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1 **Do Dry GCMs Generate QBO-like**
2 **Oscillation?**

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Abstract

9

10 The effect of vertical discretization methods and vertical resolution on Quasi-
11 Biennial Oscillation (QBO)-like oscillations that can occur in mechanistic
12 General Circulation Models (dry GCMs) is investigated. Two models are
13 compared. One model uses the spectral method in the horizontal direc-
14 tion but the finite difference method in the vertical direction (VFD model),
15 while the other is a three-dimensional spectral model that uses the spectral
16 method for discretization in both the horizontal and vertical directions (3DS
17 model). Both models include horizontal hyperdiffusion, simple Newtonian
18 cooling and Rayleigh friction, but as they are dry models, they do not in-
19 clude the effects of moist convection, and no explicit vertical diffusion is
20 used, following a previous study. Long-term numerical integrations of these
21 models show that the 3DS model does not generate QBO-like oscillations
22 at the vertical resolution settings used. On the other hand, the VFD model
23 generates QBO-like oscillations at low vertical resolution, but no QBO-like
24 oscillations at higher vertical resolution. Wavenumber-frequency spectral
25 analyses of wave disturbances show that, in the VFD model, the amplitude
26 of the waves at the sigma-level near the central altitude of the QBO-like
27 oscillations is highly dependent on the vertical resolution of the model.
28 Analyses of the wave contribution to the vertical momentum fluxes and ad-
29 ditional numerical experiments show that in the higher vertical resolution

30 setting, steady eastward zonal winds form above the altitude corresponding
31 to the tropopause, and these zonal winds suppress the upward propagation
32 of eastward moving waves. Transformed Eulerian mean analyses are also
33 done for the results of the VFD models to investigate the contribution of
34 the residual circulation and the wave-mean-flow interaction to the QBO-like
35 oscillation.

36 **Keywords**

37 dry GCM; three-dimensional spectral model; quasi-biennial oscillation

38 **1. Introduction**

39 The Quasi-Biennial Oscillation (QBO) is a periodic reversal in the di-
40 rection of the zonal mean zonal wind with a downward shift in the phase of
41 the zonal wind in the tropical stratosphere. This phenomenon was discov-
42 ered by Ebdon (1960), Ebdon and Veryard (1961), and Reed et al. (1961),
43 and subsequent theoretical work by Lindzen and Holton (1968) proposed
44 the basic mechanism of this phenomenon, namely that such oscillations are
45 caused by the wave-mean-flow interaction. The waves considered in that
46 theory were gravity waves, but Holton and Lindzen (1972) updated the
47 theory to consider Kelvin waves and mixed Rossby-gravity (MRG) waves,
48 which had by then been identified. Regarding which types of waves actually
49 drive the QBO, Dunkerton (1997) used a two-dimensional numerical model
50 to show that Kelvin waves and MRG waves alone cannot provide enough
51 momentum flux to cause oscillations with realistic period and structure due
52 to the presence of Brewer-Dobson upwelling in the real atmosphere, and that
53 gravity waves play an important role in driving the QBO. Indeed, recent
54 studies using modern high-resolution General Circulation Models (GCMs)

55 and reanalysis data have shown that not only large-scale waves such as
56 Kelvin and MRG waves, but also small-scale gravity waves are important
57 in driving the QBO, and that for the zonal mean zonal wind acceleration,
58 both Kelvin waves and small-scale gravity waves contribute to the eastward
59 acceleration, while gravity waves are mainly responsible for the westward
60 acceleration (see Anstey et al., 2022, the latest comprehensive review of the
61 QBO, and references therein).

62 The waves driving the QBO are thought to be mainly excited by moist
63 convection in the tropics, and the studies using modern GCMs mentioned
64 above have naturally included the moist process in their models. In con-
65 trast, Yao and Jablonowski (2015, hereafter referred to as YJ2015) per-
66 formed numerical experiments with four dry dynamical cores, which are
67 options in NCAR’s Community Atmosphere Model, version 5 (CAM5), and
68 showed that QBO-like oscillations can be obtained. The four dynamical
69 cores were, the spectral transform semi-Lagrangian (SLD) model, finite-
70 volume (FV) model, spectral transform Eulerian (EUL) model, and spectral
71 element (SE) model, and there simple Newtonian temperature relaxation
72 and Rayleigh friction were incorporated (hereafter, we refer to dynamical
73 cores with these effects as dry GCMs). YJ2015 investigated how the choice
74 of dynamical core affects the generation of QBO-like oscillations. Whereas
75 three of the four dry GCMs used in that study, SLD, EUL, and SE models,

76 generated QBO-like oscillations with largely different amplitudes and peri-
77 ods, the FV model did not generate QBO-like oscillations. In YJ2015, the
78 cause of the difference in the occurrence/non-occurrence of QBO-like oscil-
79 lations was explored to some extent, with the FV model where QBO-like
80 oscillations did not occur clearly having less wave activity than the other
81 three models, and this was thought to be the cause of the difference in the
82 occurrence/non-occurrence of QBO-like oscillations. However, it is not clear
83 what was responsible for such a difference in the magnitude of wave activity,
84 although some analysis of the instability of the calculated fields was done
85 there. Furthermore, there is a possibility that the solutions without QBO-
86 like oscillations, i.e., with less wave activity, could be the true solution if we
87 were able to solve the governing equation of the YJ2015 setting analytically
88 without any discretization error at all. This possibility was also mentioned
89 in YJ2015, but each of the four dynamical cores used there differs signifi-
90 cantly in aspects other than vertical discretization, and how the methods
91 and accuracy of vertical discretization actually affect the degree of wave ac-
92 tivity and the occurrence/non-occurrence of QBO-like oscillations has not
93 been explored.

94 Based on the above background, the purpose of the present manuscript
95 is not to explore a realistic QBO but to investigate how the vertical dis-
96 cretization method and accuracy affect the occurrence/non-occurrence of

97 QBO-like oscillations in dry GCMs, by conducting numerical experiments
98 using two dry GCMs with different vertical discretization methods, and
99 comparing them with different vertical discretization points. Two models
100 are used here: one is a model similar to the EUL used in YJ2015, using the
101 spectral method with the spherical harmonic expansion in the horizontal di-
102 rection and the finite difference method in the vertical direction; the other
103 is a three-dimensional spectral model based on the formulation of Ishioka et
104 al. (2022). The remainder of the present manuscript is organized as follows.
105 In Section 2, descriptions of two models and experimental settings are pre-
106 sented. Results are shown in Section 3. In Section 4, results of additional
107 numerical experiments are shown. Summary and discussion are given in
108 Section 5.

109 **2. Methods**

110 *2.1 Model description*

111 We use two dynamical cores to study the effect of different vertical dis-
112 cretizations on the generation/non-generation of QBO-like oscillations. One
113 is a three-dimensional spectral model (3DS model) constructed on the basis
114 of the formulation presented by Ishioka et al. (2022). The model uses the
115 primitive equations in σ -coordinate on a rotating sphere as the governing

116 equations and the dependent variables are expanded by spherical harmonics
117 in the horizontal direction and Legendre polynomials in the vertical direc-
118 tion. Time integration is performed using AM2*/AX2* method, which is an
119 IMEX(Implicit-explicit linear multistep) method proposed by Durran and
120 Blossey (2012), with a combination of the implicit trapezoidal scheme and
121 a third-order three-stage Runge-Kutta method for the computation of ini-
122 tial two steps. In time integration, the semi-implicit method where linear
123 gravity wave components are computed separately is adopted. For more
124 details on the system of equations and numerical implementation, as well
125 as the results of validation with standard test cases such as Held and Suarez
126 (1994) and Polvani et al. (2004), see Ishioka et al. (2022). In the present
127 manuscript, numerical experiments are carried out using the 3DS model
128 with two settings of 85 and 170 vertical truncation wavenumbers of the
129 Legendre polynomials (the number of vertical grid points is 128 and 256,
130 respectively). Each setting is referred to as M85 and M170, respectively, in
131 the following. The horizontal truncation wavenumber is set as T63, that is,
132 the triangular truncation wavenumber of the spherical harmonic expansion
133 is 63. The horizontal grid setting used for the spectral transform method
134 is 256×128 (longitude \times latitude). The other is a vertical finite difference
135 (VFD) model, which has the same formulation as the three-dimensional
136 spectral model above, except that it uses the finite difference method for

137 vertical discretization. The implementation of vertical difference in this
 138 VFD model is based on the description in Chapter 8 of Durran (2010),
 139 which is similar to Hoskins and Simmons (1975) and Bourke (1974). En-
 140 ergy conservation in discrete form is also considered following Corby et al.
 141 (1972). The vertical grid point distribution of the VFD model is based
 142 on YJ2015. However, as YJ2015 uses a hybrid coordinate, the following
 143 procedure is used to approximate and convert them to the corresponding
 144 σ -coordinates. All YJ2015 numerical experiments were carried out with 55
 145 vertical layers, where the pressure p at each interface of the layers was given
 146 by the following equation (Eq. A1 of YJ2015),

$$p_{k+1/2} = a_{k+1/2}p_0 + b_{k+1/2}p_s \quad (k = 0, 1, \dots, 55), \quad (1)$$

147 where p_s is the surface pressure, $p_0 = 1000$ hPa, and $(a_{k+1/2}, b_{k+1/2})$ are
 148 predetermined constants (Table A1 of YJ2015). In the present manuscript,
 149 assuming $p_s \approx p_0$, each of the model interface σ -level of the VFD with 55
 150 layers is set as follows:

$$\sigma_{k+1/2} = p_{k+1/2}/p_s = a_{k+1/2} + b_{k+1/2} \quad (k = 0, 1, \dots, 55). \quad (2)$$

151 The VFD model with 55 layers, where the σ -levels are defined in this way, is
 152 referred to as the L55 in the present manuscript as the standard setting. In
 153 the present manuscript, numerical experiments are also carried out for the
 154 VFD model in a setting with higher vertical resolution. In those, 110 layers

155 are used, which will be referred to as L110. There, each layer defined by the
156 L55 is divided into two layers using cubic spline interpolation in the $\log \sigma$
157 coordinate, so that the layer depths vary smoothly. As with the 3DS model,
158 the VFD model has also been tested in standard test cases such as Held
159 and Suarez (1994) and Polvani et al. (2004) (not shown). The vertical grid
160 point distributions are shown in Fig. 1, including those of the 3DS model.
161 Note that for the VFD models, the σ -coordinates of the midpoint of each
162 layer are shown.

Fig. 1

163 *2.2 Experimental settings*

164 The basic setup of the numerical experiments in the present manuscript
165 follows that of YJ2015. First, the forcing introduced in Held and Suarez
166 (1994) is imposed. That is, a Rayleigh friction is added to the equations
167 of motion at low levels ($\sigma > 0.7$), and a Newtonian relaxation towards
168 an equilibrium state is imposed on the evolution equation of the tempera-
169 ture field. The atmosphere is a dry atmosphere and the effects of diabatic
170 processes other than the Newtonian relaxation, topography and seasonal
171 variations are not included. Furthermore, a sponge layer is introduced in
172 the upper layer of the model according to YJ2015. There, Rayleigh fric-
173 tion is applied only to the zonal flow; YJ2015 imposed a Rayleigh friction
174 coefficient $K_r(p/p_0)$ determined as a function of p/p_0 in the region where

175 the air pressure p is smaller than 1 hPa. As the models used in the present
176 manuscript use the σ -coordinate system in the vertical direction, full com-
177 pliance with this setting would incur some extra computational cost, so
178 a form of this setting is mimicked, where the Rayleigh friction coefficient
179 K_r is set as $K_r(\sigma)$ for $\sigma < 0.001$. For the horizontal diffusion term, the
180 models in the present manuscript use the same setting as Held and Suarez
181 (1994), i.e., a hyperviscosity of a Laplacian raised to the fourth power (∇^8)
182 is imposed on both the wind and temperature fields, and the e-folding time
183 for the horizontal truncation wavenumber is 0.1 days. This setting differs
184 from the EUL model of YJ2015, but this difference will be mentioned again
185 in Section 4.1. Note that no explicit vertical diffusion is used, following
186 YJ2015.

187 All numerical experiments in the present manuscript are performed as
188 follows. The initial state is an isothermal stationary atmosphere at 300 K
189 with a small temperature perturbation, from which a time integration is
190 performed for a spin-up period of 200 days. The results of the last day of
191 the spin-up period (day 200) are stored and used again as the initial condi-
192 tions for numerical time integration. Whereas YJ2015 performed ensemble
193 numerical experiments and constructed some initial states using reanalysis
194 data, we do not use such an approach. One reason for this is that the sys-
195 tem of equations is essentially chaotic and the states after sufficient time

196 evolution are not sensitive to the initial state, at least statistically. Another
197 reason is that the initial state based on reanalysis data should contain waves
198 that cause the QBO in the real atmosphere, and such waves could cause
199 QBO-like oscillations in the numerical experiments in the dry GCMs used
200 in the present manuscript, at least in the early stages of time integration.
201 As the aim of the present manuscript is to investigate whether QBO-like
202 oscillations are spontaneously induced by waves excited in the dry GCMs,
203 reanalysis data are not used.

204 The time step used for the time integration is 720 s for all vertical
205 resolutions (L55 and L110) for the VFD model, but for the 3DS model,
206 it is 120 s for the spin-up phase for M85 and 300 s thereafter, 60s for the
207 spin-up phase for M170 and 150 s thereafter. This is because the vertical
208 resolution in the 3DS model is higher than in the VFD model near the
209 lower atmospheric boundary, and also because during the spin-up period,
210 the deviation from the balanced state is large and gravity waves with higher
211 frequencies are strongly excited. It is certainly possible that differences in
212 the time steps could affect the experimental results. However, we have
213 confirmed that the results show little difference depending on the the time
214 steps by actually performing the calculation with a time step of 150 s for all
215 the models, so the above time steps are adopted to reduce the calculation
216 time.

217 *2.3 Spectral analysis methods*

218 *a. Temperature disturbances*

219 Wavenumber-frequency spectral analyses of the temperature field at low
 220 latitudes are performed as follows. After 3600 days of time integration,
 221 900 days of hourly data are stored. We then calculate the temperature of
 222 specified σ -levels by linear interpolation, as the VFD and the 3DS model
 223 have different vertical grid point distributions and it is not appropriate to
 224 compare the results obtained in the original σ -levels. The symmetric part T_s
 225 and the antisymmetric part T_a of the temperature field with respect to the
 226 equator are calculated and the discrete 2D Fourier transform is performed
 227 at each latitude as follows:

$$T_s(\lambda, \phi, t) = (T(\lambda, \phi, t) + T(\lambda, -\phi, t))/2, \quad (3)$$

$$T_a(\lambda, \phi, t) = (T(\lambda, \phi, t) - T(\lambda, -\phi, t))/2, \quad (4)$$

$$\hat{T}_{s/a}(m, \nu, \phi) = \frac{1}{JK} \sum_{j=0}^{J-1} \sum_{k=0}^{K-1} T_{s/a}(\lambda_j, \phi, t_k) e^{-im\lambda_j + 2\pi i\nu t_k}. \quad (5)$$

228 Here, $T(\lambda, \phi, t)$ is the temperature field, λ is the longitude, ϕ is the latitude,
 229 t is the time, $\lambda_j = 2\pi j/J$ ($j = 0, 1, \dots, J-1$) is the grid points in the
 230 longitudinal direction, and $J = 256$ is the number of grid points, t_k ($k =$
 231 $0, 1, \dots, K-1$) is the time of the hourly data, $K = 900 \times 24$ is the number of
 232 the stored hourly data, $m = -N, \dots, 0, 1, \dots, N$ is the zonal wavenumber,
 233 $N = 63$ is the horizontal truncation wavenumber, $\nu = 0, \Delta\nu, \dots, (K/2 -$

234 1) $\Delta\nu$ is the frequency, and $\Delta\nu = 1/900 \text{ day}^{-1}$. Power spectral density
 235 $P_{s/a}(m, \nu, \phi)$ is defined as follows:

$$P_{s/a}(m, \nu, \phi) = |\hat{T}_{s/a}(m, \nu, \phi)|^2. \quad (6)$$

236 By averaging $P_{s,a}(m, \nu, \phi)$ in 15°S–15°N, we obtain the wavenumber-frequency
 237 spectra of the temperature field for the low latitude region.

238 *b. Vertical momentum flux*

239 We also conduct spectral analyses of the vertical momentum flux in the
 240 following steps. After 3600 days of time integration, we stored 900 days of
 241 hourly velocity data interpolated to a prescribed p -coordinate. Each level of
 242 this p -coordinate is set to $p = p_0\sigma$ for each level of the original σ -coordinate.
 243 The following discrete 2D Fourier transform is then performed at each pres-
 244 sure level and latitude grid, analogous to the previous subsection:

$$\hat{u}(m, \nu, \phi, p) = \frac{1}{JK} \sum_{j=0}^{J-1} \sum_{k=0}^{K-1} u'(\lambda_j, \phi, t_k, p) e^{-im\lambda_j + 2\pi i\nu t_k}, \quad (7)$$

$$\hat{\omega}(m, \nu, \phi, p) = \frac{1}{JK} \sum_{j=0}^{J-1} \sum_{k=0}^{K-1} \omega'(\lambda_j, \phi, t_k, p) e^{-im\lambda_j + 2\pi i\nu t_k}. \quad (8)$$

245 Here, the primed quantities are the perturbations from the zonal mean field,
 246 u is the zonal velocity, and ω is the vertical pressure velocity. The zonal
 247 wavenumber $m = -N, \dots, -1, 1, \dots, N$, and the number of the hourly data
 248 points is set as either $K = 900 \times 24$ or 120×24 , depending on whether the

249 period of days to be analyzed is 900 or 120 days. The vertical momentum
 250 flux is then summed for each phase velocity as follows:

$$S_{\overline{u'\omega'}}(m, \nu, \phi, p) = \frac{2}{\Delta\nu} \operatorname{Re} \{ \hat{u}(m, \nu, \phi, p) \hat{\omega}^\dagger(m, \nu, \phi, p) \}, \quad (9)$$

$$C_{\overline{u'\omega'}}(c, \phi, p) = \sum_{c-\Delta c \leq c_p(m, \nu) < c+\Delta c} S_{\overline{u'\omega'}}(m, \nu, \phi, p) \Delta\nu. \quad (10)$$

251 Here, the overbar denotes the zonal mean, $\operatorname{Re}\{\cdot\}$ denotes the real part of
 252 the complex number, the dagger denotes the complex conjugate, $c_p(m, \nu)$
 253 is the zonal phase velocity of each wave, defined by $c_p(m, \nu) = 2\pi a\nu/m$,
 254 where a is the radius of the earth. In (10), the bins for the phase velocities
 255 are determined by setting $c = 0, \pm 1, \pm 2, \dots$ (the unit is m s^{-1}), and $\Delta c =$
 256 0.5 m s^{-1} . Also note that in (9) and (10), $\Delta\nu$ is set as either $1/900 \text{ day}^{-1}$ or
 257 $1/120 \text{ day}^{-1}$, depending on whether the period of days to be analyzed is 900
 258 or 120 days. By averaging $S_{\overline{u'\omega'}}$ and $C_{\overline{u'\omega'}}$ in an equatorial region of 2°S – 2°N
 259 at each pressure level, we can examine how the wavy components contribute
 260 to the vertical momentum flux in the wavenumber-frequency space and the
 261 zonal phase velocity space, respectively. Such an analysis was also carried
 262 out, for example, in Fig. 18 in Horinouchi and Yoden (1998).

263 3. Results

264 3.1 Zonal wind profile and static stability distribution

Fig. 2

265 In order to investigate whether different vertical discretization methods
266 and different resolutions affect whether QBO-like oscillations occur or not,
267 the time evolution of the zonal mean zonal wind over the equatorial re-
268 gion (averaged in 2°S–2°N) is shown for the VFD (L55 and L110) and 3DS
269 (M85 and M170) models (Fig. 2). The time evolution of the zonal mean
270 zonal wind obtained by the VFD models is strongly dependent on the ver-
271 tical resolution, with the L55 VFD model showing a QBO-like oscillation
272 (Fig. 2a), while the L110 VFD model shows no oscillatory variation after
273 5 years of time integration (Fig. 2c). In the L55 VFD model, the period
274 of the QBO-like oscillation is about 3 years, which is considerably shorter
275 than that obtained in the EUL of YJ2015 (about 15 years). Although the
276 periods do not match, the downward phase propagation of QBO-like oscil-
277 lations in the L55 VFD model is similar to that of the EUL in YJ2015.
278 That is, the region with eastward wind descends to around $\sigma = 0.02$, while
279 the region with westward wind only descends to around $\sigma = 0.01$. The 3DS
280 model also shows dependence on vertical resolution, although not as much
281 as the VFD, and no QBO-like oscillations occur at either vertical resolu-
282 tion. The M85 3DS model (Fig. 2b) shows dominant eastward winds in

283 $0.0002 < \sigma < 0.05$, while the M170 3DS model (Fig. 2d) shows westward
284 winds in $0.001 < \sigma < 0.002$. Even in models that did not generate QBO-like
285 oscillations, however, the phase propagates downwards in the early stages
286 of time evolution in the first year (M85 3DS) or for about four years (L110
287 VFD and M170 3DS). This downward phase propagation may be due to the
288 shortness of the spin-up period to reach the equilibrium state. The 200 day
289 of spin-up period simply follows the setting of Held and Suarez (1994), who
290 mainly focused on the troposphere where 200 days of spin-up period seems
291 to be sufficient. In this study, however, we focus mainly on the stratospheric
292 QBO-like oscillations, which have much longer periods than 200 days. How-
293 ever, the data analyses in the following sections are carried out well after
294 the equilibrium state has been reached, so the analyses are not affected by
295 this short spin-up period and the initial downward phase propagation. The
296 initial downward propagation of the phase in these models is similar to the
297 results of the FV model of YJ2015, where no QBO-like oscillations occur.

298 Whereas the occurrence of QBO-like oscillations at the equator depends
299 on the resolution as well as the vertical discretization method, the merid-
300 ional structure of zonal winds depends mainly on the vertical discretization
301 method. Figure 3 shows latitude- σ cross section of the zonal and monthly
302 mean zonal wind, where the average is taken over day 2970–3000. In the
303 region where $\sigma > 0.1$, the location and strength of the eastward jets do not

304 change and its structure is similar in all models. On the other hand, in the
305 region where $\sigma < 0.1$, the latitudinal extent of the strong zonal wind over the
306 equator depends less on the vertical resolution and mainly on the vertical
307 discretization method. Specifically, the eastward wind region is 10°S – 10°N
308 in the VFD model (Fig. 3a, 3c), but 20°S – 20°N in the 3DS model (Fig. 3b,
309 3d). For the correspondence with the EUL model in YJ2015, the result
310 of the L55 VFD model is consistent in that the latitudinal extent of the
311 eastward wind area over the equator is 10°S – 10°N and the extent of the
312 westward wind area is 20°S – 20°N (Fig. 3a).

Fig. 3

313 As shown above, the zonal mean zonal wind structure differs signifi-
314 cantly depending on both the vertical discretization method and the vertical
315 resolution, but the zonal mean static stability fields show little difference
316 among the models. The result of the L55 VFD model is shown in Fig. 4
317 as a representative example. In this experimental setting, there are two
318 distinct regions over the equator; a low static stability region in the lower
319 atmosphere ($\sigma > 0.2$) and a high static stability region in the upper atmo-
320 sphere ($\sigma < 0.1$). Accordingly, we hereafter refer to this σ -level of 0.1 as
321 the “tropopause” in the equatorial region.

Fig. 4

322 3.2 Spectral analysis

323 a. Wavenumber-frequency spectra of the temperature disturbances

Fig. 5

324 Next, we analyze the wave activity to investigate the causes of the differ-
325 ences in the occurrence/non-occurrence of QBO-like oscillations in different
326 models. Figure 5 shows the results of the wavenumber-frequency analyses
327 of the temperature field at $\sigma = 0.002$, the central altitude of the QBO-like
328 oscillations observed in the L55 VFD model in Fig. 2a, using the method
329 described in Section 2.3 for four different model setups. There, the negative
330 (positive) zonal wavenumber in each panel represents westward (eastward)
331 propagation. Comparing the four models, the L55 VFD model has the
332 strongest wave activity among the four models, especially for the symmetric
333 component (Fig. 5a), which is consistent with the finding that only the L55
334 VFD model generates QBO-like oscillations. A closer look at the spectra
335 shows that, regardless of the vertical resolution, in both the VFD and 3DS
336 models there are spectral peaks at low frequencies and low zonal wavenum-
337 bers ($|m| < 6, \nu < 0.3$ cpd), especially for the symmetric component, near
338 dispersion curves corresponding to Rossby and Kelvin waves (Fig. 5a, 5c,
339 5e, and 5g). However, the extent and intensity of such spectral peaks vary
340 with model settings, particularly for the symmetric components. Below we
341 will focus on the regions where the value of the common logarithm shown
342 in Fig. 5 is larger than -4 . For the symmetric components of the L55

343 VFD model (Fig. 5a), the intensity and extent of the spectral peak of the
344 westward and eastward components are comparable. However, for the L110
345 VFD model (Fig. 5e), the peak for eastward components is weaker. For the
346 L55 VFD, a strong and broad peak is seen in the region of the dispersion
347 curve corresponding to the Kelvin wave of $h > 12$ m, while for the L110
348 VFD it is mainly seen only near the region corresponding to the Kelvin wave
349 of $h \approx 200$ m. Here h is the equivalent depth of the waves. With regard
350 to the 3DS models, the M85 and M170 3DS models have similar intensity
351 and extent of the spectral peaks for the westward symmetric component,
352 whereas the M85 3DS model has a narrower spectral peak for the eastward
353 symmetric component (Fig. 5c and 5g).

Fig. 6

354 As seen above, depending on whether QBO-like oscillations were present
355 or not, the level of wave activity was indeed very different, but care must
356 be taken when interpreting the results, as the upward propagating waves
357 may be filtered by the mean zonal wind, which varies considerably among
358 the models. Since the vertical wavelengths become short and the vertical
359 group velocity approaches zero as the gravity waves approach their critical
360 level, where their phase velocity is equal to the background horizontal wind
361 velocity, the different wave activities seen above may be the result of the
362 different zonal wind profiles obtained in the different models. To investigate
363 this effect, we perform spectral analyses at $\sigma = 0.1$, below which the zonal

364 wind profiles are similar in all the models used here. The results are shown
 365 in Fig. 6. All the models show similar wave activity at $\sigma = 0.1$, except
 366 that the L55VFD model has slightly weaker wave activity than the oth-
 367 ers, with spectral peaks at lower frequencies and lower zonal wavenumbers
 368 ($|m| < 6, \nu < 0.2$ cpd) for the symmetric components. This result suggests
 369 that the strength of the upward propagating wave that enters from below
 370 $\sigma = 0.1$ is not significantly different among all the models. Thus, there are
 371 at least three possible reasons for the different wave activities at $\sigma = 0.002$;
 372 different wave filtering by the background zonal wind, different wave dis-
 373 sipation depending on the vertical discretization or resolution, and wave
 374 generation at $0.002 < \sigma < 0.1$. That is, we cannot deduce which one(s) is
 375 (are) occurring from the results shown so far.

376 *b. Wave contribution to the vertical momentum flux in the zonal phase*
 377 *velocity space*

Fig. 7

378 To clarify the reason why the wave activities are different at $\sigma = 0.002$
 379 but not significantly different at $\sigma = 0.1$ as shown in the previous subsec-
 380 tion, the wave contributions to the vertical momentum flux $\overline{u'\omega'}$ in the zonal
 381 phase velocity space are examined using the method described in Section 2.3.
 382 The result for day 3600–4500 is shown in Fig. 7. Note that a positive (neg-
 383 ative) zonal phase velocity represents an eastward (westward) propagating

384 wave and that negative values of the contributions to the momentum fluxes
385 indicate upward fluxes of eastward momentum, as the fluxes are shown in
386 the forms of $\overline{u'\omega'}$. Note also that in Fig. 7, the vertical coordinate is the
387 pressure. In Fig. 7, there is no significant difference in the dependence of
388 the momentum fluxes on the wave phase velocity among the models in the
389 region of $p > 100\text{hPa}$, which is consistent with the fact that there was no
390 significant difference in the spectral distribution of the temperature distur-
391 bances among the models at $\sigma = 0.1$. On the other hand, in the region
392 of $p < 100\text{ hPa}$, the vertical propagation of the waves differs significantly
393 among the models. In particular, it is clear that the differences depend on
394 the realized zonal mean zonal wind profile (indicated by the solid curves in
395 Fig. 7). The vertical momentum fluxes shown in Figs. 7a-d indicate that, in
396 all the cases, the vertical propagation of the waves is significantly suppressed
397 at the critical levels, and that the positive and negative upward flux of the
398 eastward momentum roughly corresponds to the sign of the difference be-
399 tween the wave phase velocity and the zonal mean zonal wind velocity. This
400 result suggests that the upward energy fluxes are all positive for $p < 100\text{hPa}$
401 for each phase velocity wave components. One notable difference in Fig. 7
402 is that in all cases except for Fig. 7a for the L55 VFD, the zonal mean zonal
403 wind is eastward around $p = 50\text{hPa}$, where the upward momentum fluxes of
404 the eastward waves are sharply reduced, i.e. the upward propagation of the

405 eastward waves is suppressed (Figs. 7b-d). This suppression of the vertical
 406 propagation of the eastward moving waves is also visible in Fig. 7a around
 407 $p = 20\text{hPa}$, but is not as sharp as in Figs. 7b-d. These results indicate that
 408 the formation of strong eastward winds around $p = 50\text{hPa}$ and at higher al-
 409 titudes, and the presence of sharp suppression of the vertical propagation of
 410 eastward moving waves there, determine the degree of penetration of east-
 411 ward propagating waves to higher altitudes. Such suppression is not strong
 412 in Fig. 7a and the momentum fluxes of eastward propagating waves remain
 413 large up to around 2 hPa, which may be the reason why L55 VFD (Fig. 5a)
 414 shows a larger eastward component of the temperature disturbance than
 415 L110 VFD (Fig. 5e), M85 3DS (Fig. 5c) and M170 3DS (Fig. 5g).

Fig. 8

416 In the L55VFD model, the mean zonal wind profile changes significantly
 417 over the 900-day period for which the spectral analysis was performed in
 418 Fig. 7. Therefore, it is necessary to consider the effect of temporal changes
 419 in the zonal wind profile, in particular to discuss the relationship between
 420 the suppression of vertical wave propagation and the zonal wind profile. For
 421 this reason, a similar analysis is performed for the L55 VFD model using
 422 120 days of data. Figure 8 shows the results for days 3960–4080 and 4380–
 423 4500. We have chosen these periods as representative of the westward and
 424 eastward shear phases around $p = 5\text{ hPa}$. Here, eastward (westward) shear
 425 refers to $\partial\bar{u}/\partial p < 0$ ($\partial\bar{u}/\partial p > 0$). In Fig. 8, in both phases (Figs. 8a and

426 8b), an eastward vertical shear of the zonal mean zonal wind appears at the
427 pressure levels around $p = 20\text{hPa}$, and the vertical profile of the momentum
428 flux shows that the upward propagation of waves is suppressed in the region
429 of the shear, again similar to that seen in Fig. 7a. However, the suppression
430 of upward propagation is not as sharp as that in the same altitude region in
431 the other models, such as the L110 VFD. In addition to this, the result of
432 the eastward phase in the L55 VFD model (Fig. 8b) shows that even in the
433 eastward shear around $p = 5\text{hPa}$, fast phase speed waves ($c_p > 16 \text{ m s}^{-1}$)
434 that do not reach their critical level are not sharply suppressed there and
435 propagate higher up. This result is in contrast to that of the L110 VFD
436 model (Fig. 7c), where eastward-moving waves are sharply reduced around
437 $p = 5 \text{ hPa}$ even if such waves do not reach their critical levels. The above
438 results suggest that the representation of the wave-mean-flow interaction is
439 altered by the vertical resolution, which may be the reason why the eastward
440 zonal wind formation around $p = 50 \text{ hPa}$ seen in the other models is not
441 seen in the L55 VFD model.

442 *3.3 Zonal wind acceleration in the VFD models*

443 *a. Transformed Eulerian mean analyses*

444 In order to quantitatively investigate the causes of the different zonal
445 wind profiles obtained in the L55 VFD model and the L110 VFD model, we

446 perform the transformed Eulerian mean analyses. The zonal wind acceler-
 447 ation is expressed by the following equation, the form of which follows that
 448 used in YJ2015:

$$\frac{\partial \bar{u}}{\partial t} = \bar{v}^* \left[f - \frac{1}{a \cos \phi} \frac{\partial}{\partial \phi} (\bar{u} \cos \phi) \right] - \bar{\omega}^* \frac{\partial \bar{u}}{\partial p} + \frac{1}{a \cos \phi} \left[\frac{1}{a \cos \phi} \frac{\partial}{\partial \phi} (F_\phi \cos \phi) + \frac{\partial}{\partial p} F_p \right] + X. \quad (11)$$

449 Here, the definitions for symbols not already mentioned so far are as follows:
 450 f is the Coriolis parameter, \bar{v}^* and $\bar{\omega}^*$ are the meridional velocity and the
 451 vertical pressure velocity of the residual circulation, respectively, and F_ϕ and
 452 F_p are the meridional and vertical components of the Eliassen-Palm (EP)
 453 flux vector, respectively, and X represents the residual. Each component of
 454 the meridional velocity and the EP flux vector is defined by the following
 455 equations:

$$\bar{v}^* = \bar{v} - \frac{\partial}{\partial p} \left(\frac{\overline{v'\theta'}}{\partial \bar{\theta} / \partial p} \right), \quad (12)$$

$$\bar{\omega}^* = \bar{\omega} + \frac{1}{a \cos \phi} \frac{\partial}{\partial \phi} \left(\cos \phi \frac{\overline{v'\theta'}}{\partial \bar{\theta} / \partial p} \right), \quad (13)$$

456

$$F_\phi = a \cos \phi \left(\frac{\overline{v'\theta'}}{\partial \bar{\theta} / \partial p} \frac{\partial \bar{u}}{\partial p} - \overline{u'v'} \right), \quad (14)$$

$$F_p = a \cos \phi \left[\frac{\overline{v'\theta'}}{\partial \bar{\theta} / \partial p} \left(f - \frac{1}{a \cos \phi} \frac{\partial}{\partial \phi} (\bar{u} \cos \phi) \right) - \overline{u'\omega'} \right], \quad (15)$$

457 where, θ is the potential temperature and v is the meridional velocity. The
 458 four terms on the right-hand side of (11) are the horizontal and the verti-
 459 cal advection terms of the angular momentum by the residual circulation,

460 the EP flux divergence term, which represents the acceleration due to the
461 wave-mean-flow interaction resolved in the models, and the residual term,
462 which represents the effect of explicitly and implicitly added diffusion, in-
463 cluding at least the effects of the horizontal diffusion and Rayleigh friction
464 in the experimental setting of this study. Note that when evaluating these
465 terms from the output of the models, the derivatives are evaluated in the
466 σ -coordinate, then transformed and interpolated to the p -coordinate and
467 zonally averaged.

Fig. 9

468 The results of the TEM analyses for day 3600–5000 are shown in Fig. 9,
469 and this period captures approximately 1.3 cycles of the QBO-like oscillation
470 in the L55 VFD model. In the remainder of this section, we focus mainly on
471 the region above 20 hPa, where the QBO-like oscillation is seen in the L55
472 VFD model. Figures 9c and 9d show the acceleration due to the vertical
473 and meridional advection, where the L55 VFD model and the L110 VFD
474 model show similar features in the eastward shear region. That is, the
475 eastward shear regions show strong eastward accelerations ($0.2\text{--}0.4 \text{ m s}^{-1}$
476 day^{-1}). In the L55 VFD model, this strong eastward acceleration mainly
477 contributes to the descent of the eastward phase of the zonal mean zonal
478 wind. It should also be mentioned that, although not shown in the figure,
479 the meridional advection term for both the VFD models shows westward
480 accelerations in the eastward shear region with values of -0.1 to -0.2 m

481 $\text{s}^{-1} \text{ day}^{-1}$, in contrast to the YJ2015 results where the acceleration due to
482 the meridional advection was weak (with its values in the range of ± 0.03
483 $\text{m s}^{-1} \text{ day}^{-1}$). However, this acceleration by the meridional advection is
484 overwhelmed by the strong acceleration due to the vertical advection, so
485 that the net advection term shows the above feature.

486 In contrast to the advection term, the EP flux divergence term shows
487 different features between the L55 VFD model (Fig. 9e) and the L110 VFD
488 model (Fig. 9f). For the L55 VFD model, westward accelerations (~ -0.1
489 $\text{m s}^{-1} \text{ day}^{-1}$) can be seen in the westward shear regions, which is a major
490 contributor to the descent of the westward phase of the QBO-like oscillation.
491 In the eastward shear regions for the L55 VFD model, since both westward
492 and eastward accelerations occur, this term does not directly contribute to
493 the descent of the eastward phase. On the other hand, in the eastward
494 shear regions for the L110 VFD model, the acceleration due to the EP
495 flux divergence is almost entirely westward and the westward acceleration
496 has large absolute values ($\sim -0.4 \text{ m s}^{-1} \text{ day}^{-1}$), almost canceling out the
497 eastward accelerations of the advection term.

498 The residual term (Fig. 9g for L55 VFD model and 9h for L110 VFD
499 model), obtained by subtracting the advection and the EP flux divergence
500 terms seen so far from the net acceleration, has small absolute values (± 0.01
501 $\text{m s}^{-1} \text{ day}^{-1}$) and it is accounted for by the diffusion tendency, both for the

502 L55 VFD model and the L110 VFD model. This is because the residual term
503 minus the diffusion tendency is close to zero (Fig. 9i for the L55 VFD model
504 and 9j for the L110 VFD model), which means that the momentum budget is
505 almost completely balanced. The diffusion term includes the hyperdiffusion,
506 the Rayleigh friction introduced by following YJ2015 in the sponge layer
507 ($\sigma < 0.001$), and the Rayleigh friction of Held and Suarez (1994) near the
508 lower boundary ($\sigma > 0.7$). In the altitude region shown here, the effect of
509 the horizontal hyperdiffusion is dominant, and the diffusion tendency has
510 much smaller absolute values than the result of YJ2015. This is because
511 the hyperdiffusion used in the present manuscript is the Laplacian raised to
512 the fourth power (∇^8), which mainly dissipates waves of small horizontal
513 scale and has practically no effect on the large scale zonal wind.

Fig. 10

514 The acceleration characteristics of the L55 and L110 VFD models differ
515 mainly in the EP flux divergence in the eastward shear regions, and these
516 differences lead to the differences in the total acceleration in the eastward
517 shear regions. To clarify what causes this difference in the eastward shear
518 regions, the EP flux divergence is decomposed into the meridional (Figs. 10a
519 and 10b) and vertical components (Figs. 10c and 10d). For both the L55
520 and L110 VFD models, the meridional component (Figs. 10a and 10b) shows
521 similar strong westward acceleration in the eastward shear regions. The
522 vertical component (Figs. 10c and 10d), on the other hand, shows a much

523 stronger eastward acceleration in the L55 VFD model, while it is not as
524 strong in the L110 VFD model. The result that the acceleration due to
525 the vertical component of the EP flux divergence is significantly different
526 between the two models in the altitude range we are now considering ($p < 20$
527 hPa) is consistent with the finding in Section 3.2b that the negative $\overline{u'\omega'}$
528 associated with eastward propagating waves has larger absolute values for
529 the L55 VFD model than for the L110 VFD model in the altitude range we
530 are now focusing on.

531 Before closing this subsection, we would like to mention a point that
532 some readers may wonder about, namely, why strong shear can be main-
533 tained in the L110 VFD model, especially for the eastward shear at 10–6
534 hPa and the westward shear at 6–2 hPa. We believe that this is due to the
535 weakness of the diffusion term. In this experimental setting, the diffusion
536 term is very weak and the zonal mean zonal wind is hardly weakened directly
537 by the diffusion. Therefore, the zonal mean winds can be maintained once
538 they are formed in the early stages of time evolution by wave-mean-flow
539 interactions with then abundant gravity waves.

540 To summarize this subsection, the downward phase propagation ob-
541 served in the L55 VFD model is mainly supported by the acceleration due
542 to the advection in the eastward wind shear region, and the wave-mean-
543 flow interaction in the westward shear region. In the L110 VFD model, the

544 advection term causes large eastward accelerations in the eastward shear
545 region, but these accelerations are almost canceled out by the EP flux con-
546 vergence, resulting in small net accelerations. In both the L55 and L110
547 VFD models, the characteristics of the advection term and the meridional
548 component of the EP flux divergence are similar. It is the vertical compo-
549 nent of the EP flux divergence term that is mainly different between the
550 L55 and L110 VFD models, and this difference can be attributed to the
551 differences in the upward propagating waves.

552 *b. Residual circulations*

Fig. 11

553 The TEM analysis revealed that the advection term contributes to the
554 descent of the downward phase propagation in the L55 VFD model. How-
555 ever, in YJ2015, this term did not contribute significantly to the descent
556 of the zonal wind phase. Therefore, in this subsection we will focus on the
557 residual circulation, particularly on the vertical component, since the ac-
558 celerations due to the vertical advection are larger than those due to the
559 meridional advection. The latitude-pressure cross sections of the zonal mean
560 residual circulation averaged over days 3600–4680 are shown in Fig. 11. This
561 period captures almost one cycle of the QBO-like oscillation obtained in the
562 L55 VFD model. Note that the residual pressure velocity $\bar{\omega}^*$ is converted
563 to the residual vertical velocity \bar{w}^* using the following approximation to see

564 the differences clearly;

$$\bar{w}^* = -\frac{RT_0}{pg}\bar{\omega}^*. \quad (16)$$

565 Here, $R = 287 \text{ J K}^{-1} \text{ kg}^{-1}$ is the gas constant for the dry atmosphere,
566 $g = 9.806 \text{ m s}^{-2}$ is the gravitational acceleration, and $T_0 = 200 \text{ K}$ is the
567 reference temperature, which is a representative value in the altitude range
568 shown in Fig. 11. The result for the L55 VFD model (Fig. 11a) shows a
569 dominant downward motion ($\sim -0.2 \text{ mm s}^{-1}$) over 5°S – 5°N in $p \lesssim 30 \text{ hPa}$.
570 This altitude range corresponds to that in which the QBO-like oscillation
571 occurred. This downward motion is more similar to the result of the YJ2015
572 SLD than the YJ2015 EUL.

573 On the other hand, for the L110 VFD model (Fig. 11b), while strong
574 downward motions ($\sim -0.3 \text{ mm s}^{-1}$) are seen in the eastward shear regions
575 over the equator, especially for $4 \text{ hPa} < p < 9 \text{ hPa}$ and $p < 1.5 \text{ hPa}$, weak
576 upward motions ($\sim 0.1 \text{ mm s}^{-1}$) are seen in the westward shear regions.
577 This pattern of upward (downward) motion in the westward (eastward)
578 shear zone is similar to the QBO-induced secondary circulation in the real
579 atmosphere (Choi et al., 2002). That is, the vertical wind field is modulated
580 by the zonal wind shear associated with the QBO through the thermal
581 wind balance, resulting in upward (downward) motions in the westward
582 (eastward) shear regions. Therefore, the strong eastward accelerations by
583 the advection term in the eastward shear regions for the L110 VFD model

584 shown in Fig. 9d appear to be due to this downward motion, induced by
585 the secondary circulation.

Fig. 12

586 To investigate the time evolution of the vertical motion in the L55 VFD
587 model, the time series of the residual vertical pressure velocity $\bar{\omega}^*$ at $p = 4.9$
588 hPa, the central pressure level of the QBO-like oscillation in the L55 VFD
589 model, is shown in Fig. 12, where the time evolution of the zonal wind at the
590 same level is also shown. It is clear that the time evolution of the vertical
591 pressure velocity is followed by that of the zonal wind velocity, suggesting
592 that the vertical motion is modulated by the zonal wind shear in a manner
593 similar to the QBO-induced secondary circulation in the real atmosphere.
594 In addition, the vertical velocity is predominantly downwelling, with a brief
595 period of a weak upwelling. Taken together with the results in Fig. 11a, the
596 downwelling, which is different from the Brewer-Dobson upwelling in the real
597 equatorial atmosphere, is dominant in the 5°S – 5°N for the L55VFD model
598 and this downwelling is modulated by the secondary circulation. These
599 results show that the strong acceleration due to vertical advection in the
600 eastward shear region for the L55 VFD model is caused by the enhancement
601 of the downwelling by the secondary circulation.

602 4. Additional numerical experiments

603 4.1 Numerical experiment as close as possible to the YJ2015

604 *EUL setting*

Fig. 13

605 As described in Section 2, the present manuscript uses a hyperdiffusion
606 of the Laplacian raised to the fourth power (∇^8) as the horizontal diffusion,
607 following Held and Suarez (1994), while the YJ2015 EUL uses a hyperdiffu-
608 sion of the Laplacian raised to the second power (∇^4), and considering the
609 value of the hyperdiffusion coefficient used there, for the total wavenumber
610 of spherical harmonic functions below 50, the time constant of dissipation
611 is smaller than the setting in the present manuscript. In addition, the
612 Robert-Asselin time-filtered Leapfrog scheme was used as the time integra-
613 tion method in the YJ2015 EUL, which can also contribute to the differences
614 in the results. To check the effects of these differences with YJ2015 EUL, a
615 numerical experiment is carried out using the same hyperdiffusion and time
616 integration method as for YJ2015 EUL case. The rest of the numerical
617 setup is the same as for the standard L55 VFD model. The time evolution
618 of the zonal mean zonal winds in this case is shown in Fig. 13. In this case,
619 a QBO-like oscillation with a period of about 4 years is generated, which is
620 longer than in the case of the standard L55 VFD (Fig. 2a). However, the
621 period is still shorter than in the YJ2015 EUL case (about 15 years). This

622 remaining difference may be due to differences in specific implementations,
623 such as the semi-implicit method for the following reason: In order to im-
624 plement the semi-implicit method, we need to specify a reference state from
625 which the governing equations are linearized. In the present manuscript,
626 this reference state is set to a 300 K isothermal atmosphere at rest, while
627 such a reference state is not explicitly described in YJ2015. If the reference
628 state is different, the behavior of the linear waves calculated with a finite
629 time step may be different. However, pursuing this difference is beyond the
630 scope of this paper and will not be considered further.

631 *4.2 Partially higher resolution setting*

Fig. 14

632 As seen in Section 3.1, doubling the number of vertical layers from L55
633 to L110 in the VFD model eliminated QBO-like oscillations, but it was
634 not clear at which altitude the increased resolution led to the suppression
635 of these QBO-like oscillations. Numerical experiments were therefore con-
636 ducted with a VFD model in which the vertical resolution was doubled
637 from the L55 VFD model only at a certain altitude range. The number
638 of the vertical grids at the altitude range shown below is doubled by cubic
639 spline interpolation, in a similar manner to the L110 VFD model. The cor-
640 responding σ ranges and the total numbers of the vertical layers are, (a):
641 $0.12 < \sigma < 0.49$, L66, (b): $0.028 < \sigma < 0.12$, L62, (c): $0.028 < \sigma < 0.49$,

642 L74, (d): $0.0067 < \sigma < 0.49$, L81, and (e): $0 < \sigma < 0.49$, L96, respec-
643 tively. Figure 14 shows the time evolution of the zonal mean zonal wind for
644 these additional numerical experiments. The experimental setup is iden-
645 tical to that of the L55 VFD model, except that the number of vertical
646 grid points is increased. In Fig. 14a, a QBO-like oscillation is observed,
647 with a period of about 3 years. In Fig. 14b and 14c, QBO-like oscillations
648 do not occur. These two settings (Figs. 14b and 14c) show two altitude
649 regions with eastward zonal wind and two altitude regions with westward
650 zonal wind. In Fig. 14d, in the region with increased vertical resolution
651 ($0.0067 < \sigma < 0.49$), a steady eastward wind region and a westward wind
652 region are formed. However, in regions where the vertical resolution is not
653 increased, QBO-like oscillations occur with a period of about 8 years. In
654 Fig. 14e, where the vertical resolution is increased to the top of the model,
655 three eastward and three westward wind regions are formed and no QBO-
656 like oscillations occur. It is worth noting here that in the setting for Fig. 14b,
657 the number of vertical grid points is only increased by seven compared to
658 the L55 VFD, but steady and strong eastward wind region above 16 m s^{-1}
659 is formed in $0.01 < \sigma < 0.05$ and QBO-like oscillations are no longer ob-
660 served at all. In addition to this, a comparison of the results in Figs. 14e
661 and 14d shows that in the region of $0.02 < \sigma < 0.05$, where the vertical
662 resolution is similarly high, an eastward wind of about 8 m s^{-1} is commonly

663 formed. However, the behavior in the region of $\sigma < 0.0067$, where the ver-
664 tical resolution is different, is significantly different. In Fig. 14d, in which
665 the vertical resolution in the $\sigma < 0.0067$ region is lower, QBO-like oscilla-
666 tions occur, whereas in Fig. 14e, in which the vertical resolution is higher,
667 strong eastward winds of more than 16 m s^{-1} are formed and no QBO-like
668 oscillations occur. These results suggest that the vertical resolution in the
669 $\sigma < 0.12$ region is strongly related to the formation of a steady eastward
670 wind region, and in particular the formation of a strong eastward wind re-
671 gion above 16 m s^{-1} is in turn strongly related to the non-occurrence of
672 QBO-like oscillations.

673 *4.3 VFD model with additional Rayleigh friction over a spe-* 674 *cific range of altitudes and latitudes*

675 As we have seen, from comparing the results of the L55 VFD model with
676 those of the other models, and from calculations with partially increased ver-
677 tical resolution, it appears that the difference between the occurrence/non-
678 occurrence of QBO-like oscillations is strongly related to the formation/absence
679 of a strong steady eastward zonal wind near or above 50 hPa. To investigate
680 the influence of this eastward wind region on the occurrence of QBO-like os-
681 cillations, an experiment is performed in which additional Rayleigh friction
682 is applied to weaken the eastward wind in the L110 VFD model. Specifically,

683 the friction is applied to the zonal mean zonal wind only as follows:

$$\frac{\partial \bar{u}}{\partial t} = -k_0 K_v(\sigma) K_h(\phi) \bar{u}(\phi, \sigma, t) + (\text{the other standard terms}), \quad (17)$$

684

$$K_v(\sigma) = \begin{cases} \frac{1}{2} \left[\tanh\left(\frac{z_* - z_b}{H}\right) + \tanh\left(-\frac{z_* - z_t}{H}\right) \right] & (0.02 < \sigma < 0.05) \\ 0 & (\text{otherwise}). \end{cases}, \quad (18)$$

685

$$K_h(\phi) = \begin{cases} \frac{1}{2} \left[\tanh\left(\frac{\sin \phi + \sin \phi_0}{L}\right) + \tanh\left(-\frac{\sin \phi - \sin \phi_0}{L}\right) \right] & (-25^\circ < \phi < 25^\circ) \\ 0 & (\text{otherwise}). \end{cases}. \quad (19)$$

686 Here, k_0 is the Rayleigh friction coefficient, the value of which is set to
 687 $1/3 \text{ day}^{-1}$, the same as for the sponge layer, and $z_* = -h_0 \log \sigma$ is an
 688 approximate logarithmic pressure height, where $h_0 = 7 \text{ km}$ is the scale
 689 height. The values of the parameters are set as follows: $z_b = 21.5 \text{ km}$,
 690 $z_t = 26.5 \text{ km}$, $H = 1 \text{ km}$, $\phi_0 = \arcsin 0.3$, and $L = 1/15$. With such a setup,
 691 the zonal mean zonal wind is selectively weakened at $-25^\circ < \phi < 25^\circ$ and
 692 $0.02 < \sigma < 0.05$, but the wave components ($m \neq 0$) are not affected. Note
 693 that in (17), the part written as “(the other standard terms)” is the zonal
 694 mean of the terms included in the time evolution of u in the original model
 695 equations.

Fig. 15

696 The result of day 3600 of the standard L110 VFD model without the
697 additional Rayleigh friction term is used as an initial state, then the time
698 evolution is calculated thereafter with the friction term. The resulting time
699 evolution of the equatorial zonal wind (Fig. 15) shows that, above the alti-
700 tude range where the additional Rayleigh friction term is added, the zonal
701 wind phase slowly descends for the first four years or so, after which the
702 zonal mean zonal wind profile settles down to an almost steady state. In
703 this case, it appears that the strong eastward zonal wind is eventually main-
704 tained above the altitude region where the additional Rayleigh friction term
705 is added, thereby preventing the occurrence of the QBO-like oscillations in
706 the higher altitude region. This result, together with the results of the L110
707 VFD model and those of Section 4.2, suggests that the occurrence/non-
708 occurrence of the QBO-like oscillations is strongly related to the formation
709 of a strong eastward wind region: once a strong eastward wind region is
710 formed, the upward propagation of eastward moving waves is inhibited,
711 suppressing the occurrence of the QBO-like oscillations.

712 5. Summary and discussion

713 5.1 *Summary*

714 In the present manuscript, with reference to YJ2015, we investigated
715 whether QBO-like oscillations occur spontaneously in dry GCMs in a bench-
716 mark experimental setting of the Held and Suarez (1994) with an addi-
717 tional sponge layer, by performing several numerical experiments with dif-
718 ferent vertical discretization methods (finite difference method and spectral
719 method) and different vertical resolutions. As a result, the model that uses
720 the finite difference method for vertical discretization (VFD model) gener-
721 ated QBO-like oscillations when the number of vertical layers was set to
722 the same number as in YJ2015 (L55), but did not generate QBO-like oscil-
723 lations when it was doubled (L110). The model using the spectral method
724 for vertical discretization (3DS model) did not generate QBO-like oscilla-
725 tions for either of the two vertical resolution settings used. These results
726 suggest that the occurrence or non-occurrence of the QBO-like oscillation
727 in the dry GCM with the YJ2015 setting is strongly influenced by the ver-
728 tical discretization method and the vertical resolution of the model. This is
729 consistent with the fact that QBO-like oscillations were observed in three of
730 the four types of dynamical cores used in YJ2015 that used finite difference
731 discretization in the vertical direction, but not in the FV model, which used

732 floating vertical coordinates.

733 To investigate what determines the occurrence or non-occurrence of
734 QBO-like oscillations, wavenumber frequency spectral analyses of the tem-
735 perature field over the tropics were also performed. As shown in Figs. 5 and
736 6, the wave activity in the L55 VFD model was significantly higher than
737 in the other models at $\sigma = 0.002$, but there was no significant difference
738 among the models at $\sigma = 0.1$. This result suggests that the strength of the
739 upward propagating wave entering from below $\sigma = 0.1$, where the profiles of
740 the zonal mean zonal wind are similar for all the models, is not significantly
741 different among the models.

742 The wave contribution to the vertical momentum flux $\overline{u'\omega'}$ in the pressure-
743 zonal phase velocity space was also investigated in order to clarify the reason
744 for the different wave activities at $\sigma = 0.002$ (Figs. 7 and 8). For all the
745 models except for the L55 VFD model, there was a sharp suppression of the
746 upward penetration of eastward moving waves in the eastward zonal wind
747 region around 50 hPa. The L55 VFD model result also showed such sup-
748 pression around 20 hPa, but the eastward wind was weaker than those for
749 the other models, resulting in weaker suppression of the upward penetration
750 of eastward moving waves. This weaker suppression allowed the eastward
751 moving waves to propagate to higher altitudes, and this is thought to be one
752 reason why the L55 VFD model showed strong wave activity at $\sigma = 0.002$.

753 Transformed Eulerian Mean (TEM) analyses were performed to quanti-
754 tatively investigate the causes of the different zonal wind profiles obtained
755 in the L55 and L110 VFD models. In the L55 VFD model, eastward ac-
756 celerations due to the advection term mainly contributed to the descent of
757 the eastward phase, and westward accelerations due to the EP flux con-
758 vergence contributed to the descent of the westward phase. The strong
759 eastward acceleration in the L55 VFD model was produced by the predomi-
760 nant downward motion in the equatorial region, enhanced by the downward
761 secondary circulation induced by the vertical shear of the zonal wind. In
762 the L110 VFD model, there were strong eastward accelerations due to the
763 advection term, but these accelerations were almost completely canceled out
764 by the acceleration due to the EP flux convergence, resulting in an almost
765 steady zonal wind. Note that the acceleration feature due to the advection
766 term and the meridional EP flux divergence are similar between the L55
767 and the L110 VFD models. It is the vertical EP flux divergence in the
768 eastward shear region that is mainly different: in the L55 VFD model, the
769 eastward acceleration due to the vertical EP flux divergence is stronger than
770 in the L110 VFD model. This is consistent with the vertical momentum flux
771 analysis, where the L55 VFD model showed stronger upward propagation
772 of eastward moving waves.

773 The results seen above suggested that in the L55 VFD model and the

774 L110 VFD model, the difference in the formation of the eastward zonal wind
775 around $0.01 < \sigma < 0.05$ caused the differences in the upward penetration
776 of eastward propagating waves, which in turn caused the differences in the
777 occurrence of the QBO-like oscillation. To clarify the effect of this east-
778 ward zonal wind and the vertical resolution on the QBO-like oscillation,
779 we performed two additional experiments on this point. The first addi-
780 tional experiment was to double the density of the vertical grids of the L55
781 VFD model in certain altitude ranges, rather than simply doubling the to-
782 tal number of layers. Doubling the density of the vertical grid in the range
783 of $0.028 < \sigma < 0.12$ and increasing the total number of layers to 62, i.e.,
784 increasing the total number of layers by only 7 from the default setting
785 of L55, resulted in the formation of a steady eastward zonal wind around
786 $0.01 < \sigma < 0.05$ and the elimination of QBO-like oscillations (Fig. 14b).
787 Note that when a steady and strong eastward wind region (above 16 m/s)
788 formed, QBO-like oscillations did not occur, but when this steady eastward
789 wind was weak (around 8 m/s), QBO-like oscillations occurred in the coarse
790 vertical resolution region above this eastward wind region. The second ad-
791 ditional experiment was to apply Rayleigh friction at an altitude range in
792 the L110 VFD model to remove the eastward steady zonal wind around
793 $0.02 < \sigma < 0.05$. With this setting, the pattern of the zonal wind profile
794 above the altitude where the additional Rayleigh friction was applied prop-

795 agated downwards for the first four years, but then a strong eastward zonal
796 wind was formed just above the additional Rayleigh friction layer and the
797 zonal wind profile became almost steady (Fig. 15). These two additional
798 experiments clarified that steady eastward zonal wind was formed at a cer-
799 tain altitude when the vertical resolution of the model was high (including
800 the partially high case), and that once it was formed and strong, QBO-like
801 oscillations above it were suppressed.

802 From the above, the conclusion of the present manuscript is that the
803 QBO-like oscillations generated by the dry GCMs according to the setting of
804 YJ2015 will occur only when the vertical resolution is coarse or the vertical
805 discretization error is large in the altitude region higher than the altitude
806 corresponding to the tropopause, and when the resolution is sufficiently
807 high, strong eastward zonal winds will form somewhere in this altitude
808 region, which will suppress the vertical propagation of eastward waves and
809 thus prevent the QBO-like oscillations from occurring.

810 *5.2 Discussion*

811 As we have seen, it is clear that the differences in the wave activity
812 among the models in the altitude region where QBO-like oscillations are
813 observed in the L55 VFD model are primarily due to differences in the un-
814 derlying zonal mean zonal wind, but why the formation of the zonal mean

815 zonal wind depends on the vertical resolution of the model was not fully
816 explored in this manuscript and is a topic for future work. However, one
817 speculation for the reason is that the difference in vertical resolution caused
818 a difference in the wave-mean-flow interaction: In the YJ2015 experimental
819 setup, no explicit vertical diffusion is used, and it is only the nonlinear in-
820 teraction that contributes to the damping of the waves with short vertical
821 wavelength near the critical level. Therefore, increasing the vertical resolu-
822 tion can lead to a more correct representation of this nonlinear interaction,
823 leading to a larger acceleration of the zonal wind in the L110 VFD and 3DS
824 models. This difference in the representation of the vertical propagation of
825 waves depending on the vertical resolution can come into play if the vertical
826 wavelength of the wave is shortened, even if the waves do not necessarily
827 reach the critical level. Consistent with this argument, a comparison of the
828 wave contribution to the $\overline{u'\omega'}$ in the L110 VFD model (Fig. 7c) and in the
829 eastward phase of the L55 VFD model (Fig. 8b) shows that the vertical
830 penetration of fast phase velocity waves that do not reach the critical level
831 is sharply suppressed in the L110 VFD model (around 5 hPa), while the
832 suppression is less significant in the L55 VFD model (around 5 hPa). This
833 trend of smaller vertical momentum fluxes due to waves in the higher vertical
834 resolution setting was also observed in a realistic GCM study (Watanabe et
835 al., 2015) and may be consistent with the results of the present manuscript.

836 However, of course, note that the model used here is a bare dynamical core
837 and is far from a realistic GCM. As seen above, it seems certain that the
838 difference in vertical resolution alters the representation of the filtering ef-
839 fect of zonal winds, and affects the formation of the zonal winds themselves
840 and the vertical propagation of the waves. However, comparing the results
841 between the L55 VFD model and the L110 VFD model (Figs. 7a and 7c),
842 the vertical momentum flux due to eastward propagating waves with fast
843 phase velocities (e.g. 32 m s^{-1}) at about 20 hPa is larger in the L55 VFD
844 model. This difference may not be fully explained by the above mechanism
845 alone.

846 Another possible factor that could explain the difference in wave activity
847 with different vertical discretization methods or different vertical resolutions
848 could be overestimations of the waves connected to the vertical discretiza-
849 tion methods, which is a numerical artifact as mentioned in YJ2015. We
850 suspect that one mechanism by which waves could be excited due to vertical
851 discretization errors is as follows: For example, when a physical quantity
852 such as potential vorticity is adiabatically advected, if the isentropic sur-
853 face is tilted east-west, the error in the vertical partial differentiation of the
854 physical quantity may increase the error in the zonal advection calculation.
855 This leads to a deviation from the geostrophically balanced state, which
856 may lead to gravity wave excitation. Since the Brunt-Väisälä frequency in

857 the tropics has larger values in $\sigma < 0.1$ than in the lower levels, gravity
858 wave excitation is more likely to occur when the atmosphere is perturbed in
859 this way. Related to this point, Lindzen and Fox-Rabinovitz (1989) showed
860 that the vertical grid spacing required to resolve balanced quasi-geostrophic
861 dynamics and gravity wave modes is inversely proportional to the Brunt-
862 Väisälä frequency. Therefore, high Brunt-Väisälä frequencies in the $\sigma < 0.1$
863 region are more likely to produce large discretization errors, and this point
864 may be important, not only for the artificial wave generation issue but for
865 the overall accuracy in resolving flows. It should be noted, however, that no
866 clear evidence of the wave excitation due to vertical discretization error is
867 obtained in this study and these points, together with the speculation that
868 the representation of the wave-mean-flow interaction may be altered by the
869 vertical resolution, require further investigation .

870 In addition to the need for further investigation of the relationship
871 among zonal wind, wave excitation, and vertical resolution, it should be
872 noted here that there is also a need to investigate the occurrence/non-
873 occurrence of QBO-like oscillations in different horizontal resolution set-
874 tings, since all experiments in this manuscript are conducted in T63 hor-
875 izontal resolution and horizontal resolution setting could also change the
876 occurrence/non-occurrence of QBO-like oscillations. For example, Kawatani
877 et al. (2010) showed that the westward acceleration of the QBO is mainly

878 driven by small-scale gravity waves in the lower stratosphere. In the present
879 manuscript, the horizontal resolution is T63 and such small-scale waves are
880 not considered from the beginning. Therefore, if the horizontal resolution
881 is higher, the westward acceleration may be strong. Indeed, in the L55
882 VFD model, the westward wind regions only descend to about $\sigma = 0.01$,
883 which is a higher altitude than the altitude to which the eastward wind
884 regions descend. This result should reflect the weaker westward forcing in
885 $0.01 < \sigma < 0.02$ (corresponding to the lower stratosphere) due to the ab-
886 sence of small-scale gravity waves. However, in the real atmosphere such
887 small-scale waves are thought to be mainly excited by moist convection.
888 Therefore, whether such waves can be strongly excited by dry GCMs is
889 another question to be investigated in our future work.

890 **Data Availability Statements**

891 The datasets generated and analyzed in the present manuscript are avail-
892 able from the corresponding author on reasonable request.

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904 The GFD-DENNOU Library (<https://www.gfd-dennou.org/arch/dcl/>)
905 was used to draw the figures.

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List of Figures

969	1	Vertical grid distributions for VFD models (L55 and L110)	
970		and 3DS models (M85 and M170). In the top panel, σ is	
971		shown on a linear scale, whereas in the bottom panel, σ is	
972		shown on a logarithmic scale.	55
973	2	Time- σ cross section of zonal mean zonal wind averaged in	
974		2°S – 2°N (unit is m s^{-1}). The monthly (30-day) mean data	
975		after the spin-up period is shown. The horizontal axis is the	
976		model year (360 days) and the vertical axis is σ . Contour	
977		interval is 4 m s^{-1} . (a): L55 VFD, (b): M85 3DS. (c): L110	
978		VFD, and (d): M170 3DS model.	56
979	3	Latitude- σ cross section of zonal and monthly mean zonal	
980		wind, where the average is taken over day 2970–3000 (unit is	
981		m s^{-1}). The horizontal axis is latitude and the vertical axis	
982		is σ . Contour interval is 4 m s^{-1} . (a): L55 VFD, (b): M85	
983		3DS. (c): L110 VFD, and (d): M170 3DS model.	57
984	4	Latitude- σ cross section of Brunt-Väisälä frequency of zonal	
985		mean field at day 3000 in the time evolution of the L55 VFD	
986		model (unit is s^{-1}). The horizontal axis is latitude and the	
987		vertical axis is σ . Contour interval is $2.5 \times 10^{-3} \text{ s}^{-1}$	58

988	5	Wavenumber-frequency spectra of the temperature field $P_{s/a}$ at $\sigma = 0.002$ averaged over 15°S – 15°S . The units are K^2 and drawn in their common logarithm. (a) and (b): L55 VFD, (c) and (d): M85 3DS, (e) and (f): L110 VFD, (g) and (h): M170 3DS model. The left panel of each pair is for the symmetric components and the right one is for the anti-symmetric components. The dispersion curves are drawn overlaid with the following three equivalent depths (h): $h = 200$ m (solid line), 50 m (dashed line), and 12 m (dotted line). For the symmetric component, the eastward dispersion curve through the origin is for Kelvin waves; for the antisymmetric component, the dispersion curve leading from a relatively high frequency eastward to a low frequency westward is for Rossby-gravity waves; the relatively low frequency ($\nu < 0.15$ cpd) westward component of the symmetric and antisymmetric correspond to Rossby waves, and parabolic dispersion curves at relatively high frequencies ($\nu > 0.3$ cpd) in the symmetric and antisymmetric components correspond to gravity waves. . . .	59
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1003			
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1005			
1006	6	Same as Fig. 5, but results for $\sigma = 0.1$ are shown.	60
1007	7	Zonal phase velocity-pressure distributions of the vertical momentum flux $C_{\overline{u'\omega'}}$ in day 3600–4500 averaged over 2°N – 2°S . The units are m Pa s^{-2} . (a): L55 VFD, (b): M85 3DS, (c): L110 VFD, (d): M170 3DS model. The solid curves show the zonal mean zonal velocity profiles averaged over the corresponding time and latitude. The value of $C_{\overline{u'\omega'}}$ shown here is multiplied by 2×10^7 . The color gradations are drawn on a logarithmic scale.	61
1008			
1009			
1010			
1011			
1012			
1013			
1014			
1015	8	Same as Fig. 7, but analyses are conducted using 120 days data of the L55 VFD model. (a): Westward phase around $p = 5$ hPa for days 3960–4080, (b): Eastward phase around $p = 5$ hPa for days 4380–4500.	62
1016			
1017			
1018			

1019	9	Time-pressure cross section of zonal and monthly mean zonal wind accelerations due to each TEM component, averaged over 2°N–2°S. (a) and (b): total acceleration, (c) and (d): advection term, (e) and (f): EP flux divergence, (g) and (h): residual term, (i) and (j): residual term minus diffusion. The left panel of each row is for the L55 VFD model and the right panel of each row is for the L110 VFD model. The units are $\text{m s}^{-1} \text{ day}^{-1}$. The contours represent the zonal and monthly mean zonal wind averaged over the corresponding latitudes. The contour interval is 8 m s^{-1}	63
1020			
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1024			
1025			
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1029	10	The EP flux divergence terms shown in Figs. 9e and 9f are shown split into the meridional component and the vertical component. (a) and (b): the meridional component, (c) and (d): the vertical component.	64
1030			
1031			
1032			
1033	11	Latitude-pressure cross section of zonal mean residual circulation (vectors), zonal mean zonal velocity (contours), and vertical velocity of residual circulation (color map), averaged over days 3600–4680. (a): L55 VFD, (b): L110 VFD model. The unit vectors represent 0.1 m s^{-1} (meridional) and 0.2 mm s^{-1} (vertical). The units of the color maps are m s^{-1} . The vertical component of residual circulation is converted from the pressure velocity to the geometric velocity, using the approximate relationship (see the text). The contour interval is 8 m s^{-1}	65
1034			
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1038			
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1040			
1041			
1042			
1043	12	Time evolution of zonal mean zonal velocity (dashed curve, left axis) and residual vertical pressure velocity (solid blue curve, right axis) averaged over 2°N–2°S at 4.9hPa. The result of the L55 VFD model is shown.	66
1044			
1045			
1046			
1047	13	Same as Fig. 2a, but for the L55 VFD model with setting the hyperdiffusion and time integration scheme being the same as EUL in YJ2015.	67
1048			
1049			
1050	14	Same as Fig. 2a, but for the results for partially higher vertical resolutions. The number of vertical grid points are doubled in the following σ -ranges. (a): $0.12 < \sigma < 0.49$, (b): $0.028 < \sigma < 0.12$, (c): $0.028 < \sigma < 0.49$, (d): $0.0067 < \sigma < 0.49$, and (e): $0 < \sigma < 0.49$	68
1051			
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1055 15 Same as Fig. 2c, but for the results for the L110 VFD model
1056 with an additional Rayleigh friction over a specific range of
1057 altitudes and latitudes ($0.02 < \sigma < 0.05$ and $-25^\circ < \phi <$
1058 25°). The result of day 3600 of the L110 VFD model is used
1059 as the initial state, which corresponds to the leftmost year 10
1060 in this figure. 69

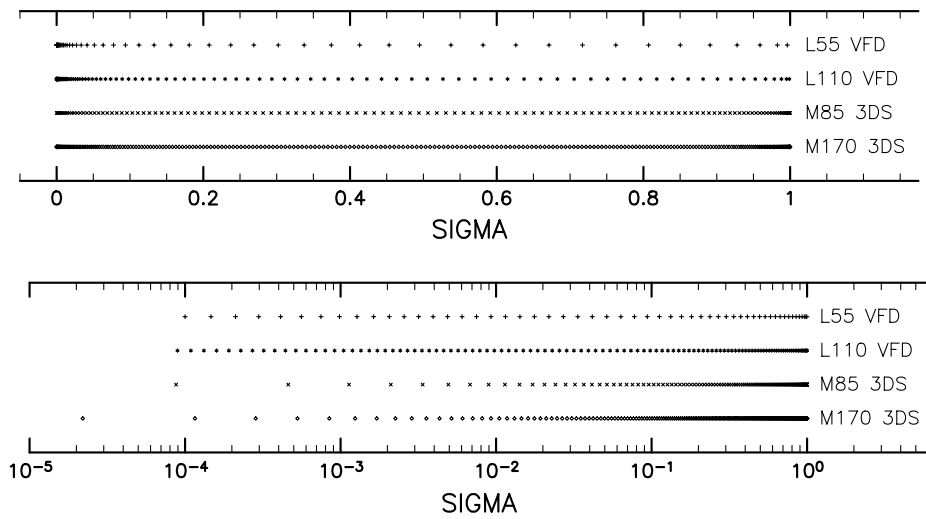


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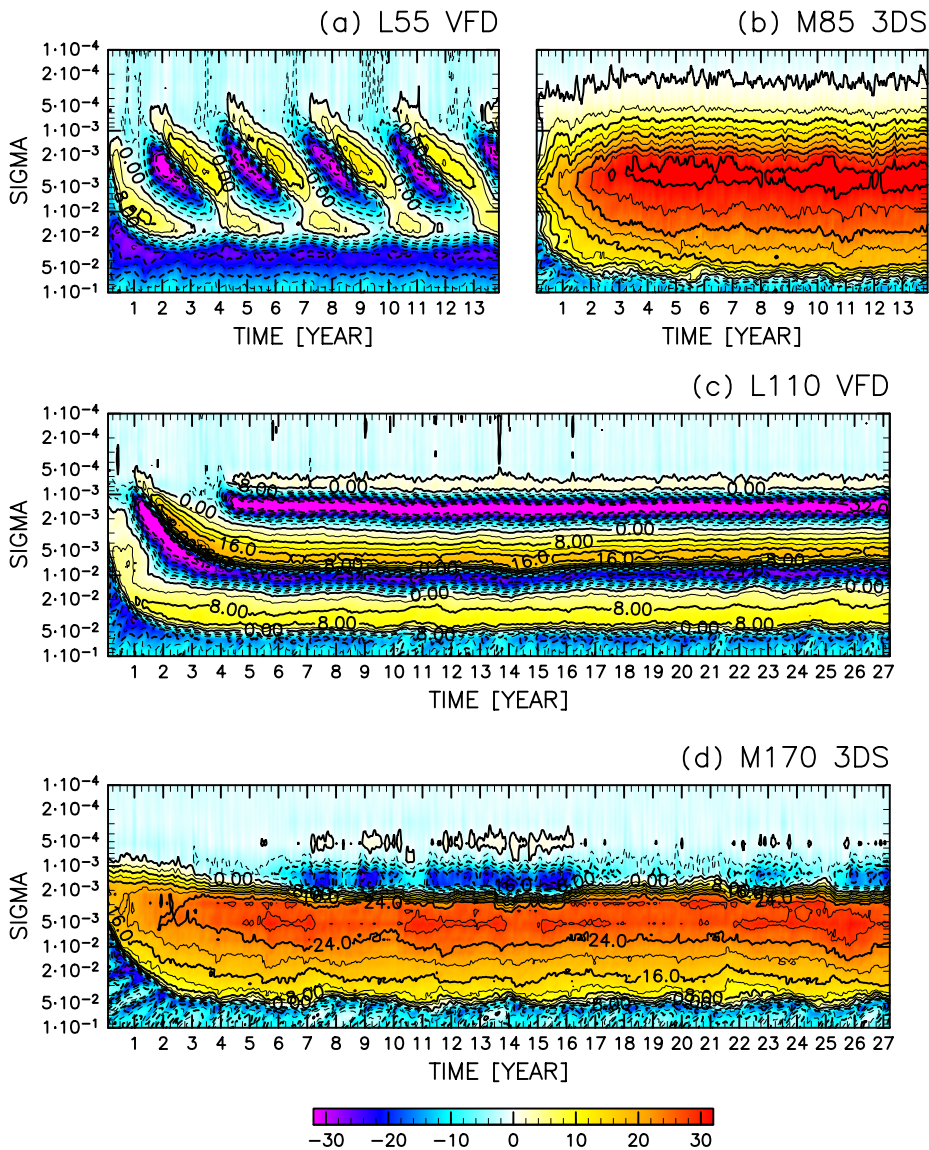


Fig. 2. Time- σ cross section of zonal mean zonal wind averaged in 2°S – 2°N (unit is m s^{-1}). The monthly (30-day) mean data after the spin-up period is shown. The horizontal axis is the model year (360 days) and the vertical axis is σ . Contour interval is 4 m s^{-1} . (a): L55 VFD, (b): M85 3DS. (c): L110 VFD, and (d): M170 3DS model.

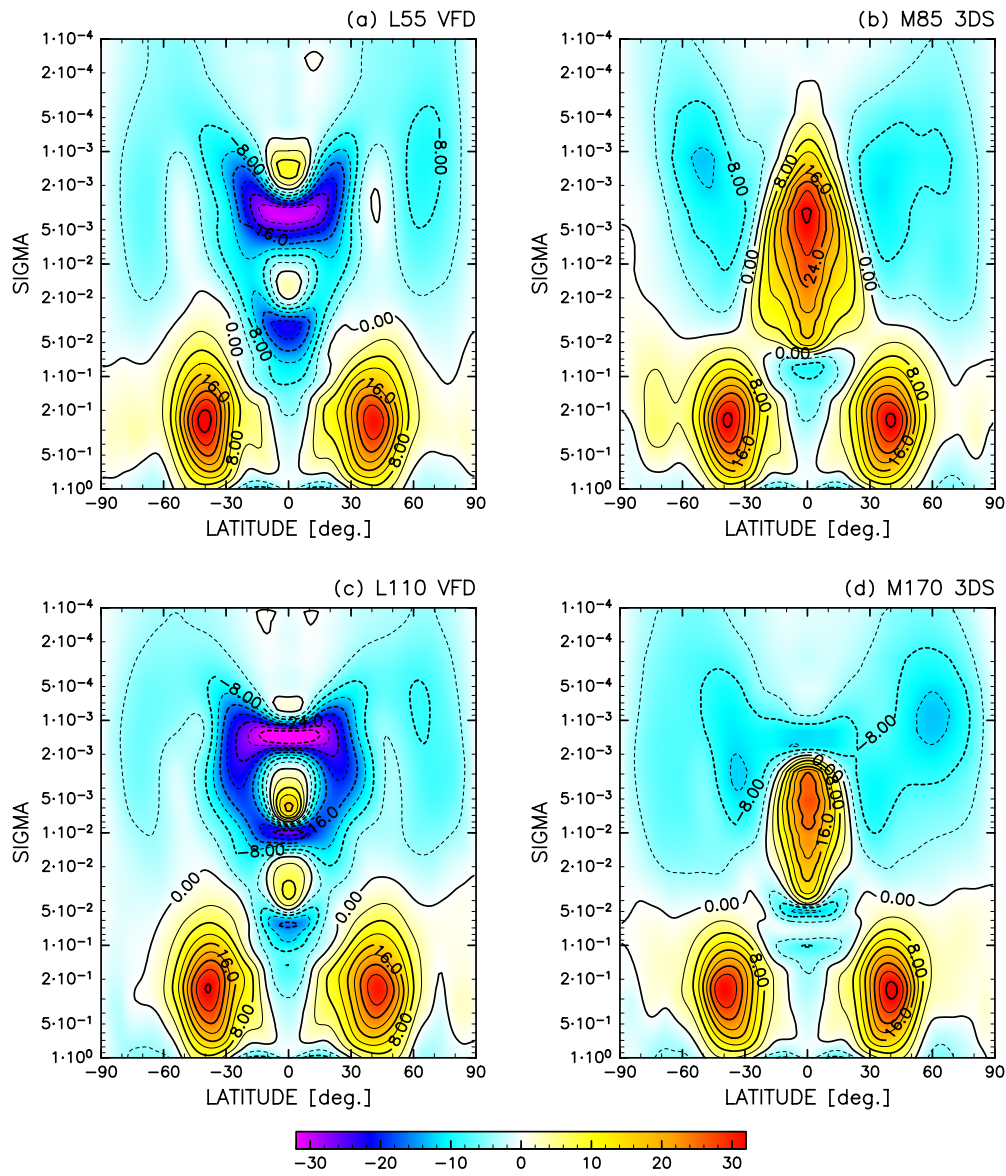


Fig. 3. Latitude- σ cross section of zonal and monthly mean zonal wind, where the average is taken over day 2970–3000 (unit is m s^{-1}). The horizontal axis is latitude and the vertical axis is σ . Contour interval is 4 m s^{-1} . (a): L55 VFD, (b): M85 3DS. (c): L110 VFD, and (d): M170 3DS model.

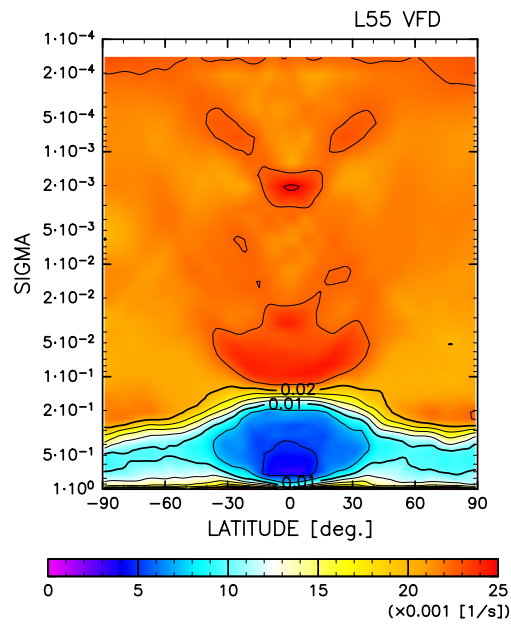


Fig. 4. Latitude- σ cross section of Brunt-Väisälä frequency of zonal mean field at day 3000 in the time evolution of the L55 VFD model (unit is s^{-1}). The horizontal axis is latitude and the vertical axis is σ . Contour interval is $2.5 \times 10^{-3} \text{ s}^{-1}$.

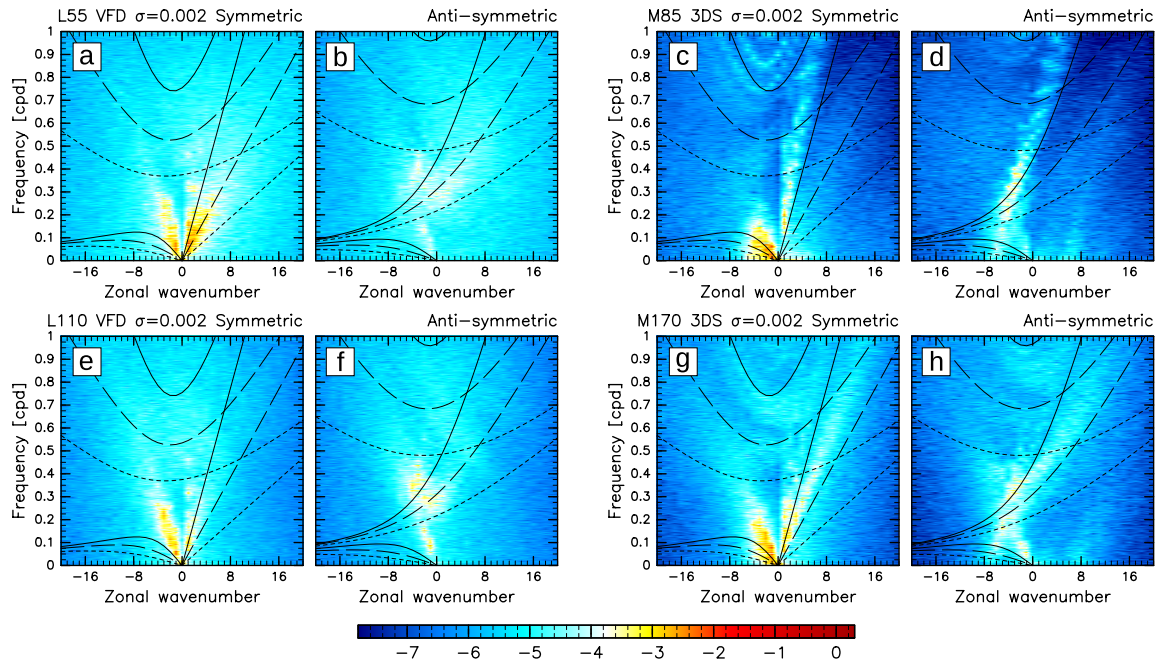


Fig. 5. Wavenumber-frequency spectra of the temperature field $P_{s/a}$ at $\sigma = 0.002$ averaged over 15°S – 15°S . The units are K^2 and drawn in their common logarithm. (a) and (b): L55 VFD, (c) and (d): M85 3DS, (e) and (f): L110 VFD, (g) and (h): M170 3DS model. The left panel of each pair is for the symmetric components and the right one is for the anti-symmetric components. The dispersion curves are drawn overlaid with the following three equivalent depths (h): $h = 200$ m (solid line), 50 m (dashed line), and 12 m (dotted line). For the symmetric component, the eastward dispersion curve through the origin is for Kelvin waves; for the antisymmetric component, the dispersion curve leading from a relatively high frequency eastward to a low frequency westward is for Rossby-gravity waves; the relatively low frequency ($\nu < 0.15$ cpd) westward component of the symmetric and antisymmetric correspond to Rossby waves, and parabolic dispersion curves at relatively high frequencies ($\nu > 0.3$ cpd) in the symmetric and antisymmetric components correspond to gravity waves.

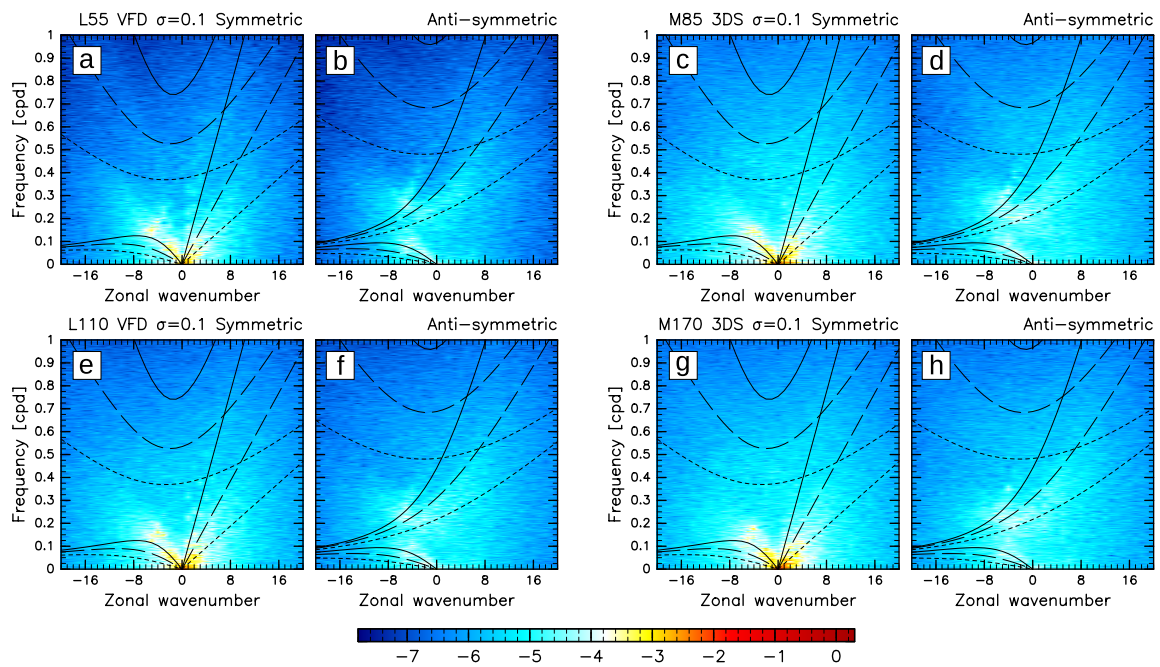


Fig. 6. Same as Fig. 5, but results for $\sigma = 0.1$ are shown.

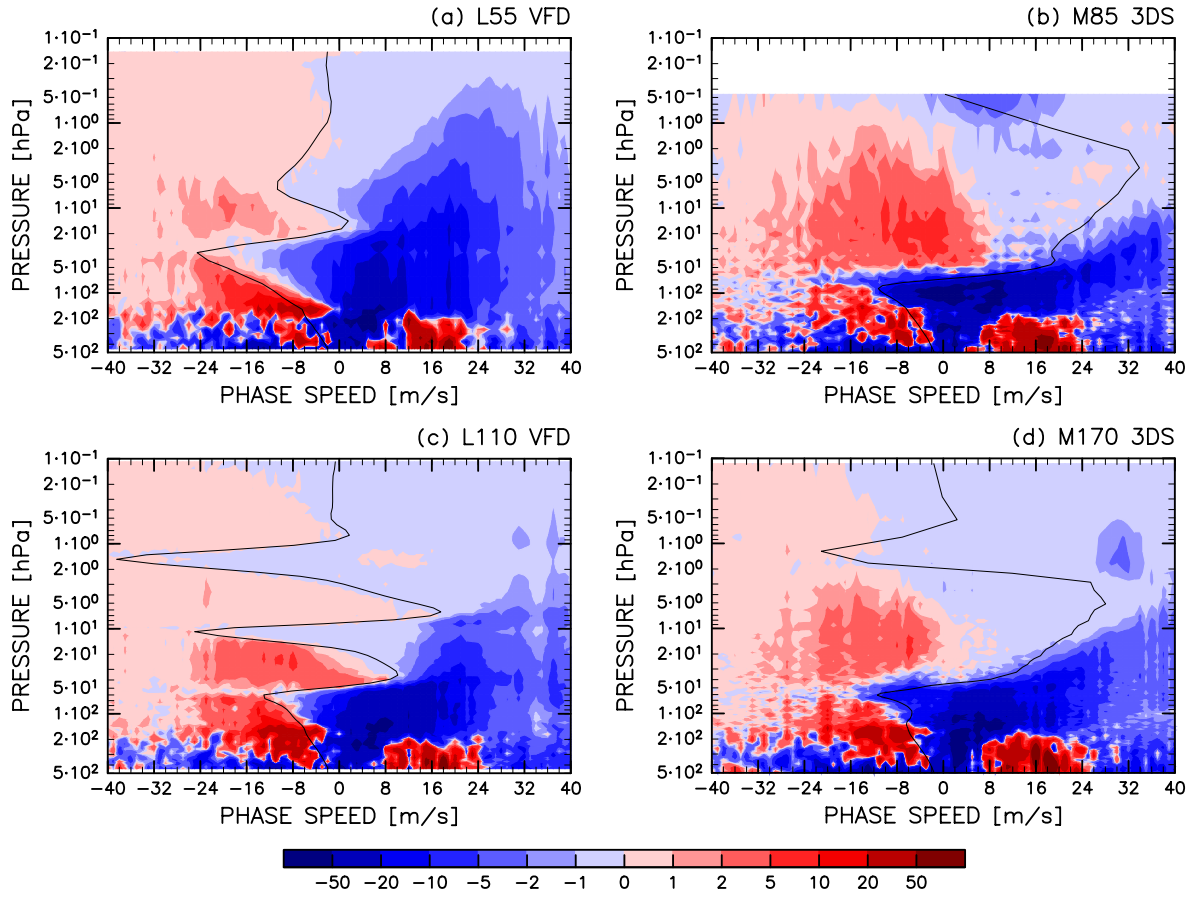


Fig. 7. Zonal phase velocity-pressure distributions of the vertical momentum flux $C'_{u'\omega'}$ in day 3600–4500 averaged over 2°N – 2°S . The units are m Pa s^{-2} . (a): L55 VFD, (b): M85 3DS, (c): L110 VFD, (d): M170 3DS model. The solid curves show the zonal mean zonal velocity profiles averaged over the corresponding time and latitude. The value of $C'_{u'\omega'}$ shown here is multiplied by 2×10^7 . The color gradations are drawn on a logarithmic scale.

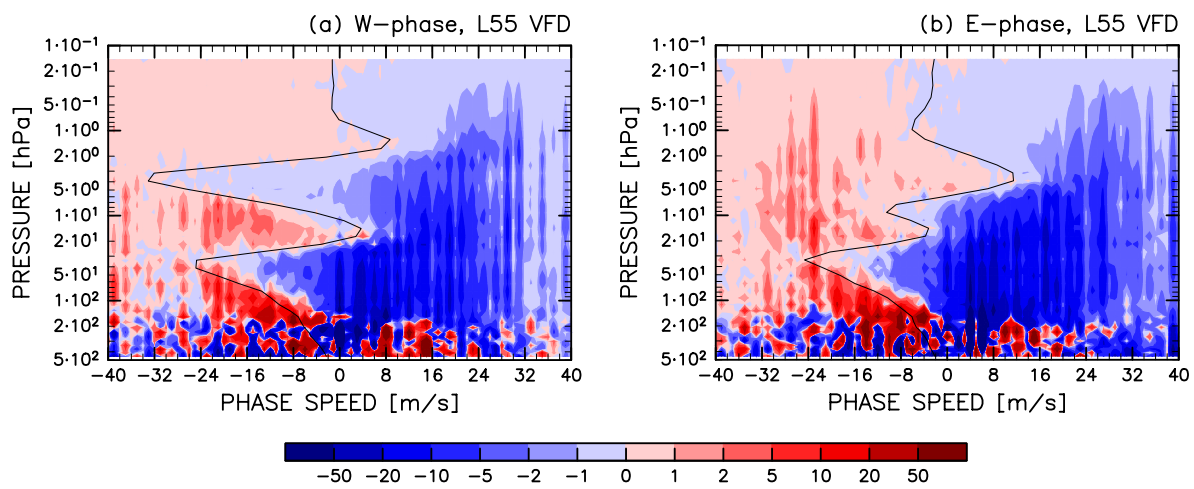


Fig. 8. Same as Fig. 7, but analyses are conducted using 120 days data of the L55 VFD model. (a): Westward phase around $p = 5$ hPa for days 3960–4080, (b): Eastward phase around $p = 5$ hPa for days 4380–4500.

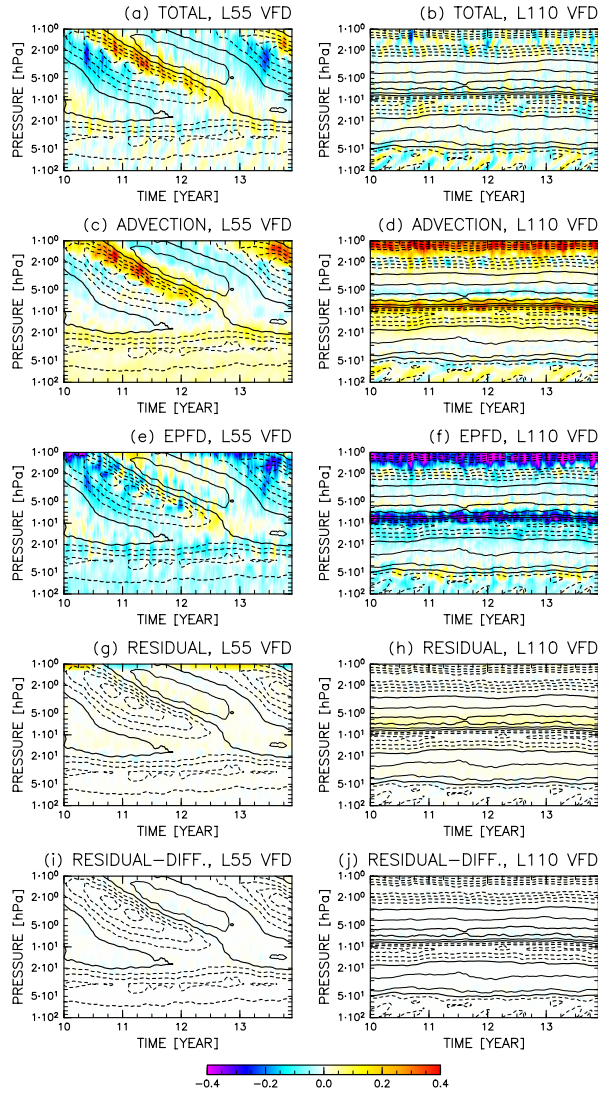


Fig. 9. Time-pressure cross section of zonal and monthly mean zonal wind accelerations due to each TEM component, averaged over 2°N – 2°S . (a) and (b): total acceleration, (c) and (d): advection term, (e) and (f): EP flux divergence, (g) and (h): residual term, (i) and (j): residual term minus diffusion. The left panel of each row is for the L55 VFD model and the right panel of each row is for the L110 VFD model. The units are $\text{m s}^{-1} \text{day}^{-1}$. The contours represent the zonal and monthly mean zonal wind averaged over the corresponding latitudes. The contour interval is 8 m s^{-1} .

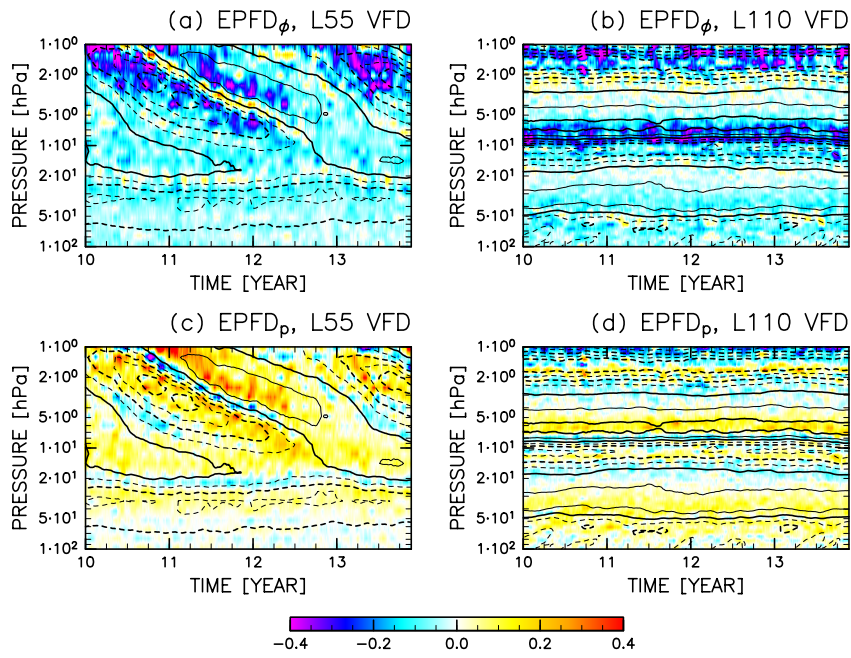


Fig. 10. The EP flux divergence terms shown in Figs. 9e and 9f are shown split into the meridional component and the vertical component. (a) and (b): the meridional component, (c) and (d): the vertical component.

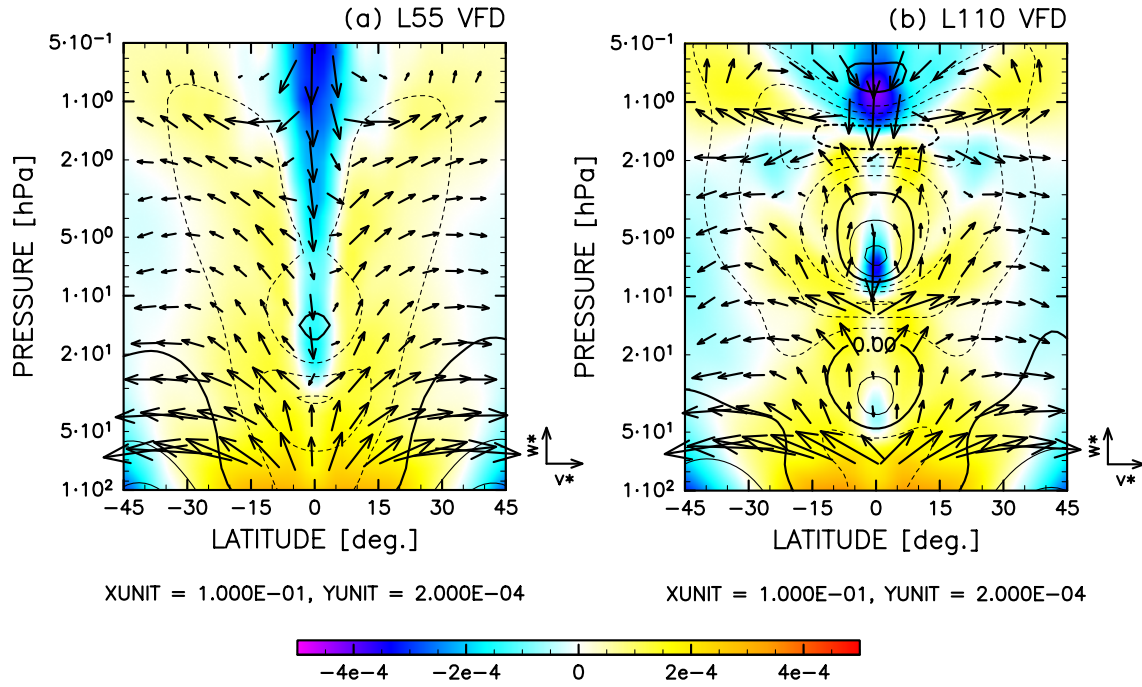


Fig. 11. Latitude-pressure cross section of zonal mean residual circulation (vectors), zonal mean zonal velocity (contours), and vertical velocity of residual circulation (color map), averaged over days 3600–4680. (a): L55 VFD, (b): L110 VFD model. The unit vectors represent 0.1 m s^{-1} (meridional) and 0.2 mm s^{-1} (vertical). The units of the color maps are m s^{-1} . The vertical component of residual circulation is converted from the pressure velocity to the geometric velocity, using the approximate relationship (see the text). The contour interval is 8 m s^{-1} .

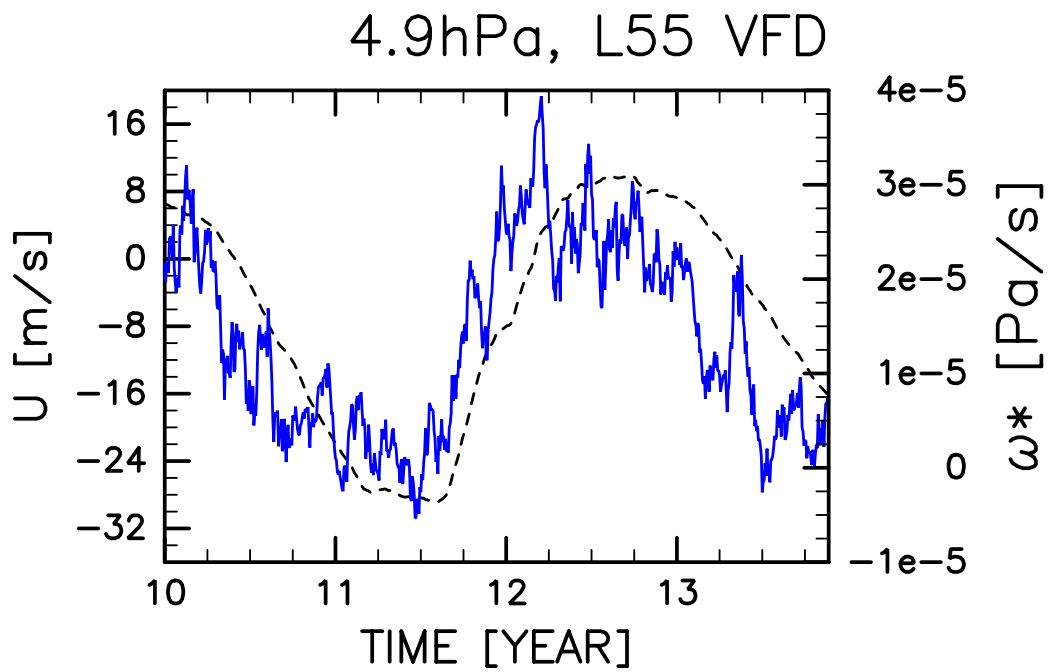


Fig. 12. Time evolution of zonal mean zonal velocity (dashed curve, left axis) and residual vertical pressure velocity (solid blue curve, right axis) averaged over 2°N - 2°S at 4.9hPa. The result of the L55 VFD model is shown.

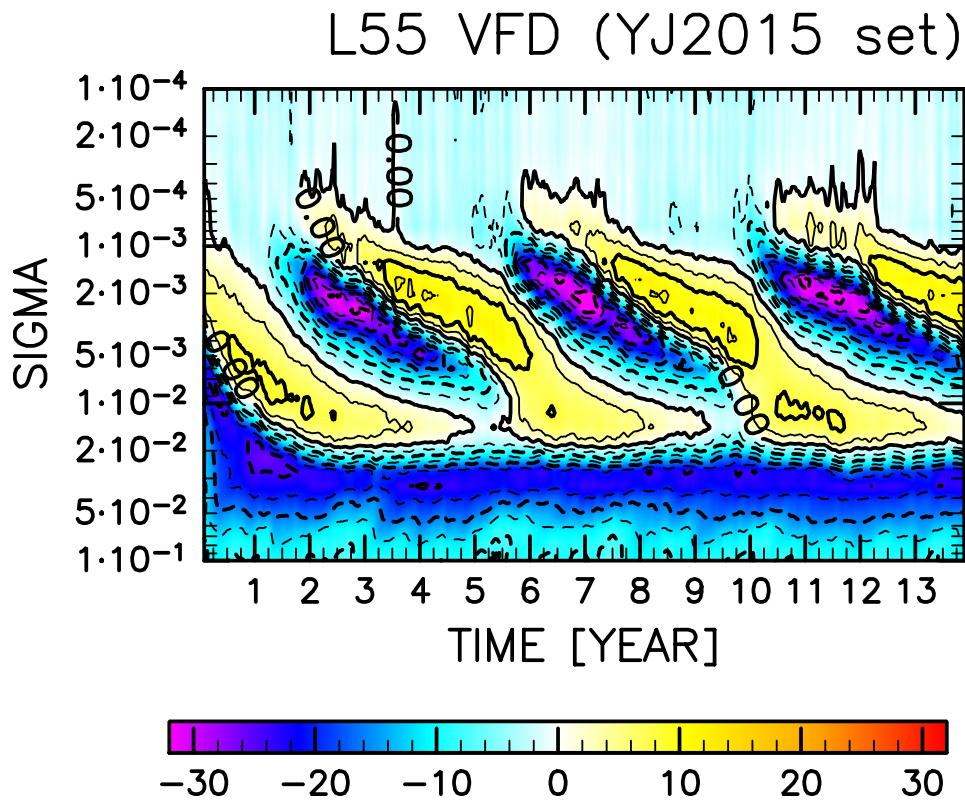


Fig. 13. Same as Fig. 2a, but for the L55 VFD model with setting the hyperdiffusion and time integration scheme being the same as EUL in YJ2015.

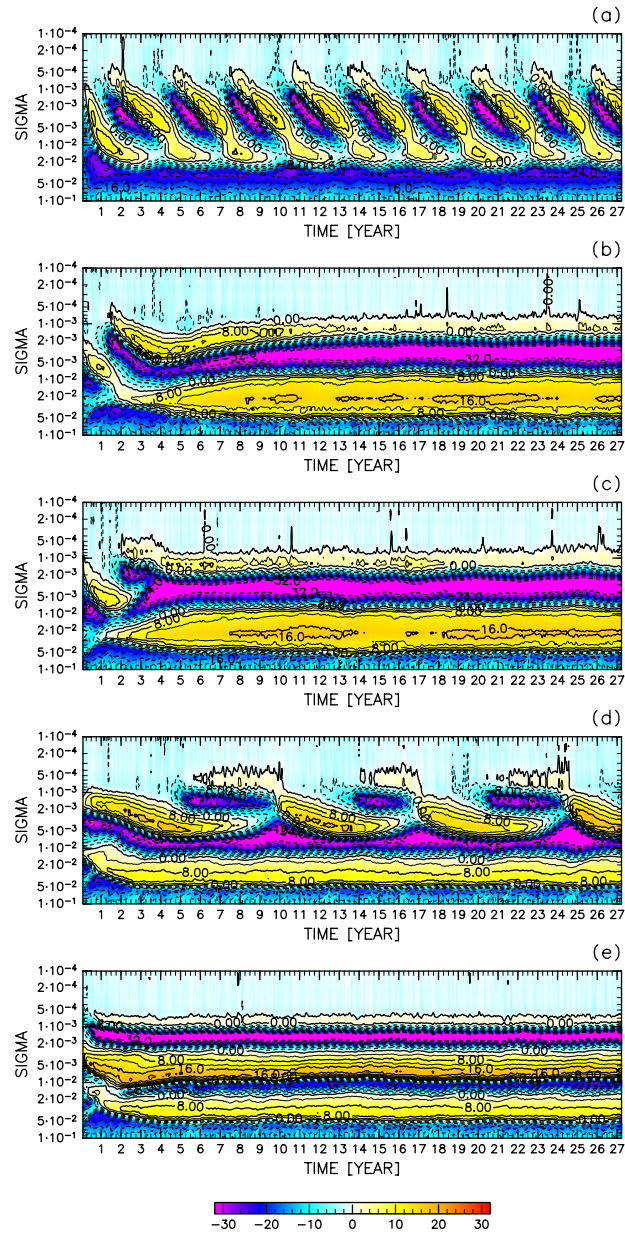


Fig. 14. Same as Fig. 2a, but for the results for partially higher vertical resolutions. The number of vertical grid points are doubled in the following σ -ranges. (a): $0.12 < \sigma < 0.49$, (b): $0.028 < \sigma < 0.12$, (c): $0.028 < \sigma < 0.49$, (d): $0.0067 < \sigma < 0.49$, and (e): $0 < \sigma < 0.49$.

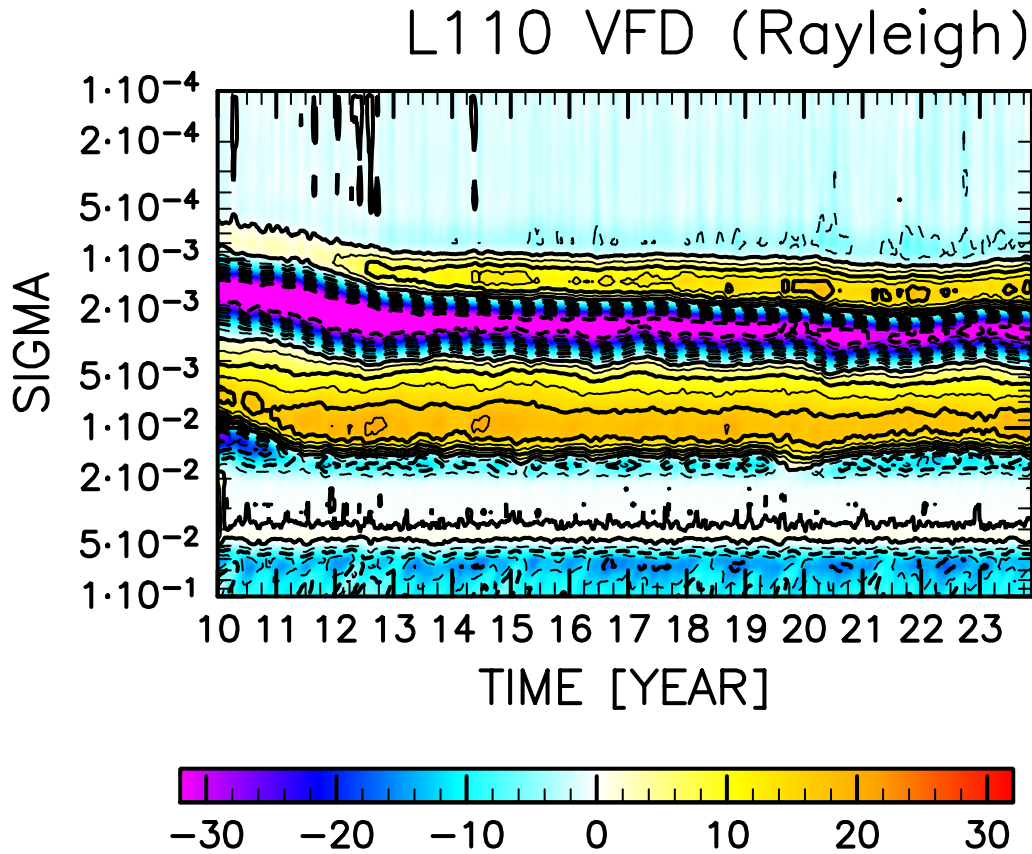


Fig. 15. Same as Fig. 2c, but for the results for the L110 VFD model with an additional Rayleigh friction over a specific range of altitudes and latitudes ($0.02 < \sigma < 0.05$ and $-25^\circ < \phi < 25^\circ$). The result of day 3600 of the L110 VFD model is used as the initial state, which corresponds to the leftmost year 10 in this figure.