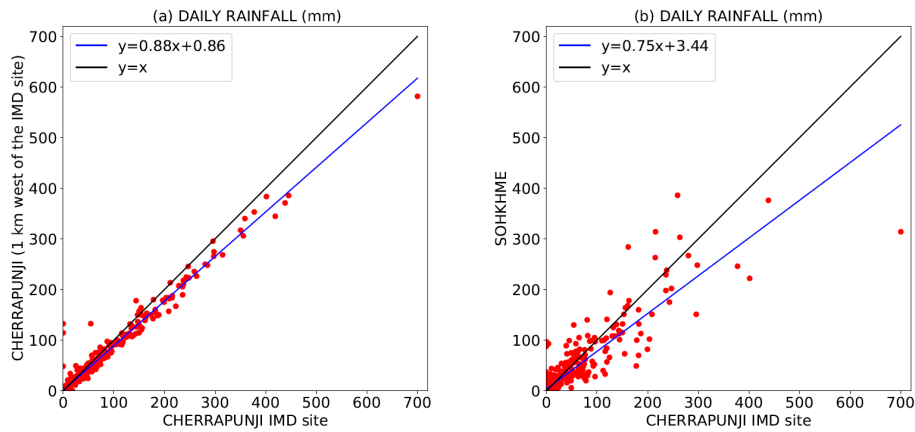


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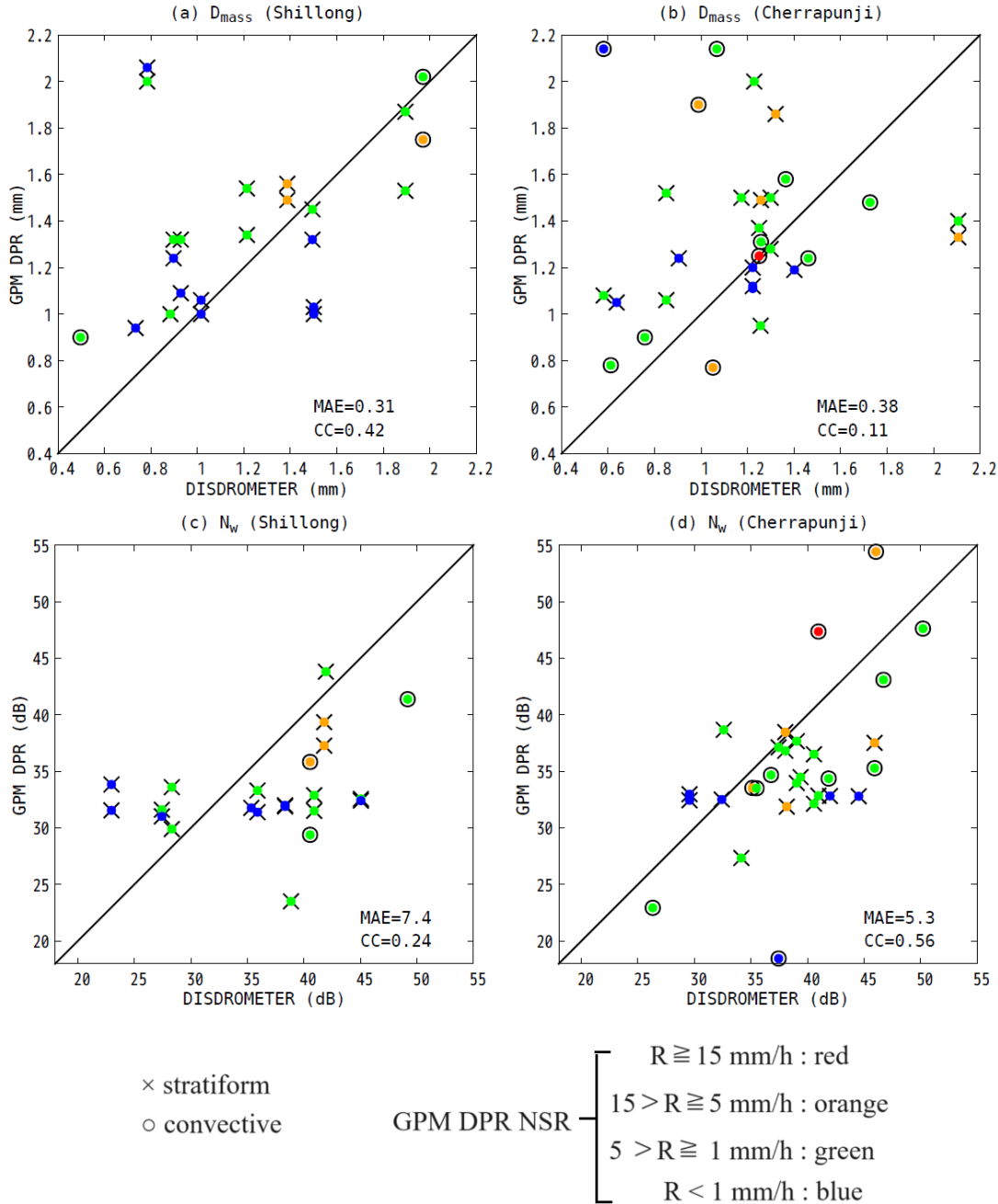


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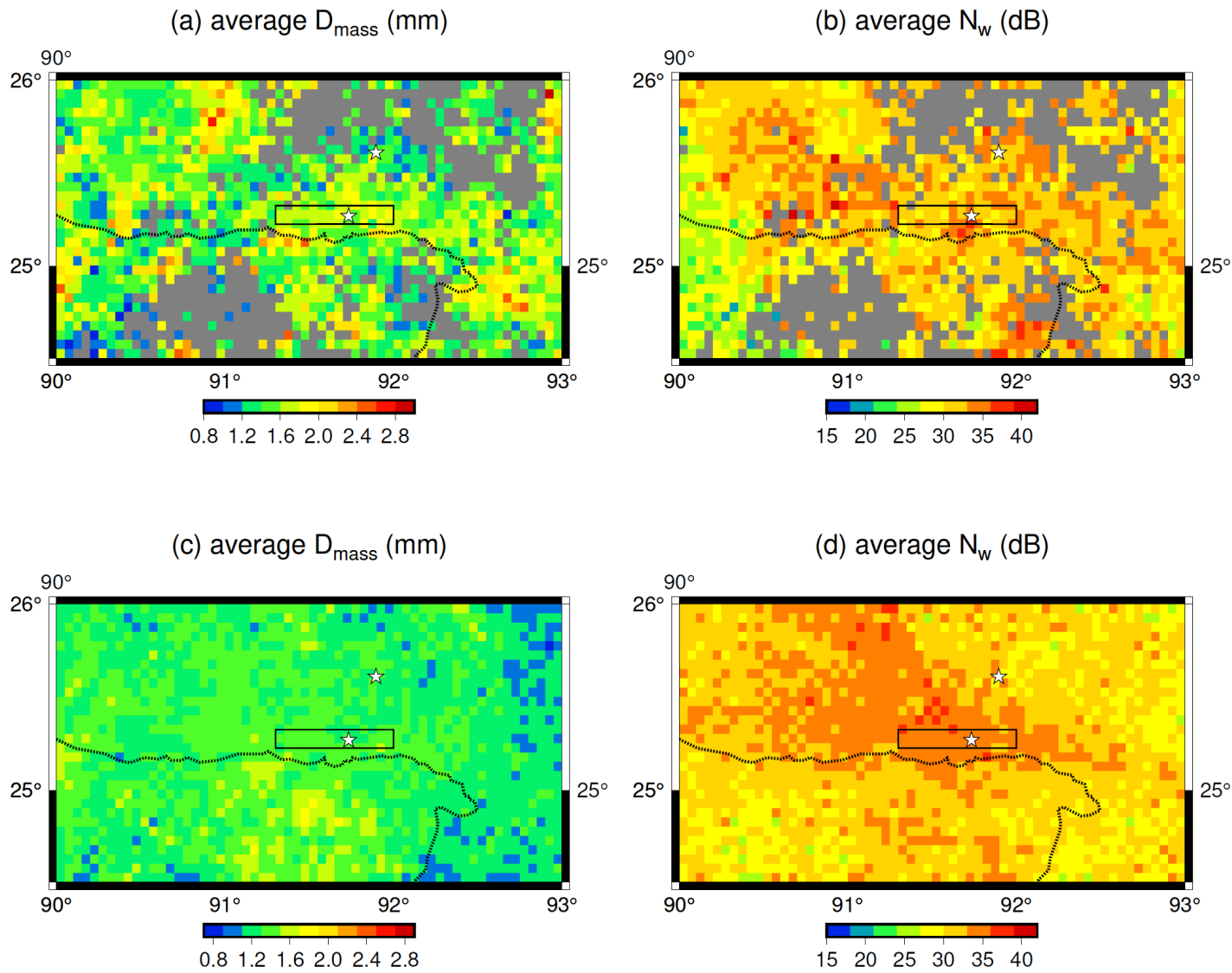


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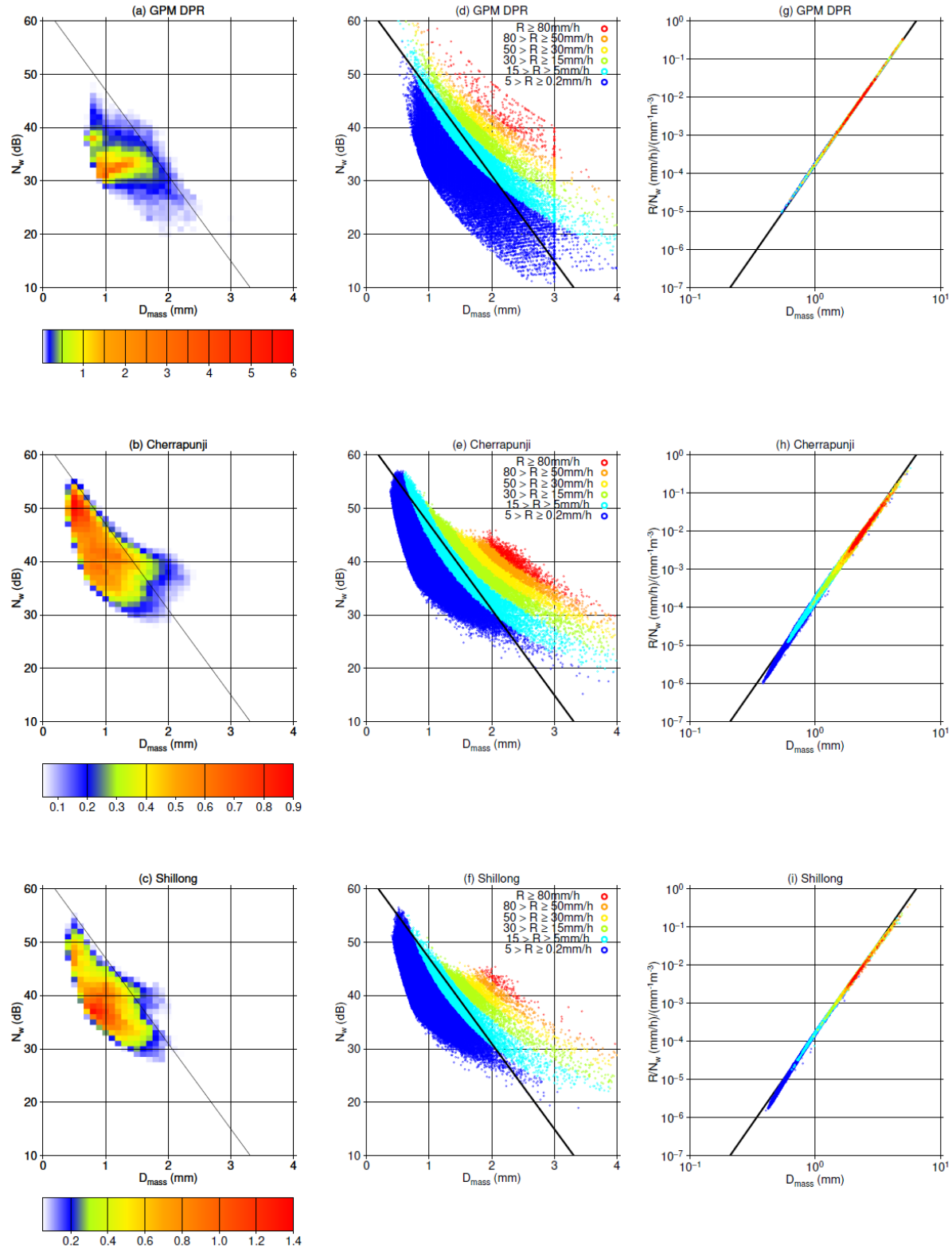


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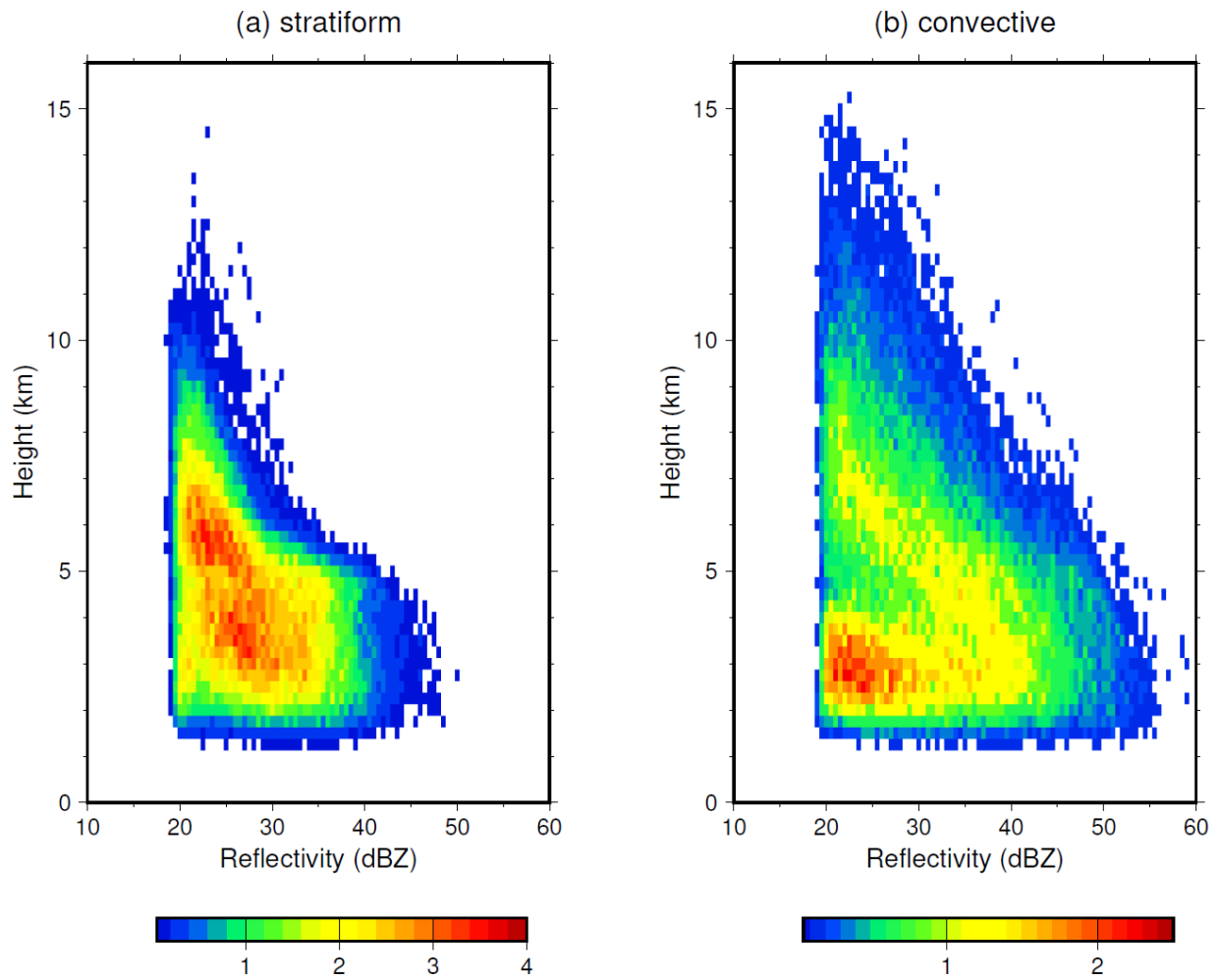


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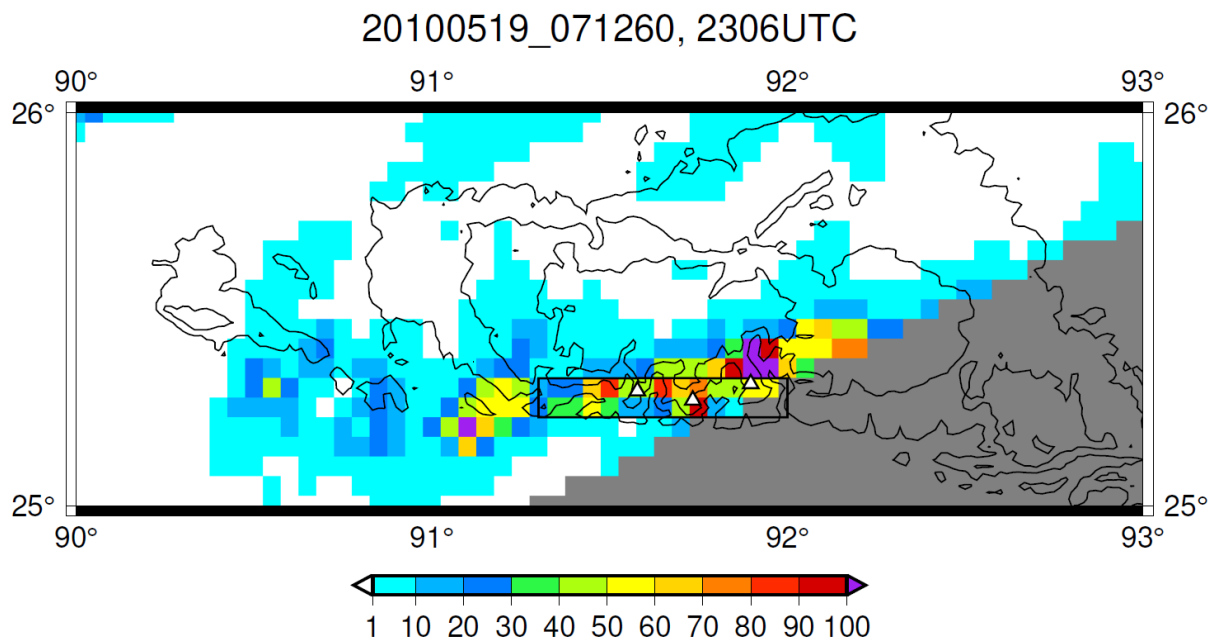


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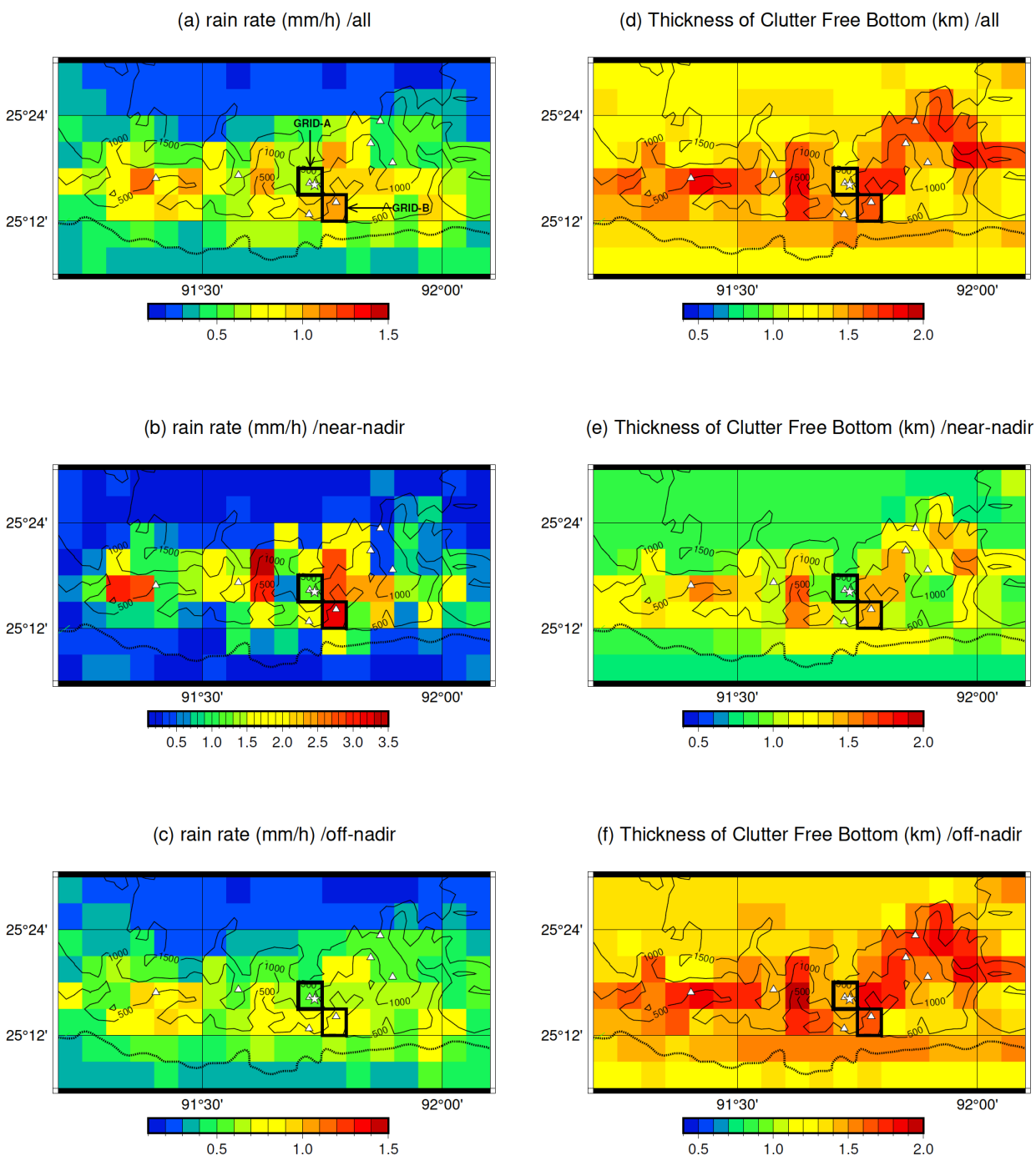


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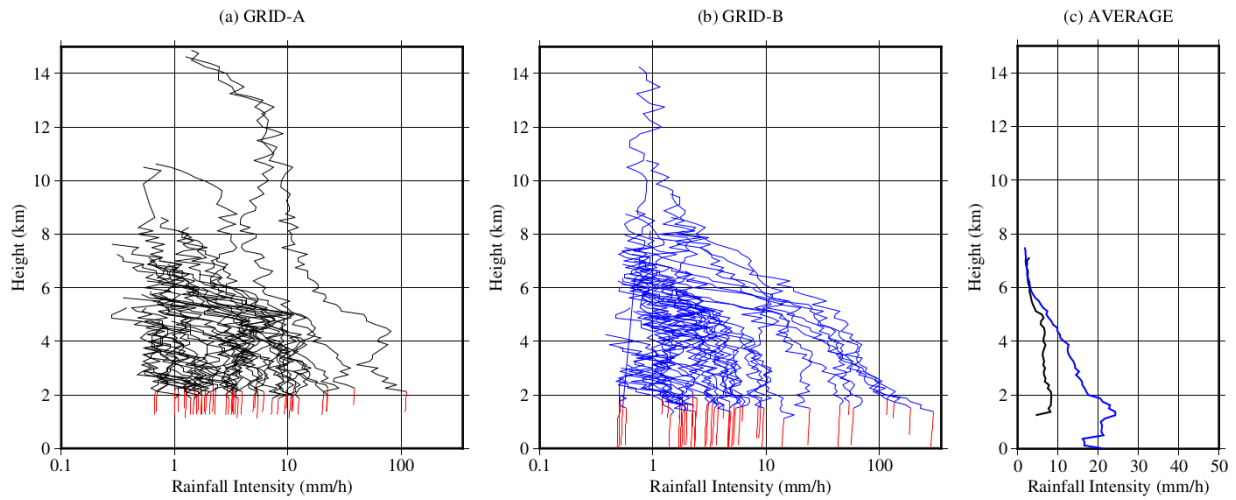


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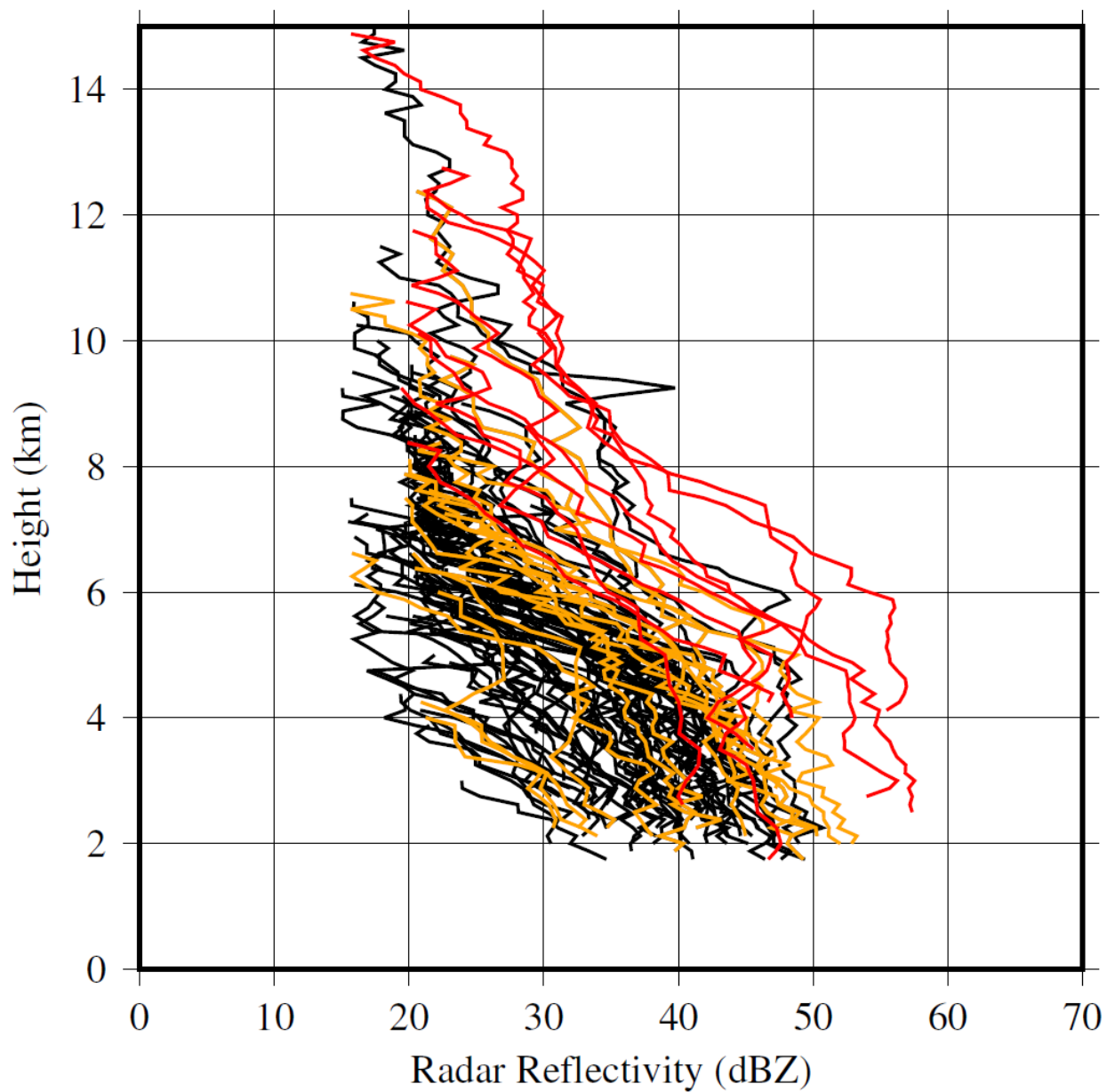


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(a) Area	\overline{RG} (mm h ⁻¹)	\overline{SAT} (mm h ⁻¹)	$\overline{SAT} - \overline{RG}$ (mm h ⁻¹)	$Bias(\%)$	N_{obs}	N_{rain}
Meghalaya	2.24	1.18	-1.06	-47**	3849	725
Assam	0.43	0.36	-0.07	-16**	8422	878
Sylhet+Barak	1.03	0.70	-0.33	-32**	7445	1210
Bengal Plain	0.44	0.33	-0.11	-25**	6186	571
(b) Area	\overline{RG} (mm h ⁻¹)	\overline{SAT} (mm h ⁻¹)	$\overline{SAT} - \overline{RG}$ (mm h ⁻¹)	$Bias(\%)$	N_{obs}	N_{rain}
Meghalaya	2.25	1.48	-0.77	-34**	3829	725
Assam	0.42	0.34	-0.08	-20**	8415	930
Sylhet+Barak	1.04	0.73	-0.32	-30**	7444	1262
Bengal Plain	0.45	0.30	-0.15	-34**	6149	595
(c) Area	\overline{RG} (mm h ⁻¹)	\overline{SAT} (mm h ⁻¹)	$\overline{SAT} - \overline{RG}$ (mm h ⁻¹)	$Bias(\%)$	N_{obs}	N_{rain}
Meghalaya	2.17	1.23	-0.93	-43**	1177	275
Meghalaya/new	0.93	1.13	0.20	+21	507	84
Assam	0.46	0.35	-0.11	-24	2893	333
Sylhet+Barak	0.92	0.75	-0.17	-19	1715	296
Bengal Plain	0.43	0.58	0.15	+36	1134	126

Table 2: Validation of (a) TRMM PR V7, (b) TRMM PR V8, and (c) GPM DPR V6A with rain gauges. Same as Table 1 except for the premonsoon season (March–May).

(a) Area	\overline{RG} (mm h ⁻¹)	\overline{SAT} (mm h ⁻¹)	$\overline{SAT} - \overline{RG}$ (mm h ⁻¹)	<i>Bias</i> (%)	N_{obs}	N_{rain}
Meghalaya	0.77	0.70	-0.07	-9	2725	221
Assam	0.16	0.23	+0.07	+45**	5652	318
Sylhet+Barak	0.28	0.34	+0.06	+23	5419	370
Bengal Plain	0.14	0.15	+0.01	+12*	4676	151
(b) Area	\overline{RG} (mm h ⁻¹)	\overline{SAT} (mm h ⁻¹)	$\overline{SAT} - \overline{RG}$ (mm h ⁻¹)	<i>Bias</i> (%)	N_{obs}	N_{rain}
Meghalaya	0.80	0.66	-0.14	-18	2752	223
Assam	0.16	0.16	+0.00	+1	5678	330
Sylhet+Barak	0.27	0.24	-0.03	-11	5454	405
Bengal Plain	0.14	0.10	-0.04	-30	4675	172
(c) Area	\overline{RG} (mm h ⁻¹)	\overline{SAT} (mm h ⁻¹)	$\overline{SAT} - \overline{RG}$ (mm h ⁻¹)	<i>Bias</i> (%)	N_{obs}	N_{rain}
Meghalaya	0.84	0.64	-0.20	-24	908	75
Meghalaya/new	0.29	0.17	-0.12	-43	393	27
Assam	0.20	0.17	-0.03	-15	2162	172
Sylhet+Barak	0.27	0.26	-0.02	-6	1259	94
Bengal Plain	0.43	0.17	-0.26	-60	854	42

Table 3: Number of rain and no rain events observed at (a–c) Cherrapunji and (d–f) Shillong by the disdrometer and the GPM DPR for (a,d) all periods, (b,e) the monsoon season, and (c,f) the premonsoon season.

(a) Cherrapunji: all

sum 360		disdrometer	
		rain	no rain
DPR	rain	30	7
	no rain	27	296

(b) Cherrapunji: monsoon

sum 130		disdrometer	
		rain	no rain
DPR	rain	23	6
	no rain	19	82

(c) Cherrapunji: premonsoon

sum 72		disdrometer	
		rain	no rain
DPR	rain	3	0
	no rain	4	65

(d) Shillong: all

sum 291		disdrometer	
		rain	no rain
DPR	rain	23	6
	no rain	8	254

(e) Shillong: monsoon

sum 99		disdrometer	
		rain	no rain
DPR	rain	9	4
	no rain	8	78

(f) Shillong: premonsoon

sum 67		disdrometer	
		rain	no rain
DPR	rain	10	1
	no rain	0	56

Table 4: Contingency tables of rainfall at rain gauges in the HRA and the matchups of the TRMM PR and GPM DPR for (a) near-nadir and (b) off-nadir data.

sum 1305		rain gauges	
		rain	no rain
radars	rain	96	62
	no rain	16	1131

sum 7603		rain gauges	
		rain	no rain
radars	rain	333	381
	no rain	175	6714

Supplement 1: List of raingauge stations used for the validation with TRMM PR (Meghalaya & Assam)

Area Name	Station Name	Longitude	Latitude	Periods	Num. samples
Meghalaya	Amlarem	92.1202	25.2943	April 11, 2006 – August 05, 2008	744
Meghalaya	Cherrapunji	91.7239	25.2723	April 19, 2006 – October 06, 2014	1918
Meghalaya	Mawsynram	91.5755	25.2869	May 23, 2006 – October 06, 2014	2405
Meghalaya	Nongtalang	92.0654	25.2077	June 20, 2009 – February 25, 2014	1167
Meghalaya	Pynursla	91.8966	25.3109	June 18, 2006 – February 25, 2014	2087
Meghalaya	Thangkharang Park1	91.7237	25.2172	November 16, 2006 – September 30, 2008	596
Meghalaya	Thangkharang Park2	91.7227	25.2128	February 21, 2009 – October 06, 2014	1580
Meghalaya/ new	Mawjng	91.8709	25.3892	-	0
Meghalaya/ new	Nongkenbah	91.4035	25.2814	-	0
Meghalaya/ new	Sohkhme	91.7789	25.2358	-	0
Meghalaya/ new	Wahkhen	91.8510	25.3474	-	0
Assam	Bokakhat	93.5891	26.6409	March 07, 2013 – September 01, 2014	519
Assam	Diphu	93.4248	25.8426	July 02, 2006 – October 07, 2014	866
Assam	Goalpara	90.6306	26.1612	June 26, 2006 – March 03, 2012	1534
Assam	Guwahati	91.6579	26.1526	March 18, 2006 – October 03, 2014	2833
Assam	Kokrajhar	90.2762	26.4044	March 07, 2008 – October 05, 2014	1561
Assam	Lumding	93.1781	25.7513	July 02, 2006 – October 07, 2014	1420
Assam	Moridhal	94.5958	27.5331	July 08, 2006 – October 05, 2014	2498
Assam	Nagaon	92.6858	26.3571	May 27, 2006 – October 07, 2014	2086
Assam	Nalbari	91.4439	26.4365	July 20, 2006 – October 07, 2014	1940
Assam	Sankardev College	93.8404	27.0768	July 09, 2006 – October 07, 2014	2460
Assam	Teok	94.4581	26.8447	March 04, 2008 – October 07, 2014	1506
Assam	Tezpur	92.8372	26.6973	May 27, 2006 – October 07, 2014	2587
Assam	Tinskia	95.3644	27.4980	February 20, 2007 – October 04, 2014	2116

Supplement 2: List of raingauge stations used for the validation with TRMM PR (Sylhet+Barak & Bengal Plain)

Area Name	Station Name	Longitude	Latitude	Periods	Num. samples
Sylhet+Barak	Amarshid	92.4764	24.8768	March 10, 2006 – October 06, 2014	2352
Sylhet+Barak	Bulaganj	91.7506	25.1361	May 12, 2006 – October 06, 2014	2595
Sylhet+Barak	Chhatak	91.6704	25.0379	March 10, 2006 – October 06, 2014	2676
Sylhet+Barak	Haflong	93.0181	25.1723	August 09, 2006 – October 06, 2014	788
Sylhet+Barak	Hailakandi	92.5651	24.6908	August 09, 2006 – October 06, 2014	1261
Sylhet+Barak	Jaflong	92.0198	25.1791	March 08, 2007 – October 06, 2014	2206
Sylhet+Barak	BARI, Jaintapur	92.1359	25.1359	May 11, 2006 – October 06, 2014	2362
Sylhet+Barak	Juri	92.1617	24.6361	March 13, 2006 – March 04, 2007	295
Sylhet+Barak	Kulaura	92.0345	24.5270	March 08, 2007 – October 06, 2014	2041
Sylhet+Barak	Naljuri	92.1617	25.1751	May 11, 2006 – March 02, 2007	252
Sylhet+Barak	Rajnagar	91.8542	24.5222	May 11, 2006 – October 06, 2014	2621
Sylhet+Barak	Sylhet	91.8842	24.9055	May 11, 2006 – October 06, 2014	2157
Sylhet+Barak	Sylhet Airport	91.8694	24.9595	March 11, 2007 – October 06, 2014	2182
Bengal Plain	Chittagong	91.8085	22.3531	August 07, 2004 – October 03, 2014	2903
Bengal Plain	Dhaka	90.3784	23.7799	August 15, 2004 – October 06, 2014	2900
Bengal Plain	Dinajpur	88.6545	25.6466	March 06, 2005 – March 05, 2014	2680
Bengal Plain	BRRI, Habiganj	91.4286	24.4149	March 08, 2007 – October 06, 2014	1704
Bengal Plain	Mymensingh	90.4263	24.7256	August 04 2004 – September 12, 2014	3175
Bengal Plain	Rajshahi	88.6549	24.3606	August 11, 2004 – March 08, 2014	2899
Bengal Plain	Srimangal	91.7438	24.2950	March 08, 2007 – October 06, 2014	2031

Supplement 3: List of rain gauge stations used for the validation with GPM DPR (Meghalaya & Assam)

Area Name	Station Name	Longitude	Latitude	Periods	Num. samples
Meghalaya	Amlarem	92.1202	25.2943	-	0
Meghalaya	Cherrapunji	91.7239	25.2723	May 03, 2014 – July 31, 2019	705
Meghalaya	Mawsynram	91.5755	25.2869	March 09, 2014 – March 05, 2020	825
Meghalaya	Nongtalang	92.0654	25.2077	October 26, 2014 – March 05, 2020	720
Meghalaya	Pynursla	91.8966	25.3109	October 26, 2014 – March 02, 2020	716
Meghalaya	Thangkharang Park1	91.7237	25.2172	-	0
Meghalaya	Thangkharang Park2	91.7227	25.2128	March 09, 2014 – March 05, 2020	690
Meghalaya/ new	Mawjng	91.8709	25.3892	April 20, 2016 – July 02, 2018	288
Meghalaya/ new	Nongkenbah	91.4035	25.2814	January 02, 2016 – March 11, 2019	433
Meghalaya/ new	Sohkhme	91.7789	25.2358	March 16, 2016 – March 03 2019,	395
Meghalaya/ new	Wahkhen	91.8510	25.3474	March 06, 2017 – March 02, 2020,	419
Assam	Bokakhat	93.5891	26.6409	March 19, 2014 – March 01, 2019,	616
Assam	Diphu	93.4248	25.8426	March 19, 2014 – February 23, 2020,	800
Assam	Goalpara	90.6306	26.1612	February 17, 2015 – March 03, 2020,	411
Assam	Guwahati	91.6579	26.1526	March 09, 2014 – March 05, 2020,	810
Assam	Kokrajhar	90.2762	26.4044	March 09, 2014 – February 18, 2019	665
Assam	Lumding	93.1781	25.7513	March 19, 2014 – February 20, 2020	826
Assam	Moridhal	94.5958	27.5331	March 19, 2014 – March 02, 2020	832
Assam	Nagaon	92.6858	26.3571	March 09, 2014 – February 23, 2020	812
Assam	Nalbari	91.4439	26.4365	March 09, 2014 – February 18, 2019	672
Assam	Sankardev College	93.8404	27.0768	March 19, 2014 – March 02, 2020	846
Assam	Teok	94.4581	26.8447	March 19, 2014 – February 26, 2019	707
Assam	Tezpur	92.8372	26.6973	March 09, 2014 – March 05, 2020	786
Assam	Tinskia	95.3644	27.4980	March 09, 2014 – February 26, 2019	621

Supplement 4: List of rain gauge stations used for the validation with GPM
DPR (Sylhet+Barak & Bengal Plain)

Area Name	Station Name	Longitude	Latitude	Periods	Num. samples
Sylhet+Barak	Amarshid	92.4764	24.8768	March 22, 2014 – March 03, 2019	717
Sylhet+Barak	Bulaganj	91.7506	25.1361	March 09, 2014 – March 05, 2020	787
Sylhet+Barak	Chhatak	91.6704	25.0379	March 09, 2014 – March 09, 2019	665
Sylhet+Barak	Haflong	93.0181	25.1723	May 25, 2014 – February 23, 2020	782
Sylhet+Barak	Hailakandi	92.5651	24.6908	May 24, 2014 – February 26, 2018	514
Sylhet+Barak	Jaflong	92.0198	25.1791	May 09, 2014 – March 05, 2020	762
Sylhet+Barak	BARI, Jaintapur	92.1359	25.1359	March 09, 2014 – March 05, 2020	460
Sylhet+Barak	Juri	92.1617	24.6361	-	0
Sylhet+Barak	Kulaura	92.0345	24.5270	March 22, 2014 – March 10, 2015	132
Sylhet+Barak	Naljuri	92.1617	25.1751	-	0
Sylhet+Barak	Rajnagar	91.8542	24.5222	March 09, 2014 – March 03, 2019	672
Sylhet+Barak	Sylhet	91.8842	24.9055	March 09, 2014 – August 11, 2019	249
Sylhet+Barak	Sylhet Airport	91.8694	24.9595	March 09, 2014 – March 03, 2019	184
Bengal Plain	Chittagong	91.8085	22.3531	March 29, 2014 – March 07, 2015	121
Bengal Plain	Dhaka	90.3784	23.7799	March 09, 2014 – November 05, 2018	521
Bengal Plain	Dinajpur	88.6545	25.6466	August 11, 2015 – March 08, 2019	471
Bengal Plain	BRRI, Habiganj	91.4286	24.4149	March 09, 2014 – August 09, 2019	696
Bengal Plain	Mymensingh	90.4263	24.7256	March 09, 2014 – September 10, 2018	539
Bengal Plain	Rajshahi	88.6549	24.3606	August 11, 2015 – March 08, 2019	471
Bengal Plain	Srimangal	91.7438	24.2950	March 09, 2014 – July 29, 2018	582