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A Comparison Between SAR Wind Speeds and Western North Pacific Tropical Cyclone Best Track Estimates

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Abstract

32

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34 Spaceborne synthetic aperture radar (SAR) for measuring high winds is expected to
35 reduce uncertainties in tropical cyclone (TC) intensity and structure estimation, yet the
36 consistency of SAR observed winds equivalent to a 1-min sustained wind speed with the
37 conventionally estimated 10-min maximum wind speed ($V_{\max 10}$) remains to be
38 assessed. This study compares SAR wind observations with western North Pacific best
39 track estimates from the Japan Meteorological Agency (JMA) and the Joint Typhoon
40 Warning Center (JTWC). Because SAR wind observations have a bias dependent on
41 SAR incidence angle, a first order corrective term is proposed and used to correct SAR-
42 derived maximum wind (SAR V_{\max}) tentatively. After this correction, conversion of SAR
43 V_{\max} into SAR $V_{\max 10}$ with Dvorak conversion tables revealed a mean difference
44 between SAR $V_{\max 10}$ and JMA $V_{\max 10}$ ($\Delta V_{\max 10}$) of -0.1 m s^{-1} and a mean absolute
45 difference of 4.8 m s^{-1} . $\Delta V_{\max 10}$ is found to be correlated with current intensities and
46 future intensity changes. Also, comparison of the JMA best track 50-kt wind radius (R50)
47 with SAR wind speeds suggests that R50 is systematically underestimated. Aside from
48 the SAR wind limitations, possible reasons for the observed discrepancies between SAR
49 wind observations and best track estimates include biases in the Dvorak analysis and
50 conventional surface wind products. Further accumulation of SAR wind observations with
51 appropriate bias correction in the future is expected to contribute to a comprehensive

52 evaluation and improvement of conventional Vmax estimation methods, which could also
53 be useful to verify TC intensity forecasts.

54

55 **Keywords:** tropical cyclone; synthetic aperture radar; SAR; best track; Dvorak technique;

56 maximum wind

58 **1. Introduction**

59 Real-time observations and forecasts of violent winds associated with tropical cyclones
60 (TCs; see the Appendix for acronyms used in this paper) are essential for effective measures
61 to be taken for disaster prevention. Aircraft reconnaissance in the North Atlantic has played
62 an important role in monitoring high winds and TC structures and improving wind forecasts
63 (e.g., Zawislak et al. 2022). Until recently, however, there have been no satellite instruments
64 able to observe high winds with high spatial resolution under TC conditions (e.g., Knaff et
65 al. 2021). In general, horizontal resolutions of conventional satellite wind products are too
66 coarse, ~10–50 km (e.g., Reul et al. 2017; Mayers and Ruf 2020), to observe TC fine
67 structures. Also, because conventional scatterometers (e.g., the Advanced Scatterometer,
68 ASCAT) saturate at high wind speeds above 18 m s^{-1} (Chou et al. 2013), the highest wind
69 speeds are not observed. As a result, it has been difficult to verify the accuracy of best track
70 estimates (maximum wind speed (V_{max}), radius of maximum wind (RMW), etc.), and the
71 resulting uncertainty is a serious issue for TC monitoring and forecasting in areas where
72 there is no aircraft reconnaissance.

73 The advent of synthetic aperture radar (SAR) has led to a breakthrough in observing high
74 winds with high spatial resolution in the inner core of TCs (e.g., Mouche et al. 2017, 2019).
75 Conventional scatterometers equipped with a co-polarization microwave active sensor
76 observe the roughness of the ocean surface by emitting, for example, vertically (resp.
77 horizontally) polarized waves and receiving vertically (resp. horizontally) polarized waves

78 after backscattering by ocean surface waves. However, the co-polarized signal begins to
79 saturate or at least to decrease significantly its sensitivity to wind speed at high winds above
80 15 m s^{-1} (Donnelly et al. 1999). Most SAR systems can now emit in one polarization (vertical
81 or horizontal) and receive in both polarizations (vertical and horizontal). The cross-polarized
82 signal is more sensitive to volume scattering by breaking waves than the co-polarized signal
83 (Zhang et al. 2017). Because the occurrence of breaking waves increases with wind speed,
84 the volume scattering, i.e., the cross-polarized signal, observed as a normalized radar cross
85 section (NRCS), also increases with wind speed (Phillips 1988; Hwang et al. 2010). Although
86 open questions remain regarding the relative importance of surface and volume scattering,
87 these two scatterings are the basic principles behind high wind speed estimates made by
88 SAR. SAR wind speeds are retrieved by using geophysical model functions (GMFs) that
89 relate the strength of the cross-polarization NRCS to 1-min sustained ocean winds observed
90 by the Stepped Frequency Microwave Radiometer (SFMR, Uhlhorn and Black 2003;
91 Uhlhorn et al. 2007). Mouche et al. (2019) and Combot et al. (2020) showed by using
92 independent observations that SAR wind speeds with a horizontal resolution of 3 km are in
93 good agreement with 1-min sustained ocean winds from the SFMR (Uhlhorn and Black
94 2003; Uhlhorn et al. 2007) with root mean squared error (RMSE) $< 5 \text{ m s}^{-1}$.

95 Radarsat-2 (RS2), Radarsat-C1/C2/C3 (Radarsat Constellation Mission, RCM), Sentinel-
96 1A (S1A), and Sentinel-1B (S1B) satellites equipped with C-band SARs with a wide swath
97 mode can observe TCs twice a day in a sun-synchronous sub-recurrent orbit with a local

98 time of ~06:00 on the descending node and ~18:00 on the ascending node (e.g., Radarsat-
99 2, European Space Agency, ESA, 2012). While these C-band SARs have the same
100 capabilities, the Sentinel (S1A and S1B), RS2, and RCM instruments are slightly different.
101 In addition, Isoguchi et al. (2021) are currently working to develop a new SAR wind product
102 that uses the Phased Array L-band Synthetic Aperture Radar-2 (PALSAR-2) aboard the
103 Advanced Land Observing Satellite-2 (ALOS-2), whose local sun time is ~12:00 on the
104 descending node and ~00:00 on the ascending node (Japan Aerospace Exploration Agency,
105 JAXA, 2024). Because SAR observations can have a 12-hourly frequency (or, 6-hourly in
106 the future if ALOS-2/PALSAR-2 joins the TC observation community), SAR wind speeds
107 have many potential uses and applications (e.g., Ricciardulli et al. 2023), including for
108 intensity estimation (Howell et al. 2022), wind radii monitoring (e.g., Center for Satellite
109 Applications and Research, 2024), and data assimilation by operational numerical model
110 systems for TC prediction (Ikuta and Shimada 2024). Even lower frequency observations
111 can be useful for constructing an ocean truth dataset for estimation of a TC wind field through
112 application of a statistical regression method to relate them to other data (e.g., Tsukada and
113 Horinouchi 2023; Avenas et al. 2023). To realize such goals in the future, comparisons
114 between conventional best track estimates and SAR wind speeds are necessary. Such
115 comparisons can lead to more effective use of SAR wind speeds and improvements to TC
116 intensity and wind radii estimates.

117 Combot et al. (2020) compared SAR Vmax with the Joint Typhoon Warning Center

118 (JTWC) and National Hurricane Center (NHC) best track estimates of 1-min maximum wind
119 (V_{max} , JTWC 2024; NHC 2024) and showed that, although SAR V_{max} is generally
120 consistent with best track 1-min V_{max} , the root mean squared difference (RMSD) is large in
121 areas where no SFMR observations are available (e.g., in the western North Pacific). It is
122 still unclear, however, how consistent 1-min SAR wind speeds are with best track 10-min
123 V_{max} (V_{max10}) values estimated by Japan Meteorological Agency (JMA, JMA 2024). JMA
124 estimates V_{max10} primarily based on the Dvorak technique and its own conversion table
125 (i.e., Koba table, Koba et al. 1991). Moreover, previous JMA V_{max10} had been derived
126 mainly from the central pressure using Takahashi's equation (Takahashi 1952) until the
127 1980s (Aizawa et al. 2024). The Koba table, used today, was created based on those JMA
128 V_{max10} values. Takahashi's equation was empirically made using maximum 20-min
129 average wind speeds observed in islands and coastal areas for TCs (Takahashi 1940).
130 Because of these historical reasons, it is well known that JMA V_{max10} values do not
131 correspond linearly with 1-min V_{max} values of JTWC by multiplying a factor of 0.88 or 0.93
132 (e.g., Mei and Xie 2016; Harper et al. 2010) as is done by several Regional Specialized
133 Meteorological Centres (RSMCs). Therefore, it is important to investigate how to convert 1-
134 min SAR V_{max} values into V_{max10} values that match the conventional JMA values.

135 The purpose of this study is to investigate the consistency and differences between SAR
136 wind speeds and conventional best track estimates for TCs in the western North Pacific,
137 where no operational TC reconnaissance flights in the inner core are conducted except near

138 Hong Kong (Hon and Chan 2022). Variables in the investigation include V_{max10} , the radius
139 of the 30-kt wind speed (R_{30}), and the radius of the 50-kt wind speed (R_{50}) from JMA best
140 track data. V_{max} and RMW from the JTWC best track data are also examined for
141 comparison. Through these examinations, we highlight the need to continue improving the
142 quality of SAR wind products and to comprehensively evaluate conventional estimation
143 techniques for future work. Section 2 describes the datasets used and the methodology in
144 this study. Section 3 presents the results of the examination. Section 4 discusses challenges
145 and potential uses and applications of SAR wind observations. Section 5 provides
146 conclusions of this study.

147

148 **2. Data and Methodology**

149 *2.1 Data used*

150 We used C-band SAR wind products from the CyclObs database (Vinour et al. 2023),
151 provided by an IFREMER (French Research Institute for Exploitation of the Sea) team, with
152 a horizontal resolution of 3 km. Table 1 provides basic information on C-band SAR
153 acquisition modes whose products were used in this study. Because 3-km SAR wind speeds
154 are in good agreement with 1-min sustained ocean winds from the SFMR (Mouche et al.
155 2019; Combot et al. 2020), the 3-km SAR wind speeds are considered to be equivalent to
156 the 1-min sustained wind speed (e.g., Ricciardulli et al. 2023). In addition, the effect of rain
157 attenuation on wind speed must be considered. In areas of strong rainfall, backscattered

158 radar power can be decreased, resulting in decreases in retrieved wind speeds (by 5–10 m
159 s^{-1} , Mouche et al. 2019). Also, it is known that C-band SAR suffers from the effect of
160 hydrometeors in the melting layer on wind speed (Mouche et al. 2019; Alpers et al. 2020),
161 which can lead to overestimated wind speeds primarily observed along the outer rainbands.
162 Furthermore, SAR wind speeds have an incidence-angle-dependent bias (e.g., Ikuta and
163 Shimada 2024). We examine this incidence-angle-dependent bias in section 3.

164 For this study, we collected 191 SAR wind observation files from 2012 to 2021 for TCs in
165 the western North Pacific. However, after exclusion of cases with large data gaps within 100
166 km of the TC center and on the right side of the storm track, and landfalling cases, V_{max}
167 could be computed for 117 cases (61%). Although this study relies on the results of Combot
168 et al. (2020), who confirmed a good agreement between SFMR and SAR wind speeds, it
169 should be noted that the maximum retrieved value of CyclObs SAR wind speeds is 80.0 m
170 s^{-1} . Given that SFMR observed a surface wind speed of more than 90 m s^{-1} during Hurricane
171 Patricia (2015) (Kimberlain et al. 2016) and JTWC V_{max} can reach 85 m s^{-1} (i.e. the highest
172 value in the Dvorak current intensity table), it is possible that the CyclObs SAR wind speeds
173 are underestimated in the case of such an extremely intense storm.

174 Because the obtained SAR wind speeds are swath data with a horizontal resolution of 3
175 km, they are transformed into polar coordinate data by using the center position obtained by
176 the method described in Section 2.2 and Cressman interpolation. The polar coordinates are
177 2 km in the radial direction and 0.7° in azimuth (i.e., 512 grid points in azimuth). Although

178 these resolutions are arbitrary, they are determined to properly obtain wind structure
179 parameters even with large TC sizes. Then, a simple quality control (QC) procedure is
180 performed, in which outliers exceeding three times the standard deviation of winds (i.e., 3-
181 sigma QC) at each radius in the polar coordinate system are removed. However, it is not
182 possible to remove all outliers using this method.

183 Other data used in this study include JMA and JTWC best track data, JMA Dvorak analysis
184 data, and sea-surface wind (ASWind) data (Nonaka et al. 2019) derived from infrared (10.4
185 μm) atmospheric motion vectors (AMVs, Shimoji 2017) at heights below 700 hPa from
186 Himawari-8 target observations (Bessho et al. 2016). The spatial resolution of ASWind data
187 is 10 km. ASWinds are calibrated against ASCAT winds by multiplying the low-level infrared
188 AMVs by a reduction factor (0.76). We use ASWind data that have passed a QC process
189 (Nonaka et al. 2019) from the start of Himawari 8 operations (July 2015) to 2021. The best
190 track estimates (V_{max} , R30, R50, center positions, and RMW) used are linearly interpolated
191 to the SAR observation time. Hereafter, the 6-hourly synoptic time closest to the SAR
192 observation time is set to $t = 0$ h.

193

194 *2.2 Center Finding Process*

195 Center finding of a TC is conducted by using an interpolated best track center as a first
196 guess position. In this finding process, the center is defined as the point where the azimuthal-
197 mean SAR wind speed is maximized, a similar definition to what TC observational studies

198 have done (e.g., Marks et al. 1992; Lee and Marks 2000; Rogers et al. 2013). Considering
 199 the effect of the environmental wind and the effect of a false SAR wind maximum seen near
 200 the center (Li et al. 2013), we do not regard the point with the minimum SAR wind in the eye
 201 region as the TC center. More specifically, the center finding process is shown in Fig. 1 and
 202 as follows:

203 Step 1. Determine center positions candidates by interpolating the SAR wind speeds
 204 (SAR_{wind}) to 40×40 grid points at 0.025° intervals within 0.5 degrees of the interpolated
 205 best track position, and keeping those points that are 45% of the maximum wind
 206 (SAR_{max}) observed within the 0.5 degree area.

207 Step 2. Using the center position candidates identified in step 1, refine those center
 208 candidates by calculating the ratio of maximum azimuthal-mean SAR wind speed (\overline{Vm})
 209 centered on the center candidate to the center's wind speed (SAR_{wind}) and excluding
 210 candidates with a ratio less than 1.5 or the 40th percentile of all ratios, whichever is
 211 more restrictive. In this study, \overline{Vm} is computed if wind data are available for more than
 212 half of the polar grids at a given radius.

213 Step 3. Using the remaining center position candidates, calculate the symmetry of SAR wind
 214 speed (v). Our definition of symmetry (γ) is,

$$215 \quad \gamma(r) \equiv \frac{\overline{v(r)^2}}{\overline{v(r)^2} + \int_0^{2\pi} v'(r, \lambda)^2 d\lambda / 2\pi}, \quad (1)$$

216 where r and λ are the radial and tangential directions, respectively; the overbar denotes
 217 the azimuthal-mean; and the prime denotes the deviation from the azimuthal-mean. The

218 symmetry is averaged within a radius of 100 km. Next, normalize $\overline{V_m}$ ($\overline{V_m'}$) obtained in
219 step 2 by the highest value among all $\overline{V_m}$ values. Then, calculate the possible center
220 index (PCI) of the symmetry multiplied by the $\overline{V_m'}$ for each center position candidate.
221 Find the center position that has the highest PCI value.

222 Step 4. Repeat steps 1–3 with position candidates by interpolating the SAR wind speeds to
223 40×40 grid points at 0.01° intervals within 0.05° of the center position found in the
224 previous step 3. Then, the position obtained in step 3 becomes the final center position.

225 Although a reasonable center position can be objectively determined by the above
226 procedure, it was not possible in two cases because of observational noise. Therefore, in
227 those two cases, the center point was determined subjectively.

228

229 **3. Results**

230 Here, we validate and compare SAR wind speeds. First, we briefly validate SAR with
231 ASWinds. Second, we compare the seven nearly coincident SAR wind speed estimate
232 cases to assess intra-SAR differences. Finally, we present the results of our comparison
233 between SAR wind speeds and best track estimates.

234 *3.1 Comparison with ASWinds*

235 ASWinds are used for the estimation of R30 by JMA (Nonaka et al. 2019). ASWinds
236 available within 1 km from SAR wind grid points and within 10 min of SAR observations are
237 compared with SAR wind speeds in a two-dimensional histogram (Fig. 2). SAR wind speeds

238 below 20 m s^{-1} are consistent with ASWinds with a standard deviation of less than 3 m s^{-1} .
239 However, SAR wind speeds greater than 20 m s^{-1} are much higher than ASWinds. This
240 result is not surprising because of three reasons: (1) ASCAT winds tend to have negative
241 biases caused by saturation at high wind speeds (e.g., Chou et al. 2013); (2) ASWinds are
242 calibrated against ASCAT winds; and (3) the spatial resolution of ASWinds (10 km) is lower
243 than that of SAR wind speeds (3 km). A more sophisticated technique to derive AMVs with
244 a finer spatial resolution in a TC environment, such as one developed by Horinouchi et al.
245 (2023), would be needed to partially resolve the negative bias issue.

246 *3.2 Intercomparison of SAR Wind Products*

247 Next, we intercompare SAR wind speeds observed nearly simultaneously (within 10 min)
248 by two C-band SARs (RS2 and S1A or S1B). There are seven match-ups that can be used
249 for this purpose. Here, two SAR wind speeds are compared between the closest swath grid
250 points. Figure 3 shows two-dimensional histograms of the match-ups. Overall, the mean
251 absolute difference (MAD) is less than 2.5 m s^{-1} , and at wind speeds below 20 m s^{-1} , the
252 wind samples are concentrated along the 1-to-1 line. For wind speeds greater than 20 m
253 s^{-1} , however, there are systematic differences between RS2 SAR wind speeds and Sentinel-
254 1 (S1) SAR wind speeds:
255 (1) When the incidence angle of RS2 SAR is in the 20° – 30° range and that of S1 SAR is in
256 the 40° – 50° range, RS2 SAR wind speeds tend to be higher than S1 SAR wind speeds
257 (Figs. 3c and g).

258 (2) When the incidence angle relationship is opposite to that in (1), RS2 SAR wind speeds
259 tend to be lower than S1 SAR wind speeds (Fig. 3d).

260 (3) When the incidence angles of RS2 and S1 are almost the same and in the mid-20°–30°
261 range, RS2 SAR wind speeds tend to be lower than S1 SAR wind speeds (Fig. 3a).

262 (4) When the incidence angle of RS2 SAR is in the 20°–30° range and that of S1 SAR is in
263 the 30°–40° range, RS2 SAR wind speeds tend to be lower than S1 SAR wind speeds,
264 except for wind speeds $> 60 \text{ m s}^{-1}$ (Fig. 3e).

265 (5) When the incidence angle of RS2 SAR is in the 30°–40° range and that of S1 SAR is in
266 the 40°–50° range, RS2 SAR wind speeds tend to be higher than S1 SAR wind speeds
267 (Figs. 3b and f).

268 In light of the angle-of-incidence-dependent bias, which will also be discussed in section
269 3.3, these results seem reasonable if we consider that (i) there is a positive bias in the 20°–
270 30° range and a negative bias in the 40°–50° range and (ii) the magnitude of the bias differs
271 between RS2 and S1. This certainly results from the accuracy of the GMFs with respect to
272 the incidence angle and the instrument, and the quality of the signal within the swath in the
273 range direction (incidence angle and elevation antenna gain pattern). To rectify this
274 shortcoming, revisiting the GMFs using a larger sample of SAR collocations with reference
275 wind measurements such as SFMR is certainly required. In addition, recent studies have
276 revealed opportunities for improving the calibration of the SAR signal (Schmidt et al. 2023)
277 and the noise correction (Korosov et al. 2022).

278 3.3 Maximum Wind

279 a. Relationship between Best Track Vmax and SAR Vmax

280 When comparing best track Vmax with SAR wind speeds, it should be noted that the best
281 track Vmax has a coarse time resolution (i.e., 6-hourly) and does not represent localized
282 wind speed maxima (e.g., Franklin 2013). In contrast, SAR wind speeds are instantaneous
283 and can reflect transient wind speed enhancements, but they also have outliers due to noise.
284 In this study, we define SAR maximum wind speed (SAR Vmax) as the 99th percentile of
285 SAR wind speeds at grid points within 200 km from the center in the polar coordinate system.
286 The 99th percentile is determined as in Combot et al. (2020), although the grid point range
287 is different. In a preliminary analysis, we found that outliers due to noise are almost always
288 located above the 99th percentile. Because transient wind speed maxima should not be
289 regarded as Vmax, the 99th percentile is a reasonable cutoff even if no outlier wind speeds
290 are included in an observation. In this study, SAR Vmax is regarded as valid if SAR wind
291 observations are available for more than half of the polar grids within 100 km from the center.
292 Eight cases, however, are excluded where SAR wind observations are missing at the RMW
293 on the right side of the storm track.

294 We first show how the difference between best track Vmax and SAR Vmax changes when
295 different thresholds are used (Table 2). SAR Vmax values from the 99th percentile or above
296 are much greater than JMA Vmax10 values. Because SAR wind speeds with a horizontal
297 resolution of 3 km are considered to be greater than 10-min sustained wind speeds (e.g.,

298 Ricciardulli et al. 2023), it is expected for SAR Vmax to have a positive bias relative to JMA
299 Vmax10. In contrast, SAR Vmax values from the 99th percentile or below are much smaller
300 than JTWC best track Vmax. The fact that the maximum available SAR wind speed is 80.0
301 m s^{-1} , whereas the maximum JTWC best track Vmax is 87 m s^{-1} (170 kt, 1 kt = 0.5144 m
302 s^{-1}), may affect the JTWC bias.

303 Figure 4 shows scatter plots of SAR Vmax versus JMA best track Vmax10 and versus
304 JTWC Vmax. JMA Vmax10 values are much smaller than SAR Vmax values (MAD = 7.4 m
305 s^{-1} , Table 2), especially in the case of strong TCs. Even if we convert SAR Vmax values into
306 Vmax10 values by a factor of 0.93, which is recommended by the World Meteorological
307 Organization (WMO, Harper et al. 2010), the converted Vmax10 values are much higher
308 than JMA Vmax10 values (MAD = 6.4 m s^{-1}). If we depict the conversion relationship
309 between Vmax10 and Vmax derived from the Dvorak conversion tables of Dvorak (1984)
310 for Vmax and Koba et al. (1991) for Vmax10, we find that the data points are concentrated
311 on the conversion line (Fig. 4a), in particular, in the case of strong TCs. Hence if we convert
312 SAR Vmax values to 10-min values (hereafter, SAR Vmax10) using this conversion
313 relationship, the differences between JMA Vmax10 and SAR Vmax10 (hereafter ΔVmax10 ,
314 $\Delta\text{Vmax10} \equiv \text{JMA Vmax10} - \text{SAR Vmax10}$) become small; the mean ΔVmax10 is 0.4 m s^{-1} ,
315 and its MAD is 5.5 m s^{-1} . For JTWC Vmax, there is a rough 1-to-1 relationship between
316 JTWC Vmax and SAR Vmax (Fig. 4b); the mean difference between JTWC Vmax and SAR
317 Vmax (hereafter ΔVmax , $\Delta\text{Vmax} \equiv \text{JTWC Vmax} - \text{SAR Vmax}$) is -3.4 m s^{-1} , and its MAD

318 is 7.9 m s^{-1} (Table 2). Considering the difference in the range of V_{max} values between
319 JBTC and JMA, the level of the MADs for JBTC and JMA can be interpreted as nearly
320 identical. JMA V_{max} varies from 35 to 125 kt, while JBTC V_{max} varies from 35 to 170 kt.
321 Thus, the MAD of JBTC V_{max} should be 1.5 times (i.e., $(170-35)/(125-35)$) larger than that
322 of JMA $V_{\text{max}10}$, which is almost the same as the actual 1.4 times (i.e., $7.9/5.5$).

323 Next, we further investigate the characteristics of $\Delta V_{\text{max}10}$ for JMA and ΔV_{max} for JBTC.
324 One possible cause of the variabilities of $\Delta V_{\text{max}10}$ and ΔV_{max} is a bias that is dependent
325 on SAR incidence angle, as described in section 3.2. A scatter plot of the incidence angle
326 at the TC center versus SAR V_{max} (Fig. 5a) suggests that SAR V_{max} values are dependent
327 on the incidence angle. Thus, it seems that SAR V_{max} derived from the current product is
328 not suitable for quantitative use without any correction. However, in the absence of any true
329 reference data (e.g., SFMR winds), it is not possible to estimate how much SAR V_{max} is
330 biased relative to a given incidence angle. Figures 5b and 5c show scatter plots of the
331 incidence angle versus $\Delta V_{\text{max}10}$ for JMA and versus ΔV_{max} for JBTC, respectively. If we
332 assume that the incidence-angle-dependent bias of SAR $V_{\text{max}10}$ and SAR V_{max} is
333 deduced from the deviation from the best track $V_{\text{max}10}$ and V_{max} , then SAR $V_{\text{max}10}$ and
334 SAR V_{max} values associated with low incidence angles may have a positive bias and SAR
335 $V_{\text{max}10}$ and SAR V_{max} values associated with high incidence angles may have a negative
336 bias. This deduction is consistent with the results of the intercomparison between SAR wind
337 products in section 3.2.

338 Using the relationships between $\Delta V_{\max 10}$ and ΔV_{\max} and the SAR incidence angles for
 339 RS2 and Sentinel-1A and -1B (S1) shown in Figs. 5b and 5c, we can tentatively correct SAR
 340 $V_{\max 10}$ and SAR V_{\max} , respectively, using the following linear relationships:

$$\Delta V_{\max 10}' = \begin{cases} \Delta V_{\max 10} - (0.55 \theta - 18.77), & \text{(for RS2),} \\ \Delta V_{\max 10} - (0.41 \theta - 14.92), & \text{(for S1),} \end{cases} \quad (2)$$

$$\Delta V_{\max}' = \begin{cases} \Delta V_{\max} - (0.76 \theta - 22.51), & \text{(for RS2),} \\ \Delta V_{\max} - (0.53 \theta - 16.92), & \text{(for S1),} \end{cases} \quad (3)$$

341 where the prime indicates the corrected value, and θ indicates the incidence angle. Although
 342 this bias correction method may be quantitatively rough, it eliminates the incidence-angle-
 343 dependent bias for the CyclObs SAR wind speeds. Figure 6a shows a scatter plot of SAR
 344 $V_{\max 10}'$ versus JMA $V_{\max 10}$. The correction makes $\Delta V_{\max 10}'$ small; the mean absolute
 345 $\Delta V_{\max 10}'$ is 4.8 m s^{-1} . As for JTWC, the mean absolute $\Delta V_{\max}'$ is 6.7 m s^{-1} (Fig. 6b). Note
 346 that some SAR $V_{\max 10}'$ and SAR V_{\max}' with relatively poor coverage of SAR wind
 347 observations at the RMW might be underestimated, although cases with large data gaps at
 348 the RMW on the right side of the storm track are excluded. Hereafter the corrected SAR
 349 observations (i.e., $\Delta V_{\max 10}'$, SAR $V_{\max 10}'$, $\Delta v_{\max}'$, and SAR V_{\max}') are used. For
 350 reference, we confirm that the conclusions of this study are not changed even if uncorrected
 351 data are used.

352 *b. Characteristics of $\Delta V_{\max 10}'$ and $\Delta V_{\max}'$*

353 Another possible cause of the variabilities of $\Delta V_{\max 10}'$ and $\Delta V_{\max}'$ is associated with
 354 best track V_{\max} . Figure 7 shows that $\Delta V_{\max 10}'$ and $\Delta V_{\max}'$ are correlated with best track
 355 $V_{\max 10}$ and V_{\max} , respectively, at the time of the SAR observations ($t = 0 \text{ h}$); best track

356 Vmax values of weak TCs tend to be lower than SAR Vmax and those of intense TCs tend
357 to be higher than SAR Vmax. Table 3 shows that their correlation coefficients (r) are 0.77
358 for JMA and 0.73 for JTWC, respectively. It is unclear, however, whether these correlations
359 are due to a bias of SAR Vmax or to a bias of best track Vmax.

360 Table 3 also shows correlations between $\Delta V_{\max 10'}$ and $V_{\max 10}$ changes for JMA and
361 between $\Delta V_{\max}'$ and V_{\max} changes for JTWC. Although it is natural for weak TCs to
362 intensify and for intense TCs to weaken, it is interesting that there is a clear relationship
363 within the range from 30 to 50 m s⁻¹ for JMA (Fig. 7a); weakening (i.e., negative V_{\max}
364 changes) and steady-state (i.e., no V_{\max} change) TCs tend to have a positive $\Delta V_{\max 10'}$
365 and intensifying (i.e., positive V_{\max} changes) TCs tend to have a negative $\Delta V_{\max 10}'$. Note
366 that weakening TCs with a negative $\Delta V_{\max 10}'$ include TCs landfalling within 24 h after the
367 SAR observations (Fig. 7a). Also, among the eight TCs that experienced extratropical
368 transition (ET) within 24 h after the SAR observations, seven were weakening TCs with
369 negative $\Delta V_{\max 10}'$ (not included in Fig. 7a because of the lack of best track $V_{\max 10}$
370 estimates since ET). For JTWC, a similar correlation is seen but it is weaker than that of
371 JMA (Fig. 7b and Table 3). The stronger correlation of JMA $V_{\max 10}$ changes with $\Delta V_{\max 10}'$
372 may suggest that JMA best track $V_{\max 10}$ has a time lag relative to SAR $V_{\max 10}$.

373 Best track V_{\max} values are primarily estimated by the Dvorak technique (Dvorak 1984)
374 with some modifications based on all available observations, including conventional satellite-
375 derived winds such as ASCAT and winds observed on islands. Knaff et al. (2010) evaluated

376 Dvorak intensity estimates with reference to aircraft observation-based best track Vmax
377 values and found that systematic biases in Dvorak intensity were a function of best track
378 Vmax, best track Vmax change trend, translation speed, latitude, and TC size. We also find
379 that R30 is correlated ($r = 0.37$) with $\Delta V_{\max 10'}$, likely because R30 is correlated with
380 intensity ($r = 0.37$). There is no correlation of $\Delta V_{\max 10'}$ with translation speed or latitude (r
381 $= -0.09, -0.01$, respectively). The characteristics of JMA Vmax10 are consistent with the
382 results of Knaff et al. (2010), except those for latitude and translation speed, and thus may
383 be attributed to the use of the Dvorak technique.

384 *c. Characteristics of $\Delta V_{\max 10'}$ stratified by intensity changes*

385 Here, we further examine characteristics of $\Delta V_{\max 10'}$ stratified by intensifying, steady-
386 state, weakening, and extratropical transitioning TCs in relation to the Dvorak analysis.
387 Intensifying, steady-state, and weakening cases are defined as cases with a positive JMA
388 Vmax10 change, no 24-h JMA Vmax10 change, and a negative JMA Vmax10 change from
389 $t = 0$ h to $t = 24$ h, respectively. Among 117 cases in the SAR dataset used, there are 102
390 cases with a Vmax10 change from $t = 0$ h to $t = 24$ h in JMA best track data; 34 intensifying
391 cases, 21 steady-state cases, and 47 weakening cases (Fig. 7a). The remaining 15 cases
392 include eight ET cases, five TCs that became a tropical depression or dissipated, and two
393 storms whose Vmax10 values are undefined at $t = 0$ h due to the lack of best track Vmax10
394 estimates near the time of ET.

395 Among 34 intensifying cases, more than half (21 cases, 62%) have negative $\Delta V_{\max 10'}$

396 less than -1 m s^{-1} , whereas only 15% (five cases) have positive $\Delta V_{\text{max}10'}$ more than 1 m
397 s^{-1} (Fig. 7a). The mean SAR $V_{\text{max}10'}$ of these 21 cases is 35.2 m s^{-1} , whereas the mean
398 JMA $V_{\text{max}10}$ is 29.3 m s^{-1} . Of the 21 cases, more than half (57%) have small RMWs of less
399 than 25 km , and 76% have RMWs less than the overall mean RMW of 41.4 km (Fig. 8a).
400 Also, the vast majority (90%) are TCs before reaching the Dvorak eye pattern, such as
401 organized cumulonimbus (Cb) clusters, central dense overcast (CDO), or a curved band-
402 type pattern (Fig. 8b). Velden et al. (2006) pointed out that Dvorak intensities of TCs with
403 such cloud patterns tend to be underestimates, and Knaff et al. (2010) found that rapidly
404 intensifying TCs tend to be underestimated by Dvorak analysis. Although SAR $V_{\text{max}10'}$ may
405 still show a bias, the result here is consistent with those of previous studies. Figure 8c shows
406 Typhoon Jongdari (2018) as an example. SAR $V_{\text{max}10'}$, though it has quantitative
407 uncertainty, is much greater than JMA $V_{\text{max}10}$ during the intensification stage of Jongdari
408 (2018). Also, Jongdari (2018) was characterized by a small RMW ($14\text{--}18 \text{ km}$) and a compact
409 structure (Figs. 8d and 8e).

410 Among 21 steady-state cases, 15 (71%) cases have JMA $V_{\text{max}10}$ greater than SAR
411 $V_{\text{max}10'}$ (Fig. 7a). Of these 15 cases, 14 have cloud patterns associated with TC eyes and
412 $V_{\text{max}10}$ values above 40 m s^{-1} in all 15 cases (not shown). In short, mature TCs tend to
413 exhibit these features.

414 Among 47 weakening cases, the majority (33 cases, 70%) have positive $\Delta V_{\text{max}10'}$
415 greater than 1 m s^{-1} , whereas only 21% (10 cases) have negative $\Delta V_{\text{max}10'}$ less than -1

416 m s^{-1} (Fig. 7a). The mean SAR $V_{\text{max}10'}$ of these 33 cases is 40.7 m s^{-1} , whereas the mean
417 JMA $V_{\text{max}10}$ is 46.2 m s^{-1} . Most of the 33 cases are TCs during and just after the mature
418 stage, and 67% of the 33 cases are associated with a TC eye (not shown). Typhoon Halong
419 (2019) is a typical example of a weakening TC (Figs. 9a and 9b).

420 For the majority of weakening cases, the positive $\Delta V_{\text{max}10'}$ might be associated with the
421 Dvorak time lag rule, according to which the current intensity (CI number) remains higher
422 than that estimated from the cloud pattern (T number) during the weakening stage (Lushine
423 1977). In fact, 61% of the 33 cases with positive $\Delta V_{\text{max}10'}$ greater than 1 m s^{-1} have a CI
424 number higher than their T number (Fig. 9c); the mean CI number is 5.7, whereas the mean
425 T number is 5.3. According to the table provided by Koba et al. (1991), a difference in the
426 CI number of 0.5 is equivalent to $\sim 3.6 \text{ m s}^{-1}$. JMA has a 12-h time lag rule, following Lushine
427 (1977). However, it has been pointed out that the 12-h lag is too long (Brown and Franklin
428 2004). Knaff et al. (2010) mentioned the possibility that the final T-number constraints of the
429 Dvorak analysis give a positive intensity bias to weakening TCs. Although it is possible that
430 the positive $\Delta V_{\text{max}10'}$ is simply caused by a negative bias of SAR $V_{\text{max}10'}$ converted from
431 SAR V_{max} using Dvorak tables for high winds, the finding here is consistent with previous
432 studies.

433 Although the number of cases is small, all six TCs that completed extratropical transition
434 without having made landfall within 24 h after SAR observations have SAR $V_{\text{max}10}$ greater
435 than best track $V_{\text{max}10}$ (Fig. 10). This result suggests that the best track V_{max} values of

436 extratropical transitioning TCs may be underestimated. This underestimation may be
437 because the Dvorak technique does not capture V_{max} at the time of the extratropical
438 transition. We also examine the relationship between the direction of vertical wind shear, the
439 direction of translation, and the position of SAR V_{max} (Fig. 11) for these six TCs. Generally,
440 the TC wind maximum is located on the front right side with respect to the translation
441 direction (Shapiro 1983; Kepert and Wang 2001). However, some extratropical transitioning
442 TCs are characterized by a wind speed maximum located on the left side with respect to the
443 translation direction (Figures 11c, e, and f), which is also the left side with respect to the
444 vertical shear direction. This feature is consistent with the findings of Ueno and Kunii (2009),
445 who showed that some TCs have a wind maximum on the left with respect to the TC
446 translation direction only when the vertical shear direction is close to the TC translation
447 direction. Furthermore, the extratropical transitioning TCs tend to have a wavenumber-2
448 asymmetric wind structure (Figs. 11d–f). The wavenumber-2 wind structure is one of the
449 typical wind distribution patterns of TCs that make landfall on the main islands of Japan
450 (Fujibe and Kitabatake 2007; Kitabatake and Fujibe 2009; Loridan et al. 2014).

451

452 *3.4 Wind radii*

453 *a. RMW*

454 We compare the RMWs between SAR and JTWC. Note that RMW values in the JTWC
455 best track are not reanalyzed following the season (JTWC 2024) and are the consequence

456 of the need to provide an RMW for TC vitals and input to numerical weather prediction. In
457 this study, the RMW is defined as the radius of maximum azimuthal-mean SAR wind, the
458 same as Tsukada and Horinouchi (2023), in consideration of the incidence-angle-dependent
459 bias and the rain attenuation bias in SAR wind speeds. This definition is slightly different
460 from that of JTWC, according to which the RMW is the radius of local V_{max} . For intense
461 TCs, however, the difference between these two definitions is not expected to result in
462 significantly different RMWs because both RMWs should be located near the eyewall. Figure
463 12a shows that the MAD between JTWC and SAR RMWs is 22.1 km with a correlation
464 coefficient of 0.34. This result is consistent with Fig. 12b of Combot et al. (2020).

465 Figures 12b and 12c show scatter plots of SAR RMWs versus JTWC V_{max} values and
466 versus SAR V_{max} values, respectively, and the frequency distribution of JTWC RMWs
467 versus JTWC V_{max} during the period of 2011–2021. Most of the observed cases in Fig. 12b
468 are concentrated on the frequency distribution of JTWC best track estimates. However, two
469 low frequency areas of the best track estimates have some observed cases. One is area I
470 defined as the area of V_{max} values with 15–30 $m s^{-1}$ and RMWs with 0–30 km. Area I has
471 13 observed cases in Fig. 12b. These 13 cases are characterized by large differences in
472 V_{max} and RMWs between SAR observations and JTWC estimates. Twelve cases among
473 the 13 cases have SAR V_{max} much greater than JTWC V_{max} ; the mean SAR V_{max} of the
474 12 cases is 33.1 $m s^{-1}$, whereas the mean JTWC V_{max} is 23.5 $m s^{-1}$. As a result, there are
475 only five cases in area I of Fig. 12c. All SAR RMWs of the 13 cases are much smaller than

476 JTWC RMWs (Fig. 12a).

477 Another is area II defined as the area of V_{\max} values with $30\text{--}60\text{ m s}^{-1}$ and RMWs with
478 $60\text{--}140\text{ km}$. Area II has 20 observed cases in Fig. 12c. Of these 20 cases, 95% have SAR
479 RMWs much larger than JTWC RMWs (Fig. 12a), and 85% are weakening TCs or TCs just
480 after eyewall replacement cycles (not shown). JTWC best track estimates, however, rarely
481 contain such large RMW cases with V_{\max} values with $30\text{--}60\text{ m s}^{-1}$. With the accumulation
482 of SAR wind observations, the climatological relationship between V_{\max} and RMWs in the
483 western North Pacific found in JTWC best track estimates may be completely updated in the
484 future.

485 *b. R30 and R50*

486 The swath range of SAR observations does not cover the entire R30 and R50 regions.
487 Therefore, in this study, to investigate the consistency between JMA best track R30 and
488 R50, and the SAR wind distribution, SAR wind speeds on the R30 and R50 circles are
489 divided into wind speed bins. In JMA, R30 is defined as the radius within which a 10-min
490 sustained wind speed greater than 30 kt ($\sim 15\text{ m s}^{-1}$) exists or “potentially” exists, and R50
491 is defined similarly. Thus, SAR wind speeds on the R30 and R50 circles are expected to be
492 lower than 15 and 25 m s^{-1} , respectively. Here, we use temporally interpolated R30 and R50
493 values. Also, we use SAR wind speeds transformed onto the polar coordinates so that the
494 number of wind samples at each radius is the same regardless of the TC size.

495 Figure 13 shows SAR wind speeds observed on the R30 and R50 circles using 2.5 and

496 5.0 m s⁻¹ bins, respectively. The best track R30 is generally consistent with the SAR wind
497 speeds; winds on the R30 circle are mostly (88%) less than 15 m s⁻¹. The cause of the
498 underestimation in 12% of the R30 samples would include the bias in SAR wind speeds and
499 the effect of strong environmental wind speeds, such as monsoon flow, as well as the actual
500 underestimation of R30. In contrast, on the R50 circle, 28% of the samples have wind
501 speeds of 25 m s⁻¹ or higher. The best track R50 tends to be underestimated even if the
502 difference between 1-min and 10-min sustained wind speeds is considered. We suspect that
503 the underestimation of the best track R50 is caused by the use of ASWinds and
504 scatterometer winds that have a low bias for winds greater than 20 m s⁻¹ (e.g., Fig. 3).

505

506 **4. Discussion**

507 Although SAR Vmax is equivalent to 1-min wind speed, it is consistent with JMA Vmax10
508 if we convert SAR Vmax into SAR Vmax10 with two Dvorak conversion tables used at JMA
509 and JTWC. This finding is helpful to use the brand-new SAR wind observations in a way
510 consistent with conventional JMA Vmax10. We should, however, be aware that this
511 conversion method is a by-product for convenience when the Dvorak technique is the main
512 tool for estimating TC intensity. According to Harper et al. (2010), the wind speed conversion
513 factor from 1-min to 10-min values is recommended to be 0.93, which is independent of wind
514 speed. This factor is derived from the relationship between mean wind and a gust factor.
515 Therefore, the wind speed conversion should essentially be done that way. It is possible that

516 SAR-based wind observations can be a main source for estimating TC intensity in the future
517 instead of the Dvorak technique if the frequency of SAR observations greatly increases.
518 Then, a time may come when a decision has to be made as to whether SAR Vmax should
519 be converted into Vmax10 that is consistent with conventional JMA Vmax10 or whether SAR
520 Vmax should be converted into Vmax10 by a factor of 0.93.

521 The comparison between JMA Vmax10 and SAR Vmax10' in section 3.3 suggests that
522 weakening and steady-state TCs that have reached a certain level of intensity may tend to
523 be overestimated in the Dvorak analysis. Also, the negative correlation between $\Delta V_{\max 10}'$
524 and future intensity changes suggests that the JMA Vmax10 lags behind SAR Vmax10'
525 during the intensifying stage and at the start of weakening stage; that is, actual Vmax10 may
526 increase earlier and start to decrease earlier than JMA Vmax10. After more SAR wind
527 observations have been accumulated and the incidence-angle-dependent bias has been
528 improved, whether these issues really exist in the Dvorak analysis and JMA Vmax10 should
529 be comprehensively investigated.

530 Currently, it is not easy to estimate wind structure parameters such as the RMW and R50
531 in the western North Pacific, where aircraft observations are not available. With the advent
532 of SAR wind observations as a truth dataset, it will be operationally possible to estimate wind
533 structure parameters by a statistical approach using infrared satellite cloud patterns (e.g.,
534 Kossin et al. 2007; Knaff et al. 2015; Tsukada and Horinouchi 2023) or a set of easily
535 available parameters including an outer wind radius, the Coriolis parameter, and Vmax

536 (Chavas and Knaff 2022; Avenas et al. 2023). The development of such a method will help
537 to further improve the best track estimates. We will perform this work in the future.

538

539 **5. Conclusions**

540 This study compared SAR wind speeds provided by CyclObs with best track 10-min V_{max}
541 and wind radii provided by JMA to examine the consistency between brand-new high wind
542 products and conventional TC best track estimates. We also examined best track 1-min
543 V_{max} and RMWs provided by JTWC for comparison. The SAR-derived maximum wind
544 (SAR V_{max}) was defined as the 99th percentile value of SAR wind speeds at grids within
545 200 km from the TC center in order to exclude outliers and transient wind speed maxima.
546 Furthermore, SAR V_{max} , which is considered to be the 1-min sustained wind speed, was
547 converted into 10-min V_{max} (SAR V_{max10}) by using Dvorak conversion tables for JTWC's
548 1-min V_{max} and JMA's 10-min V_{max} . Because SAR V_{max} shows a bias that is dependent
549 on SAR incidence angle, in this study, we tentatively corrected SAR V_{max} (SAR V_{max}') and
550 SAR V_{max10} (SAR V_{max10}') using a first order corrective term. After the correction, we
551 found that SAR V_{max10}' is consistent with JMA V_{max10} ; the mean difference between them
552 ($\Delta V_{max10}'$) is -0.1 m s^{-1} , and the mean absolute difference is 4.8 m s^{-1} . The mean
553 difference between SAR V_{max} and JTWC V_{max} ($\Delta V_{max}'$) is -0.1 m s^{-1} , and the mean
554 absolute difference is 6.7 m s^{-1} . We also found that $\Delta V_{max10}'$ was a function of current
555 intensity and intensity changes up to 24 h to 36 h in the future. Cases with negative

556 $\Delta V_{\max 10'}$ mostly include intensifying TCs or extratropical transitioning TCs. Most of the
557 intensifying TCs are at the stage before the TC eye appears in infrared satellite imagery;
558 this result may be related to the well-known negative bias of the Dvorak analysis. Also, it
559 can be seen that it is not easy to estimate $V_{\max 10'}$ for extratropical transitioning TCs by
560 conventional methods. In contrast, cases with positive $\Delta V_{\max 10'}$ mostly include steady-
561 state or weakening TCs. One possible cause of the positive bias is the 12-h time lag rule of
562 Dvorak intensity for steady-state and weakening TCs, according to which the current
563 intensity remains higher than the intensity derived from cloud patterns. There are large
564 differences in the RMWs between SAR observations and JTWC estimates. Some of the
565 cases with large RMW differences are characterized by cases with SAR V_{\max} much greater
566 than JTWC V_{\max} , cases with intense, but weakening TCs, and cases just after eyewall
567 replacement cycles. These results reveal that JTWC's RMW estimates are largely a function
568 of intensity, that is a climatology, and are, at times, much different from the observed (also
569 see Combot et al. 2020, and Avenas et al. 2024). This is not surprising due to the need to
570 provide this information for the guidance suite, but users of the existing RMW should be
571 aware of this shortcoming in the records. The comparison between JMA's R30 and R50 and
572 SAR wind speeds showed that best track R30 is generally consistent with SAR wind speeds,
573 whereas best track R50 is underestimated relative to SAR wind speeds. This
574 underestimation may be because, for winds above 18 m s^{-1} , scatterometer (e.g., ASCAT)
575 winds and AMV-derived winds (ASWinds) used to estimate R50 have a negative bias.

576 The time has come when a thorough review and revisitation of conventional methods such
577 as the Dvorak technique are both necessary and possible with the emergence of the new
578 SAR observation instrument. SAR wind observations still have some limitations. The
579 derivation of a geophysical model function to relate the ocean surface wind speed to the
580 radar signal under extreme conditions, properly accounting for the incident angle effect, is
581 an ongoing area of research. Future work, however, will allow the comprehensive evaluation
582 of conventional methods through the accumulation of SAR wind observations for many TCs.
583 These efforts will also contribute to the verification and improvement of TC intensity
584 forecasts.

585

586 **Data Availability Statement**

587 SAR wind products are provided by the CyclObs website (<https://cyclobs.ifremer.fr>). The
588 best track data from JMA are available on their website ([https://www.jma.go.jp/jma/jma-
589 eng/jma-center/rsmc-hp-pub-eg/trackarchives.html](https://www.jma.go.jp/jma/jma-eng/jma-center/rsmc-hp-pub-eg/trackarchives.html)). The best track data from JTWC are
590 available on their website (<https://www.metoc.navy.mil/jtwc/jtwc.html?western-pacific>).
591 ASWinds are obtained from the Himawari JDDS website ([https://www.jma.go.jp/jma/jma-
592 eng/satellite/jdds.html](https://www.jma.go.jp/jma/jma-eng/satellite/jdds.html)) although only National Meteorological and Hydrological Services can
593 have access to the data. The JMA Dvorak analysis data are not publicly available due to
594 restrictions.

595

596

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603

604

Appendix

605

List of acronyms and some symbols used in this paper

ALOS-2	Advanced Land Observing Satellite-2
AMV	Atmospheric motion vector
ASCAT	Advanced Scatterometer
ASWind	Sea-surface wind data derived from infrared AMVs
Cb	Cumulonimbus
CDO	Central dense overcast
CI number	Current Intensity number
ΔV_{\max}	Difference between JTWC V_{\max} and SAR V_{\max}
$\Delta V_{\max}'$	Bias-corrected ΔV_{\max}
$\Delta V_{\max 10}$	Differences between JMA $V_{\max 10}$ and SAR $V_{\max 10}$
$\Delta V_{\max 10}'$	Bias-corrected $\Delta V_{\max 10}$
ESA	European Space Agency
ET	Extratropical transition
GMF	Geophysical model function
IFREMER	French Research Institute for Exploitation of the Sea
JAXA	Japan Aerospace Exploration Agency
JMA	Japan Meteorological Agency

JTWC	Joint Typhoon Warning Center
MAD	Mean absolute difference
NHC	National Hurricane Center
NRCS	Normalized radar cross section
PALSAR-2	Phased Array L-band Synthetic Aperture Radar-2
PCI	Possible center index
QC	Quality control
QI	Quality indicator
R30	Radius of 30-kt wind speed
R50	Radius of 50-kt wind speed
RCM	Radarsat Constellation Mission
RMSD	Root mean squared difference
RMSE	Root mean squared error
RMW	Radius of maximum wind
RS2	Radarsat-2
RSMC	Regional Specialized Meteorological Centre
S1	Sentinel-1
S1A	Sentinel-1A satellite
S1B	Sentinel-1B satellite
SAR	Synthetic aperture radar
SD	Standard deviation
SFMR	Stepped Frequency Microwave Radiometer
TC	Tropical cyclone
Vmax	Maximum wind speed
Vmax'	Bias-corrected Vmax
Vmax10	Maximum 10-min sustained wind speed
Vmax10'	Bias-corrected Vmax10
WMO	World Meteorological Organization

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798

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List of Figures

800

801 Fig. 1. Schematic flow chart of the center finding process. See text for more details.

802

803 Fig. 2. Two-dimensional histogram of SAR wind speeds (m s^{-1}) versus ASWinds (m s^{-1}).

804 The mean difference is 0.76, and the standard deviation (SD) is 2.66 m s^{-1} . Only high-

805 quality ASWinds with quality indicator (QI) (Holmlund 1998) values greater than 0.6 are

806 used here. N is the total number of collocations, and Mean (X-Y) is the mean difference

807 between SAR wind speeds and ASWinds.

808

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810 (R2) SAR wind speeds are plotted on the x-axis, and Sentinel-1A (S1A) or Sentinel-1B

811 (S1B) SAR wind speeds are plotted on the y-axis. MAD is the mean absolute difference,

812 and SD is the standard deviation. The mean incidence angles of RS2 SAR wind speeds

813 and Sentinel-1 SAR wind speeds, in parentheses, are computed from match-ups with

814 RS2 SAR wind speeds greater than 20.0 m s^{-1} . Observation times are indicated in the

815 axis labels.

816

817 Fig. 4. Scatter plots of SAR Vmax (m s^{-1}) versus best track Vmax (m s^{-1}) for (a) JMA and

818 (b) JTWC. The black line indicates the 1-1 line. In (a), the orange line indicates the

819 conversion line between the 10-min and 1-min sustained V_{max} values by a factor of
820 0.93 (Harper et al. 2010). Also, the red circles are intensity points that relate the 1-min
821 sustained V_{max} Dvorak (1984) to the 10-min sustained V_{max} (Koba et al. 1991) through
822 Dvorak conversion tables, and the red line connecting them is the conversion line
823 between the 10-min and 1-min sustained V_{max} values.

824

825 Fig. 5. Scatter plots of (a) incidence angle versus ($^{\circ}$) SAR V_{max} ($m s^{-1}$) from RS2 (blue)
826 and S1A and S1B (S1, orange), (b) incidence angle ($^{\circ}$) versus ΔV_{max10} (JMA best
827 track $V_{max10} - SAR V_{max10}$, $m s^{-1}$), and (c) incidence angle ($^{\circ}$) versus ΔV_{max} (JTWC
828 best track $V_{max} - SAR V_{max}$, $m s^{-1}$). In (b), the linear lines indicate the linear
829 relationships between ΔV_{max10} and the incidence angles for RS2 SAR (blue) and S1
830 SAR (orange) expressed by Eq. (2). In (c), the linear lines indicate the linear
831 relationships between ΔV_{max} and the incidence angles for RS2 SAR (blue) and S1 SAR
832 (orange) expressed by Eq. (3).

833

834 Fig. 6. Scatter plots of (a) SAR V_{max10}' ($m s^{-1}$) versus JMA best track V_{max10} ($m s^{-1}$)
835 and (b) SAR V_{max}' ($m s^{-1}$) versus JTWC best track V_{max} ($m s^{-1}$). The black line
836 indicates the 1-1 line. MAD is the mean absolute difference. Colors indicate the
837 coverage (%) of SAR wind observations at the RMW.

838

839 Fig. 7. Scatter plots of (a) JMA $V_{\max 10}$ (m s^{-1}) versus bias-corrected $\Delta V_{\max 10}'$ (JMA
840 best track $V_{\max 10} - \text{SAR } V_{\max 10}'$, m s^{-1}) and (b) JTWC V_{\max} (m s^{-1}) versus bias-
841 corrected $\Delta V_{\max}'$ (JTWC best track $V_{\max} - \text{SAR } V_{\max}'$, m s^{-1}). Colors in (a) indicate
842 JMA $V_{\max 10}$ changes in the next 24 h from $t = 0$ h (i.e., the 6-hourly synoptic time
843 closest to the SAR observation time). Colors in (b) indicate JTWC V_{\max} changes in the
844 next 36 h from the time of SAR observations. The “x” mark with a gray circle plotted in
845 some of the circles indicates TCs that made landfall within 24 h in (a) and 36 h in (b).
846 Note that TCs that disappeared or became extratropical cyclones before the end of the
847 period of intensity changes are not included in these plots because of the lack of best
848 track V_{\max} data.

849

850 Fig. 8. Frequency histograms of (a) SAR-derived RMWs (km) and (b) Dvorak cloud
851 patterns of the 20 intensifying TCs with negative $\Delta V_{\max 10}'$. (c) Temporal evolution of
852 the JMA best track $V_{\max 10}$ (black line with black circles at 6-h intervals, m s^{-1}) and the
853 SAR $V_{\max 10}'$ (red circles, m s^{-1}) for Typhoon Jongdari (2018). (d, e) SAR wind
854 distributions for Typhoon Jongdari (2018) at (d) 2055 UTC 24 July and (e) 2045 UTC 25
855 July. The black circle in (d) and the outer black circle in (e) is JMA R30; in (e), the inner
856 red circle is the JMA R50. Cb and CDO in (b) are cumulonimbus and central dense
857 overcast, respectively.

858

859 Fig. 9. (a) Temporal evolution of JMA best track $V_{\max 10}$ (black line with black circles at 6-
860 h intervals, m s^{-1}) and SAR $V_{\max 10'}$ (red circles, m s^{-1}) of Typhoon Halong (2019), and
861 (b) the SAR wind distribution for the typhoon at 1957 UTC 5 November. The inner red
862 circle indicates JMA R50, and the outer black circle indicates JMA R30. (c) Box-and-
863 whisker plots of the $\Delta V_{\max 10'}$ (m s^{-1}) samples of 53 weakening cases stratified by the
864 difference between CI number and T number (CI number minus T number) from the
865 Dvorak analysis. The “x” marks indicate the mean, and the numbers of positive and
866 negative $\Delta V_{\max 10'}$ samples are shown in parentheses from left to right.

867

868 Fig. 10. JMA best track $V_{\max 10}$ and SAR $V_{\max 10'}$ of extratropical transitioning TCs.

869

870 Fig. 11. SAR wind distributions of the six TCs that completed extratropical transition
871 without having made landfall within 24 h after SAR observations. The direction of vertical
872 wind shear (black arrows) and the translation direction (blue arrows) are shown together
873 with their magnitudes (black and blue numerals). The vertical wind shear is defined as
874 deep-layer (850–200 hPa) shear, which is the mean shear within 500 km from the TC
875 center using the Japanese 55-year Reanalysis (JRA-55) data (Kobayashi et al. 2015).
876 The translation speed is calculated by a 12-h centered difference in the JMA best-track
877 position. The red circle indicates JMA R50, and the black circle indicates JMA R30.

878

879 Fig. 12. (a) Scatter plot of JTWC RMWs (km) versus SAR RMWs (km) below 140 km. (b)
880 Scatter plot of JTWC Vmax (m s^{-1}) versus SAR RMWs (km) and two-dimensional
881 histogram of JTWC Vmax (m s^{-1}) versus JTWC RMWs (km) during the period of 2011–
882 2021 (colors). (c) As in (b), but for scatter plot of bias-corrected SAR Vmax (m s^{-1})
883 versus SAR RMWs (km). In (a), the black circles are the same as those in area I in (b),
884 and the red circles are the same as those in area II in (c). In (b, c), the black and red
885 boxes correspond to areas I and II, respectively.

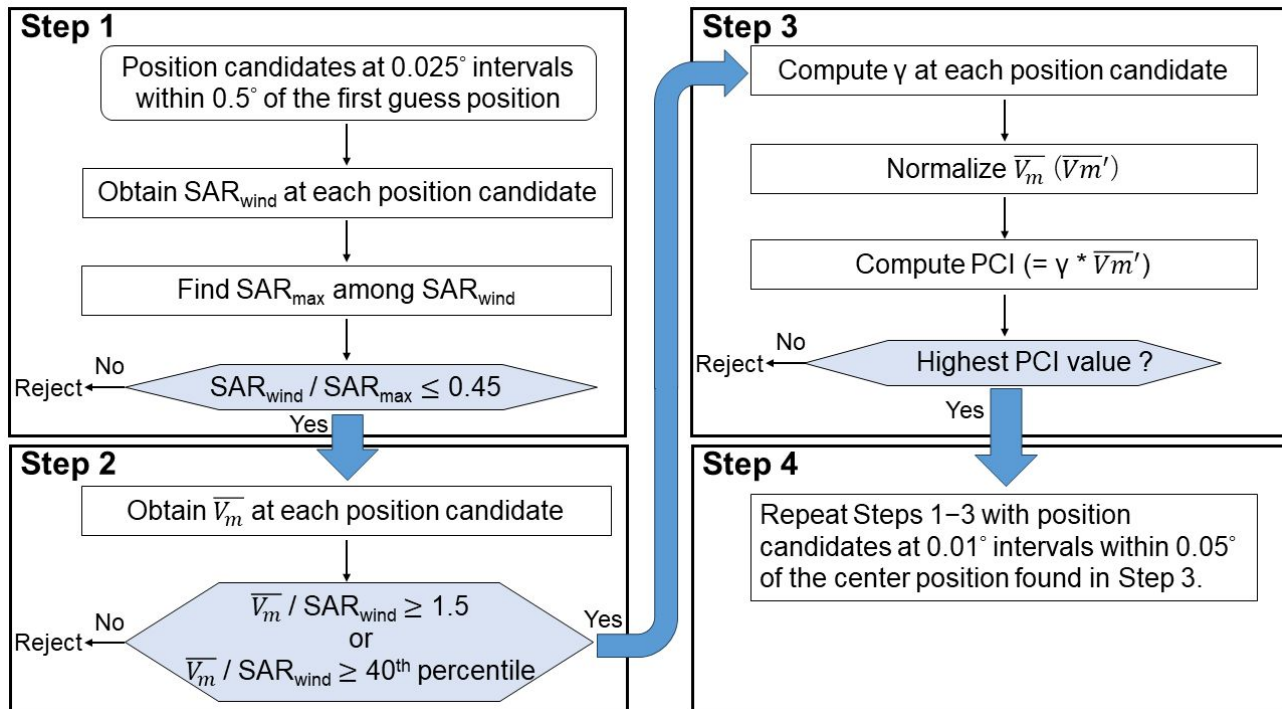
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887 Fig. 13. Frequency histograms of SAR wind speeds on the JMA (a) R30 and (b) R50

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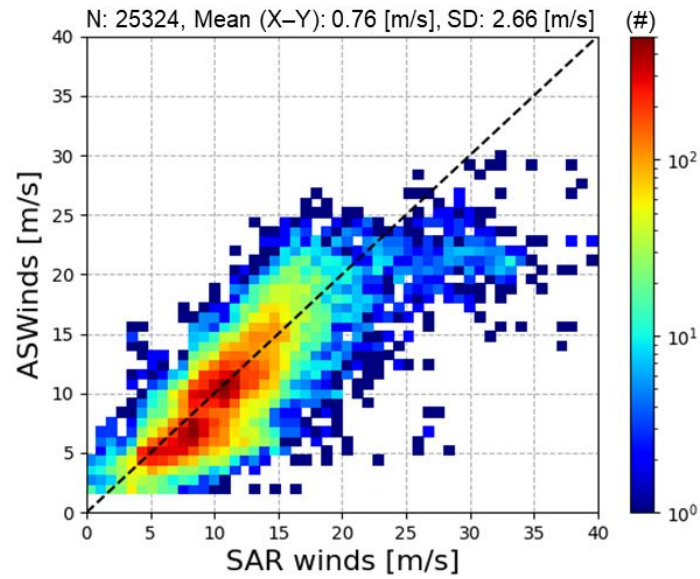
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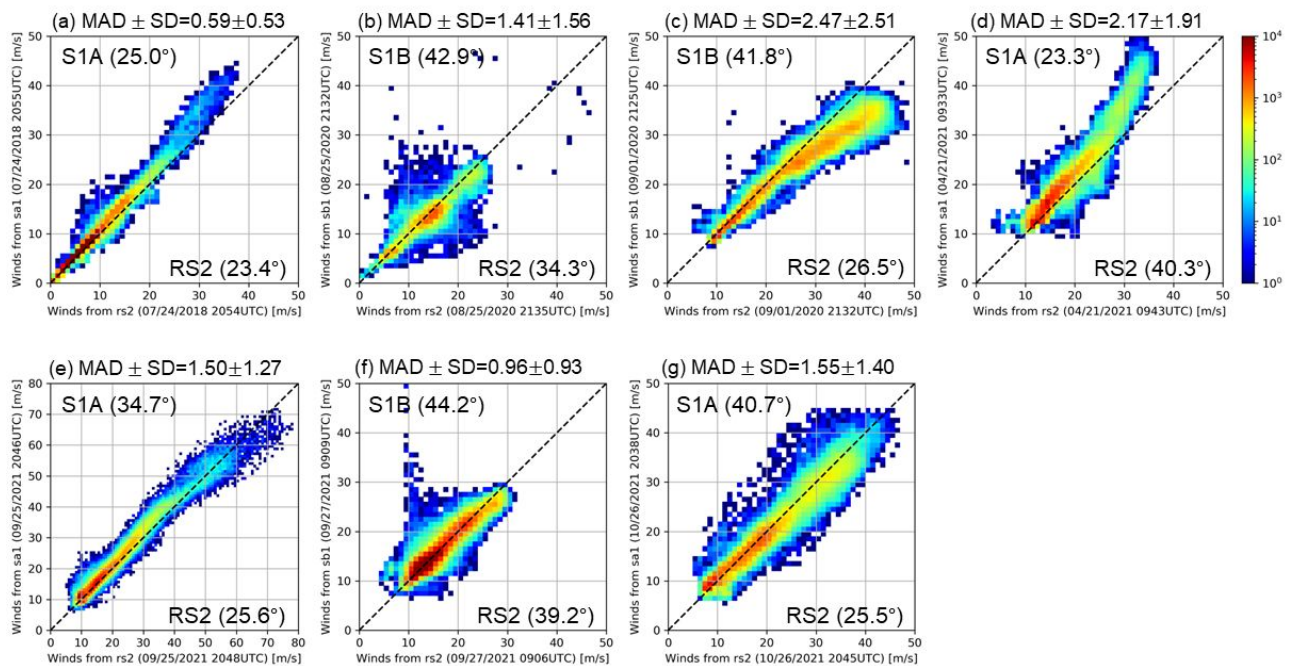


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899 ASWinds with quality indicator (QI) (Holmlund 1998) values greater than 0.6 are used
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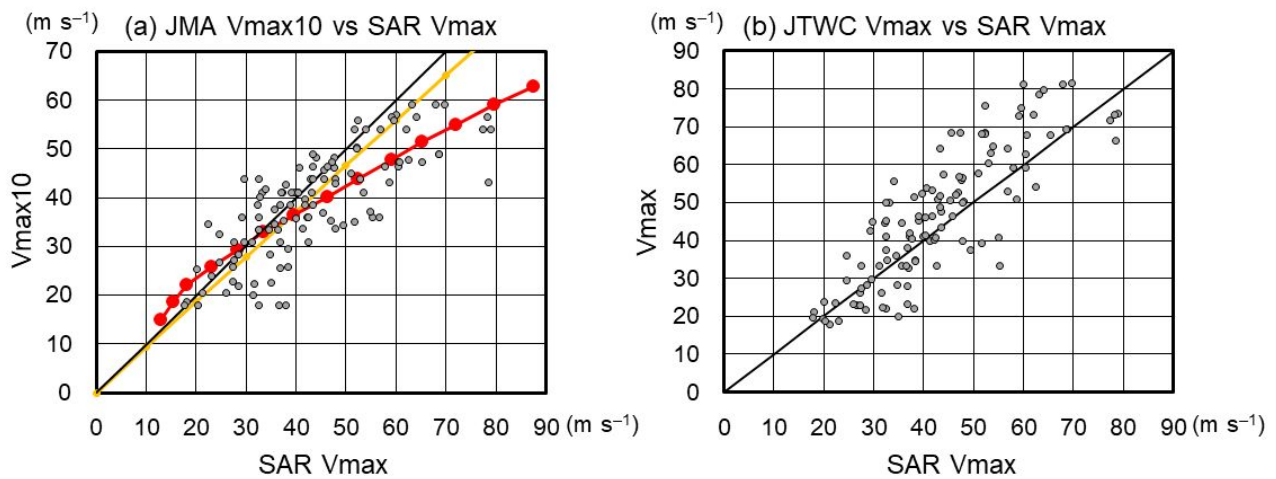
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 906 (RS2) SAR wind speeds are plotted on the x-axis, and Sentinel-1A (S1A) or Sentinel-1B
 907 (S1B) SAR wind speeds are plotted on the y-axis. MAD is the mean absolute difference,
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 909 and Sentinel-1 SAR wind speeds, in parentheses, are computed from match-ups with RS2
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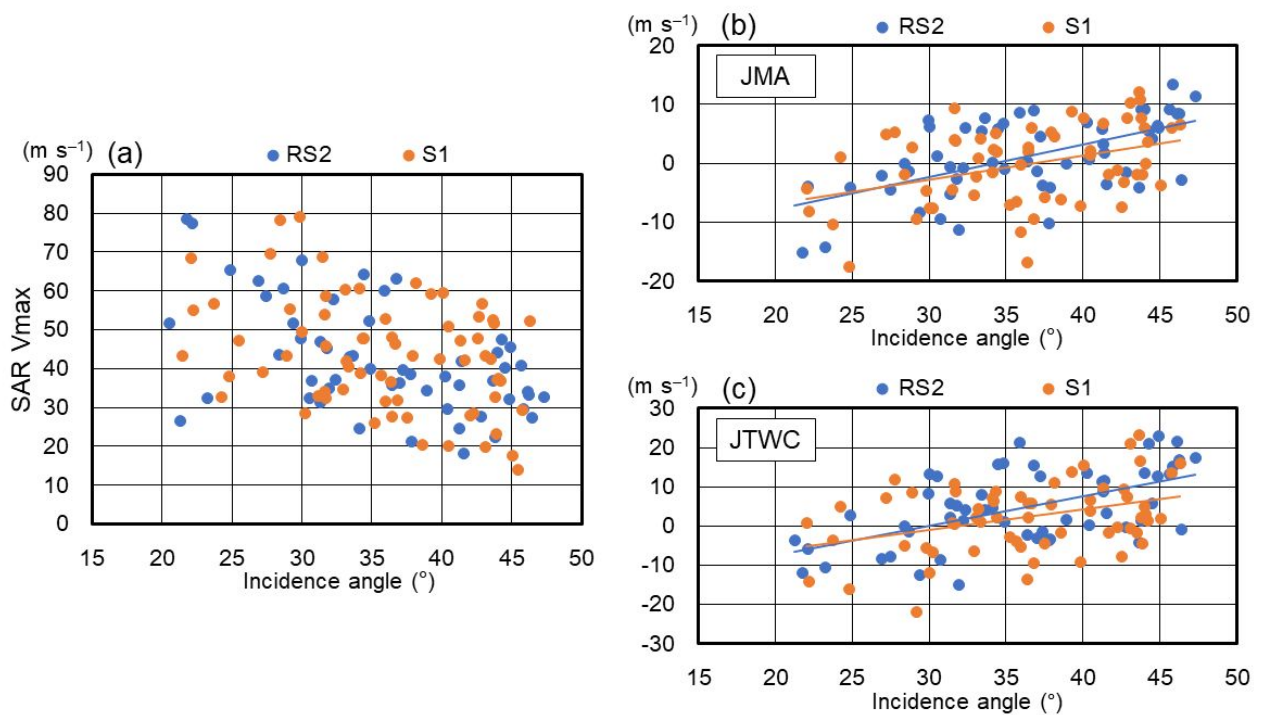
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 917 (b) JTWC. The black line indicates the 1-1 line. In (a), the orange line indicates the
 918 conversion line between the 10-min and 1-min sustained Vmax values by a factor of 0.93
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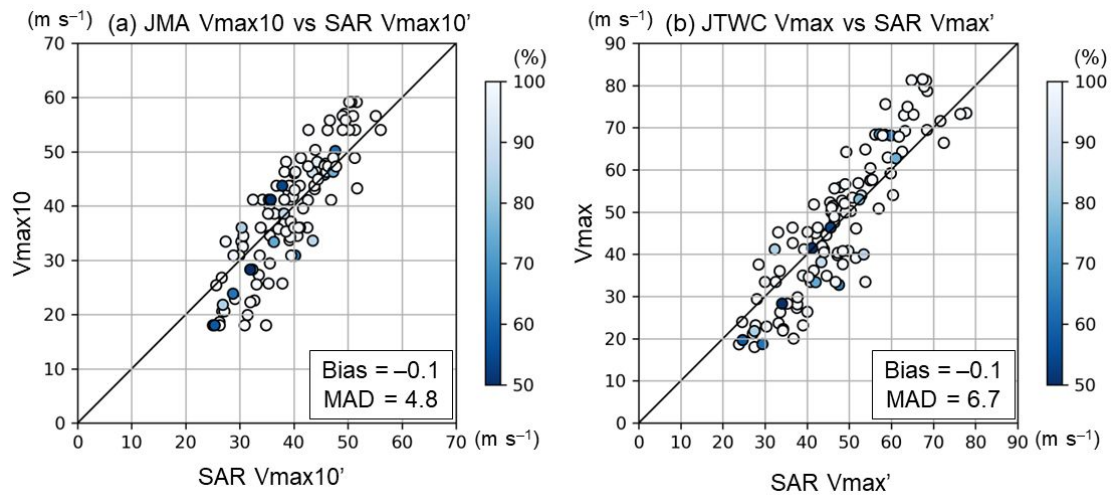
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 928 S1A and S1B (S1, orange), (b) incidence angle (°) versus $\Delta V_{\max 10}$ (JMA best track
 929 Vmax10 - SAR Vmax10, m s⁻¹), and (c) incidence angle (°) versus ΔV_{\max} (JTWC best
 930 track Vmax - SAR Vmax, m s⁻¹). In (b), the linear lines indicate the linear relationships
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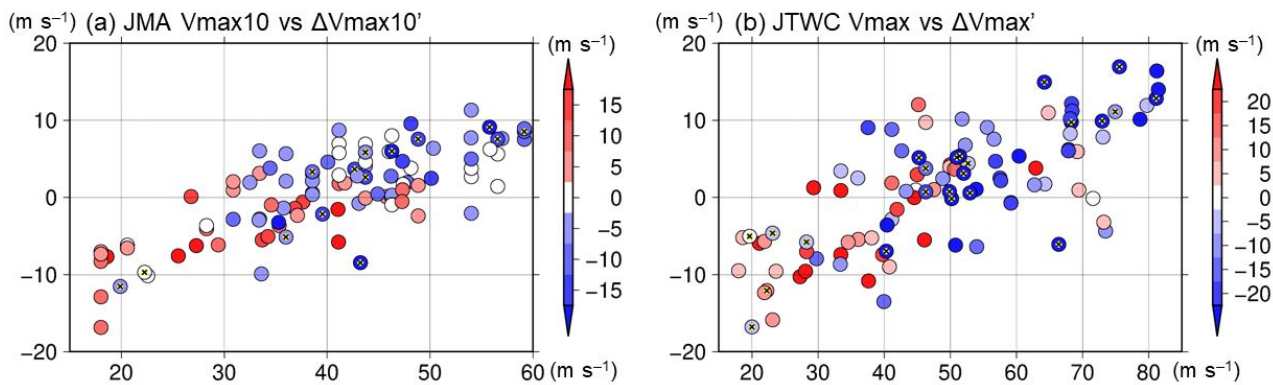


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 939 (b) SAR Vmax' (m s^{-1}) versus JTWC best track Vmax (m s^{-1}). The black line indicates the
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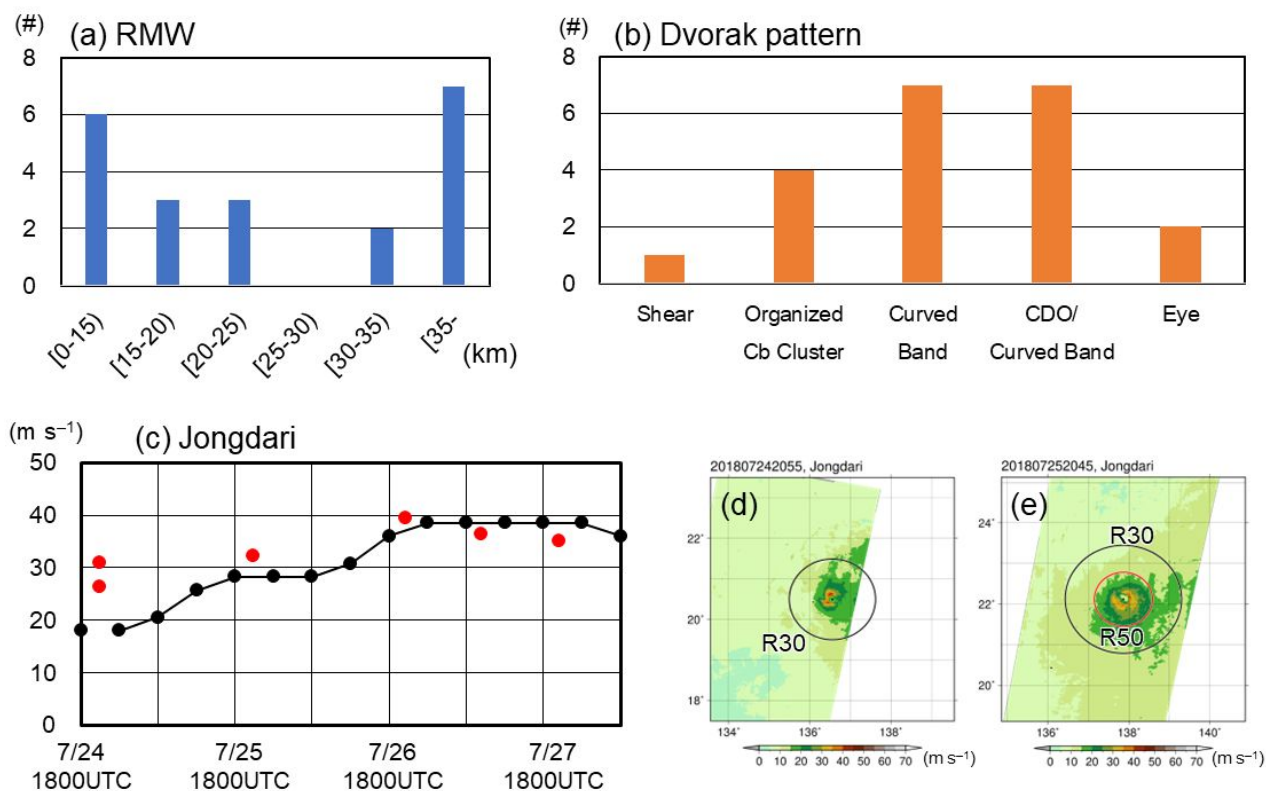
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 947 track $V_{\max 10} - \text{SAR } V_{\max 10}'$, m s^{-1}) and (b) JTWC V_{\max} (m s^{-1}) versus bias-corrected
 948 $\Delta V_{\max}'$ (JTWC best track $V_{\max} - \text{SAR } V_{\max}'$, m s^{-1}). Colors in (a) indicate JMA $V_{\max 10}$
 949 changes in the next 24 h from $t = 0$ h (i.e., the 6-hourly synoptic time closest to the SAR
 950 observation time). Colors in (b) indicate JTWC V_{\max} changes in the next 36 h from the
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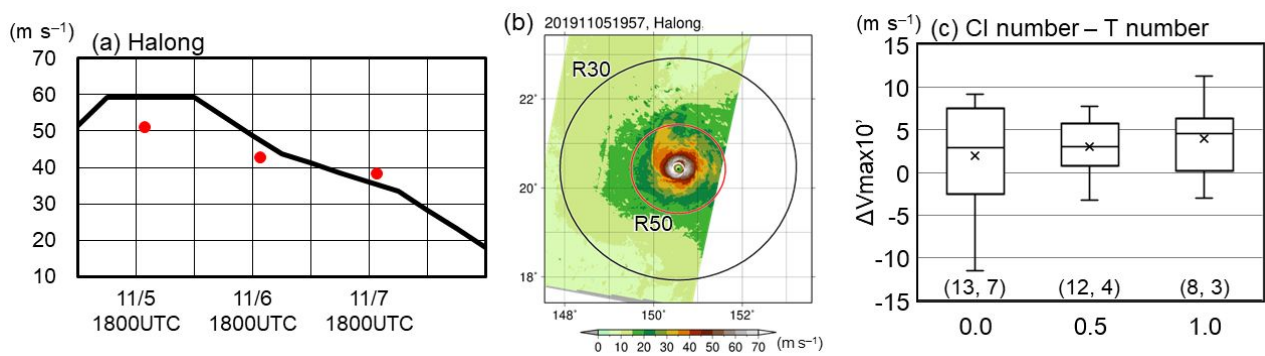
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958

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 960 of the 20 intensifying TCs with negative ΔV_{max10} . (c) Temporal evolution of the JMA
 961 best track V_{max10} (black line with black circles at 6-h intervals, $m s^{-1}$) and the SAR
 962 V_{max10} (red circles, $m s^{-1}$) for Typhoon Jongdari (2018). (d, e) SAR wind distributions
 963 for Typhoon Jongdari (2018) at (d) 2055 UTC 24 July and (e) 2045 UTC 25 July. The
 964 black circle in (d) and the outer black circle in (e) is JMA R30; in (e), the inner red circle is
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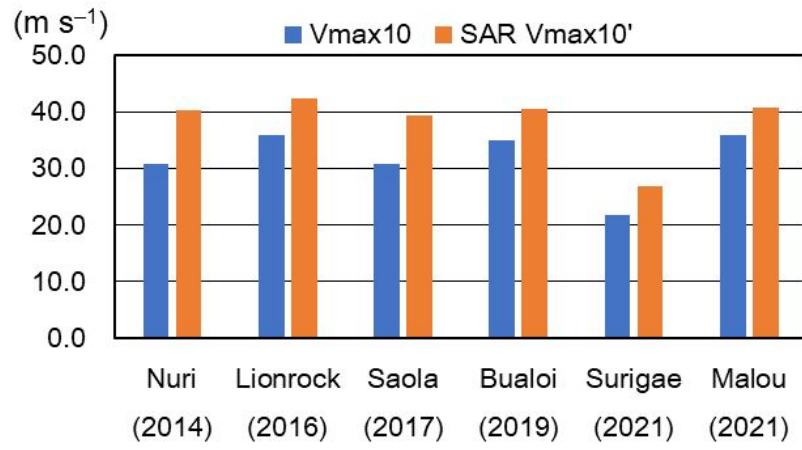


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971 Fig. 9. (a) Temporal evolution of JMA best track Vmax10 (black line with black circles at 6-
 972 h intervals, m s⁻¹) and SAR Vmax10' (red circles, m s⁻¹) of Typhoon Halong (2019), and
 973 (b) the SAR wind distribution for the typhoon at 1957 UTC 5 November. The inner red
 974 circle indicates JMA R50, and the outer black circle indicates JMA R30. (c) Box-and-
 975 whisker plots of the ΔVmax10' (m s⁻¹) samples of 53 weakening cases stratified by the
 976 difference between CI number and T number (CI number minus T number) from the
 977 Dvorak analysis. The "x" marks indicate the mean, and the numbers of positive and
 978 negative ΔVmax10' samples are shown in parentheses from left to right.

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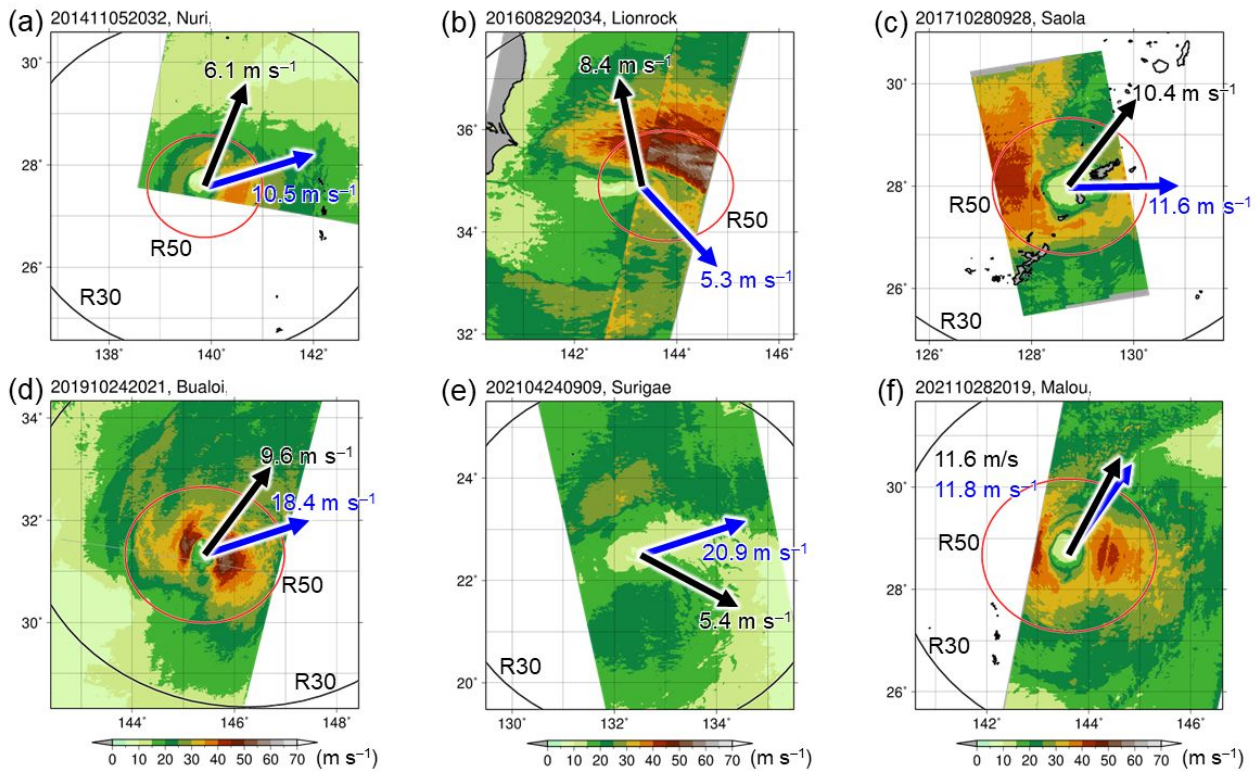


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982 Fig. 10. JMA best track Vmax10 and SAR Vmax10' of extratropical transitioning TCs.

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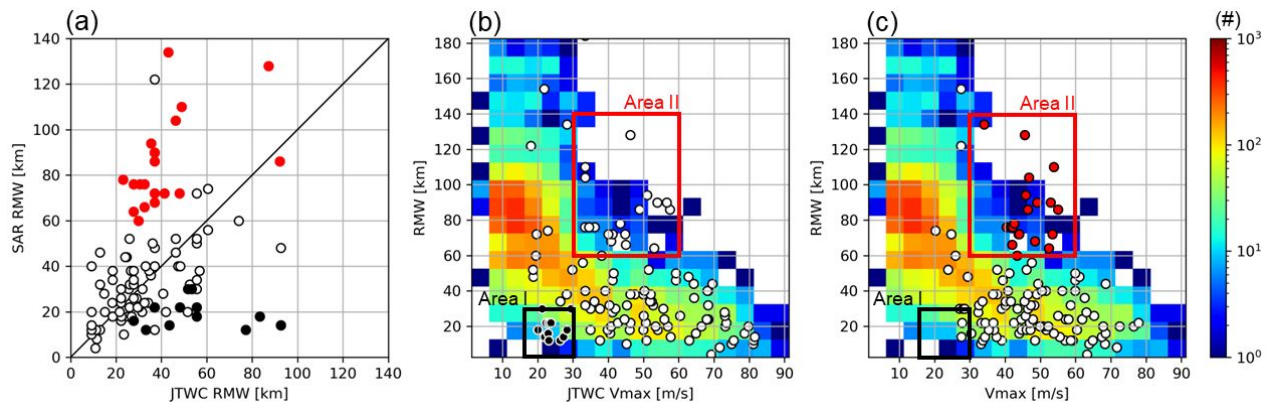


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985

986 Fig. 11. SAR wind distributions of the six TCs that completed extratropical transition without
 987 having made landfall within 24 h after SAR observations. The direction of vertical wind
 988 shear (black arrows) and the translation direction (blue arrows) are shown together with
 989 their magnitudes (black and blue numerals). The vertical wind shear is defined as deep-
 990 layer (850–200 hPa) shear, which is the mean shear within 500 km from the TC center
 991 using the Japanese 55-year Reanalysis (JRA-55) data (Kobayashi et al. 2015). The
 992 translation speed is calculated by a 12-h centered difference in the JMA best-track
 993 position. The red circle indicates JMA R50, and the black circle indicates JMA R30.

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997 Fig. 12. (a) Scatter plot of JTWC RMWs (km) versus SAR RMWs (km) below 140 km. (b)

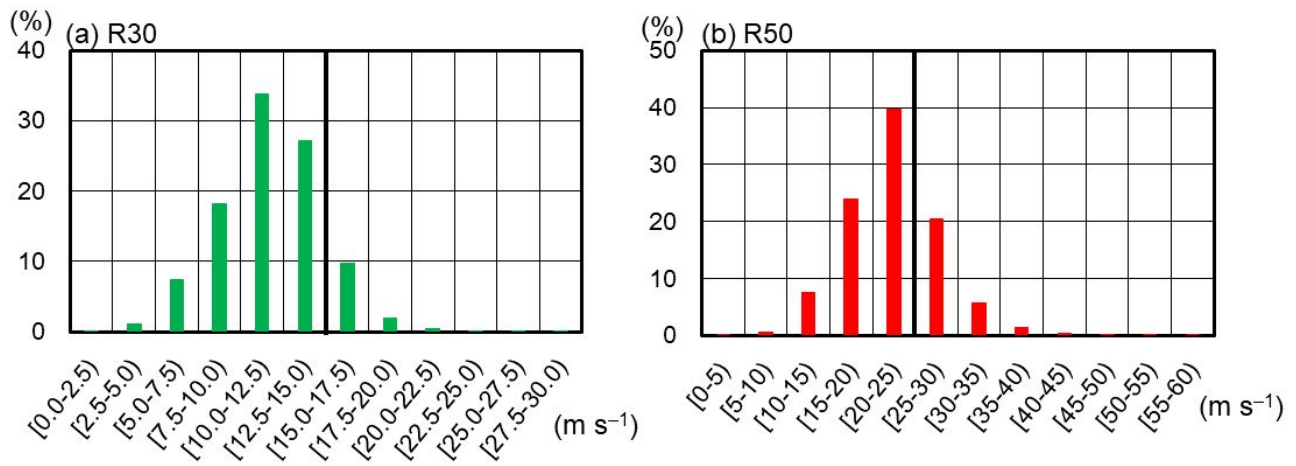
998 Scatter plot of JTWC Vmax (m s^{-1}) versus SAR RMWs (km) and two-dimensional999 histogram of JTWC Vmax (m s^{-1}) versus JTWC RMWs (km) during the period of 2011–1000 2021 (colors). (c) As in (b), but for scatter plot of bias-corrected SAR Vmax (m s^{-1}) versus

1001 SAR RMWs (km). In (a), the black circles are the same as those in area I in (b), and the

1002 red circles are the same as those in area II in (c). In (b, c), the black and red boxes

1003 correspond to areas I and II, respectively.

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1007 Fig. 13. Frequency histograms of SAR wind speeds on the JMA (a) R30 and (b) R50 circles.

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List of Tables

1010

1011 Table 1. Basic information on C-band SAR acquisition modes whose products were used

1012 in this study. The source of the information is mostly taken from Vinour et al. (2023).

1013 Azimuth is the along-track direction. Range is the cross-track direction.

1014

1015 Table 2. Biases and mean absolute differences (MADs) (m s^{-1}) between JMA best track

1016 $V_{\text{max}10}$ and SAR V_{max} and between JTWC best track V_{max} and SAR V_{max} (SAR –

1017 best track).

1018

1019 Table 3. Correlation coefficients (r) between $\Delta V_{\text{max}10}$ and best track $V_{\text{max}10}$ at $t = 0$ h

1020 and $V_{\text{max}10}$ changes for JMA and between ΔV_{max} and best track V_{max} at $t = 0$ h and

1021 V_{max} changes for JTWC. Because best track data do not include V_{max} after ET or

1022 dissipation, the number of cases is not necessarily 117.

1023

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1026 Table 1. Basic information on C-band SAR acquisition modes whose products were used
 1027 in this study. The source of the information is mostly taken from Vinour et al. (2023).
 1028 Azimuth is the along-track direction. Range is the cross-track direction.

Satellite	Acquisition mode	Swath	Incidence angle	Resolution (range × azimuth)
Radarsat-2 (RS2)	SCANSAR Wide imaging mode	450–500 km	~20°–49°	100 m × 100 m 20 m × 22 m
	Interferometric Wide swath mode	250 km	~31°–46°	(Level-1 Ground Range Detected High resolution)
Sentinel-1A (S1A) and Sentinel-1B (S1B) satellites				93 m × 87 m
	Extra Wide swath mode	400 km	~20°–47°	(Level-1 Ground Range Detected Medium resolution)

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1030

1031 Table 2. Biases and mean absolute differences (MADs) (m s^{-1}) between JMA best track
 1032 Vmax10 and SAR Vmax and between JTWC best track Vmax and SAR Vmax (SAR –
 1033 best track).

		Percentiles	95	98	99	99.5	100
Bias (m s^{-1})	JMA (2012-2021)		-0.1	3.2	4.9	6.2	10.8
	JTWC (2012-2021)		-8.4	-5.1	-3.4	-2.1	2.5
MAD (m s^{-1})	JMA (2012-2021)		6.0	6.7	7.4	8.2	12.0
	JTWC (2012-2021)		10.1	8.5	7.9	7.6	8.3

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1035

1037 Table 3. Correlation coefficients (r) between $\Delta V_{\max 10'}$ and best track $V_{\max 10}$ at $t = 0$ h
 1038 and $V_{\max 10}$ changes for JMA and between $\Delta V_{\max}'$ and best track V_{\max} at $t = 0$ h and
 1039 V_{\max} changes for JTWC. Because best track data do not include V_{\max} since ET or
 1040 dissipation, the number of cases is not necessarily 117.

Period	best track $V_{\max 10}$ or V_{\max} at $t = 0$ h	Best track $V_{\max 10}$ changes or V_{\max} changes							
		-	-6~ +6h	-12~ 0h	-12 ~ +12h	0~ +12h	0~ +18h	0~ +24h	0~ +30h
JMA $\Delta V_{\max 10}'$	0.77	-0.21	-0.10	-0.30	-0.38	-0.42	-0.48	-0.48	-0.46
# of JMA cases	115	112	112	106	109	107	102	96	90
JTWC $\Delta V_{\max}'$	0.73	-0.06	0.12	-0.07	-0.22	-0.25	-0.40	-0.45	-0.45
# of JTWC cases	117	117	117	113	113	112	106	103	99

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