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**Data quality control and calibration for mini-radiosonde  
system “Storm Tracker” in Taiwan**

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## Abstract

This study reports the correction methods of a newly introduced upper-air radiosonde instrument, “Storm Tracker” (ST), with more than one thousand co-launches of ST and Vaisala RS41-SGP (VS) data in field observations in the Taiwan area during 2016–2022. The co-launches provided more than a million comparable independent observations of wind, pressure, temperature, and humidity (PTU) data. Using the VS measurements as the reference, we use the statistical models, including the cumulative distribution function (CDF) matching method and generalized linear model (GLM), to correct the temperature and moisture fields of the ST sounding. Both approaches yield similar results. With a sounding-by-sounding comparison, the CDF-corrected ST soundings have a 1-K temperature and 7% relative humidity root mean square difference from the VS soundings. These error differences can be reduced to 0.66-K and 4.61% below the 700-hPa height. The GPS estimated a  $0.05 \text{ ms}^{-1}$  ST wind difference from the VS sounding. The biases of the corrected ST observations are slightly larger than the random errors, which were 0.24 K and 2.21% in the laboratory and 0.52 K and 2.23% in the field. The lower atmosphere in a region of complex terrain may have large wind, temperature, and moisture variations. With the relatively low cost, a high proportion of successful launches, and accuracy of wind, temperature, and moisture, ST can complement regular upper-air radiosonde observations for high-resolution observations in the lower troposphere. The high-

51 resolution lower troposphere observation is important for severe weather research in East  
52 Asia.

53 **Keywords** boundary layer; upper-air radiosonde observation; field campaign; data quality  
54 control and correction

## 56 1. Introduction

57 Upper-air radiosondes are one of the most important meteorological instruments  
58 for observing vertical profiles of atmospheric data at various altitudes. The measured  
59 pressure, temperature, and relative humidity (so-called “PTU”) data aids in weather  
60 forecasting, climate research, and the study of atmospheric dynamics. However, upper-air  
61 radiosondes are subject to certain biases due to instrument calibration, ascent rates, and  
62 environmental conditions. Collins (2001) distinguished the radiosonde observational errors  
63 into three types: random, rough, and systematic. According to Collins (2001), random error  
64 is caused by small-scale turbulence or unsystematic observational errors, and it is  
65 impossible to correct. The rough error can be introduced from observational protocol,  
66 computational error for data processing, or communication-related error. A properly defined  
67 operational procedure and automatic quality control process can minimize such errors. The  
68 third type of error, systematic error, is caused by insufficiencies in measurement devices or  
69 data processing procedures and persists in all observational data. This type of error can be  
70 detected and calibrated with statistical methods.

71 Nowadays, commercial radiosondes are often tested and corrected regarding  
72 these biases. However, they are typically characterized by their higher weight and cost,  
73 which limit the deployment of scientific field campaigns. The independently developed mini-  
74 radiosonde system – the “Storm Tracker (referred to as the ST, Figure 1b)” was created and  
75 first tested in 2016 (Hwang et al., 2020). The ST consists of a microcontroller

76 (ATMEGA328p), a GPS sensor (U-blox MAX7-Q), a pressure sensor (Bosch BMP280), a  
77 temperature–humidity sensor (TE-Connectivity HTU21D), and a transmitter (LoRa™). The  
78 sensors have an overall operation range from 1100 to 300 hPa in pressure and from -40°C  
79 to 85°C in temperature. The ST used a regular AAA battery for 2-4 hours of power; the total  
80 weight was 20g. More detailed hardware specifications can be found in Hwang et al., 2020.  
81 The design of ST aimed to leverage the low cost of sensors used in commercial electronics  
82 to enable high-frequency observations in the boundary layer. In addition, the receiver was  
83 designed to receive up to ten STs simultaneously. With such agility, using ST to gather  
84 supplemental data between regular sounding was ideal.

85         The ST was then put into intensive field observation operations for the first time  
86 during the Taipei Summer Storm Experiment (TASSE) in 2018–2020. The main goal of the  
87 field campaign is to investigate the thermal characteristics of the boundary layer in the Taipei  
88 Basin and local wind field variations to improve the forecasting ability of afternoon  
89 convection in the metropolitan area. Three advantages of using the ST for atmospheric field  
90 research were learned. First, the weight of ST with a battery is 20g, which helps to reduce  
91 the helium/hydrogen usage. Second, the commercial sensors, chips, and signal  
92 transmission components in the ST significantly reduce the cost and provide flexibility for  
93 multiple deployments and high spatial and temporal resolution observations. Lastly, the ST  
94 is easy to set up and can be quickly deployed or even mobile, which provides adaptability  
95 for different research needs and broadens the possibility for field campaign design.

96           The early work of ST by Hwang and colleagues (2020) showed an overall warm  
97   and dry bias in the troposphere compared to the VS, as shown in their figure 13. During  
98   TASSE, we discovered similar bias patterns and a typical example is shown in Figure 2.  
99   These common bias patterns motivated us to design a systematic approach to improve the  
100   data quality of ST. Our correction methods seek to align ST measurements as closely as  
101   possible with VS, enabling researchers to perform high-frequency, high-spatial resolution  
102   observations using ST with greater confidence and accuracy.

103

104           Many prior studies have recognized these biases and suggested that solar  
105   radiation can induce warm and dry bias for radiosonde measurements (Vömel et al. 2007).  
106   Similar daytime warm and dry biases have been reported in previous field experiments  
107   around the world that used relatively mature radiosonde systems (e.g., Wang et al., 2002;  
108   Ciesielski et al., 2009; Yu et al., 2015). Earlier studies indicated that radiosonde temperature  
109   biases are primarily contributed by radiative effects, with a minor proportion caused by the  
110   sensor response lag of the changing of temperatures as the radiosonde rises (e.g., McMillin  
111   et al., 1992; Sun et al., 2013).

112           The daytime temperature bias induced by solar heating was identified with various  
113   radiosonde systems (e.g., Luers, 1989, 1997; Luers and Eskridge, 1998; Sun et al., 2013;  
114   Lee et al., 2022; von Rohden et al., 2022). Their findings resulted in special surface coating  
115   over temperature sensors in most commercial radiosondes. Even though environmental

116 parameters can still affect the observed temperature, all factors influencing radiative or  
117 sensible heat flux around the sensor, such as the sensor surface temperature, solar angle,  
118 cloud fraction, and ventilation velocity, can cause the sensor temperature bias (e.g., McMillin  
119 et al., 1992; Luers and Eskridge, 1995; Mattioli et al., 2007; Lee et al., 2022). Luers and  
120 Eskridge (1998) evaluated the impact of the environmental parameters on the radiosonde  
121 in detail. Their results suggested that the temperature bias is most sensitive to solar angle,  
122 while the cloud cover has a slight impact. Also, the ventilation effect may cause bias when  
123 the sensor is in the balloon wake zone. To study the source of bias under a controlled  
124 environment, von Rohden et al. (2022) presented the Simulator for the Investigation of Solar  
125 Temperature Error of Radiosondes (SISTER), and Lee et al. (2022) proposed the Upper Air  
126 Simulator (UAS), which allows precise control over temperature, pressure, ventilation, and  
127 irradiation. Such advanced setups can help researchers measure the measurement errors  
128 more accurately and verify the cause of the errors.

129 In addition to temperature bias, the humidity bias has been discussed in many  
130 studies (e.g., Vömel et al., 2007; Yoneyama et al., 2008; Nuret et al., 2008; Kizu et al., 2018;  
131 Lee et al., 2022; Sommer et al., 2023). Vömel et al. (2007) found that the solar-heating-  
132 induced dry bias increased with altitude in the troposphere, which means the humidity bias  
133 also depended on the temperature. This resulted in the relative humidity (RH) measured in  
134 the low-temperature environment being less accurate (Miloshevich et al., 2001). Miloshevich  
135 et al. (2004) also pointed out that the response delay in humidity sensors could cause



136 measurement errors at low temperatures. The influence of these biases could be huge. For  
137 example, in the Tropical Ocean Global Atmosphere Coupled Ocean-Atmosphere Response  
138 Experiment (TOGA COARE, 1992-1993), scientists have reported the observational error  
139 induced an unrealistically dry boundary layer and caused an underestimate of convective  
140 available potential energy (CAPE) (Miller et al., 1999; Lucas and Zipser, 2000). Although the  
141 primary observation targets of ST are the lower troposphere environmental conditions, we  
142 still noticed significant warm and dry deviations in the near-surface boundary layer in TASSE  
143 (Figure 2).

144 Many studies have attempted to remedy the systematic error in radiosonde data  
145 with statistical methods. Lesht and Richardson (2002) mentioned that Vaisala accounts for  
146 the sensitivity of the RH sensor to temperature by using a high-order polynomial function  
147 with empirical coefficients. Yoneyama et al. (2008) applied a polynomial fitting function of  
148 pressure for the relative difference of RH and used the solar zenith angle as a factor for bias  
149 corrections. Other studies leveraged the thermodynamic equation and provided the  
150 temperature correction table with empirical correction factors (Wang et al., 2013; Dzambo  
151 et al., 2016).

152 In past field campaigns, scientists have also developed statistical models of  
153 humidity correction based on probability matching. For example, Ciesielski et al. (2009) used  
154 the cumulative distribution function (CDF) matching method to correct the humidity bias for  
155 nearby soundings. The advantage of the CDF-based calibration method is that the

156 calibration procedure is fast and straightforward. Building the correction table requires  
157 sufficient data to represent the statistical characteristics and questionable data can be  
158 adjusted to match the same distribution. The basic concept of the CDF matching calibration  
159 method is assuming the ambient atmospheric conditions are similar for all observation sites.  
160 In most field campaigns, the spatial distribution of upper-air radiosonde sites mostly satisfied  
161 such requirements, and hence, this method can efficiently adjust the data bias for most  
162 atmospheric conditions. However, such assumptions limit the generalizability of the CDF  
163 calibration models. Thus, the CDF models may not be directly applied to the data collected  
164 from different weather conditions, seasons, or climate regions with smaller sample sizes.

165           In this study, we focused on the calibration process of systematic error for ST  
166 temperature and moisture observations using the co-launch VS data. We use the co-launch  
167 data collected across several field campaigns in Taiwan to develop calibration methods for  
168 ST. Here, we proposed and evaluated two different calibration approaches. First, we  
169 followed the widely used CDF-matching approach and proposed a two-step CDF-based  
170 calibration scheme. Secondly, we incorporated the CDF-matching approach with modeling  
171 multivariate distributions, the central concept of machine learning, to introduce a novel  
172 correction method based on the generalized linear model (GLM). While the CDF approach  
173 discretized continuous variables, e.g., pressure and temperature, into bins to establish look-  
174 up tables, the machine-learning approach modeled a high-dimensional joint probability  
175 distribution with the same variables in their original forms. The latter approach allowed us to

176 compress complicated look-up tables into a unified mathematical representation. Hence, we  
177 can adjust the models more easily for better performance, robustness, and generalizability.

178 Section 2 describes the co-launched radiosonde data and the pre-processing.  
179 Section 3 focuses on the data correction algorithms, and data calibration processing flow.  
180 Section 4 summarizes the ST calibration results and compares them to the benchmark.  
181 Finally, Section 5 discusses the feature importance analysis and other calibration issues,  
182 and Section 6 presents the conclusions.

183

## 184 **2. Data and Preprocessing**

### 185 *2.1 Data Collection*

186 In the previous years since 2018, we have co-launched the ST with the Central  
187 Weather Administration (CWA) operational Vaisala RS41-SGP radiosonde (Figure 1c). The  
188 co-launch was conducted during field campaigns in the Taiwan area, including the Taipei  
189 Summer Storm Experiment (TASSE), the Yilan Experiment of Severe Rainfall (YESR2020),  
190 the Taiwan-Area Heavy Rain Observation and Prediction Experiment (TAHOPE), the  
191 Northern Coast Observation, Verification, and Investigation of Dynamics (NoCOVID21), and  
192 the Mountain Cloud Climatology (MCC) project. We collected 1,029 co-launches of ST and  
193 VS from these field campaigns during 2018–2022. These co-launches provided more than  
194 1,000,000 comparable independent observations of wind, pressure, temperature, and  
195 humidity (PTU) data. The co-launches of each campaign are summarized in Table 1, and

196 the geographic locations of the co-launch sites are shown in Figure 3.

197 In 2018 and 2019, based on the scientific goals of TASSE, we established a  
198 standardized procedure for the co-launches, and the observations were primarily conducted  
199 in the daytime. Once the observational procedure matured, we performed the day and night  
200 co-launches evenly in 2020, 2021, and 2022. (Table 2). Eventually, we collected 625  
201 daytime cases and 404 nighttime cases. Also, the pilot experiments were conducted in the  
202 summer, and in the latter field experiments, we performed the co-launches in other months.  
203 Though there were more cases in July and August, we still conducted at least 21 co-  
204 launches in May. As for the location, most co-launches were conducted at the Taipei weather  
205 station, while about 150 cases were in other cities in Taiwan. In these 1,029 co-launches,  
206 all STs successfully launched, and only 7 stopped sending signals after 300 seconds. The  
207 ST, designed with commercial hardware components, is reliable in field observations.

208 Note that the binding of ST and VS shown in Figure 1c differs from the instruments  
209 used in the Report of WMO's 2022 Upper-Air Instrument Intercomparison Campaign (IOM-  
210 143). The IOM-143 can be categorized into in-laboratory and in-field campaigns. The  
211 laboratory calibration techniques focus on understanding each instrument's characteristics  
212 regarding random errors, low-temperature performance, and solar radiation sensitivity. The  
213 field campaign calibration techniques emphasize ground checks. A major goal is to evaluate  
214 the observation difference between the radiosonde systems, including the VS.

215 The IOM-143 used a rig to hold multiple instruments together while avoiding

216 interferences from ventilation and signals. Accordingly, the simple binding in our study may  
217 increase the random difference between ST and VS. However, this study aims to develop  
218 correction methods for ST to behave as close to VS as possible. Our simple binding co-  
219 launches in a consistent manner for several years are the only data we have. As presented  
220 in the following session, the biases of the corrected ST observations are slightly larger than  
221 the random errors. Hence, we used a relatively simple binding design in the co-launches  
222 before 2023. Future binding co-launches will be conducted according to the WMO standard.

223

## 224 *2.2 Pre-processing of the co-launch data*

225 The ST is with the wind estimated from GPS. We analyzed the difference in wind  
226 variables with the paired data of VS and ST. The mean deviation in zonal and meridional  
227 wind components,  $u$  and  $v$ , are 0.04 and 0.03 ms<sup>-1</sup>, respectively. The difference may come  
228 from the time lag of GPS signals between two sensors, which is small enough to ignore. In  
229 this paper, we emphasize the correction of temperature and humidity calibration.

230 The co-launch's primary purpose is to understand ST's performance further and  
231 develop a data correction scheme to approximate the VS's observations. The raw data  
232 collected often contains inconsistencies, inaccuracies, and outliers that can significantly  
233 distort analytical results and impede the accuracy of predictive modeling. Therefore, we  
234 need a proper procedure to process the raw data.

235 In the work of Ciesielski et al. (2012), the authors suggested four stages for

236 developing research-quality radiosonde data (their figure 1). The first level requires a single  
237 unified data format. The second stage uses automated tools to remove unreliable data  
238 based on prior knowledge of quality control (QC) checks. Then, data biases are detected  
239 and corrected in the third level based on analysis or statistical methods. Finally, the fourth  
240 level dataset aims to be user-friendly, usually in uniform vertical resolution with QC flags.

241         Following the framework proposed by Ciesielski et al. (2012), our data correction  
242 method is applied in the third stage. Hence, we need a pre-processing scheme to derive a  
243 level 2 dataset from the raw co-launch data.

244         Figure 4 illustrates the preprocessing used in this study. In the first stage, we  
245 paired each ST and VS observation by nominal observation time and stored them in the  
246 same plain-text format, L1\_ST and L1\_VS. Then, in the second stage, we corrected known  
247 errors for both sensors, including missing values and outliers. After this stage, we derived  
248 the level 2 dataset, L2\_ST and L2\_VS. Finally, given the fact that both ST and VS  
249 radiosondes are attached during co-launch (as Figure 1c), we used “time after launch”  
250 (every second) in both profiles to pair the values of two sensors, and resulted in L2\_ST-VS.

251         Based on the prior studies of ST (Hwang et al., 2020), we performed a “ground  
252 check” procedure to correct the pressure values of ST. This procedure adjusts the  $P_{ST}$  by  
253 a constant bias  $dP_0$ , which is the difference between the surface pressure of the standard  
254 instrument and the sensor of ST. Furthermore, we filtered out profiles with inconsistent  
255 timestamps and paired records less than 250 (366 out of 1,029). Finally, we derived a

dataset of 663 merged profiles and 1,219,710 paired entries (up to every second) for further analysis.

258

### 259 **3. Data Correction Methods**

260 To develop a data correction scheme for ST, we first investigated the conventional  
261 CDF-based probability matching method (Ciesielski et al., 2009). Then, we extended this  
262 approach with direct modeling of multivariate distributions, which is the central concept of  
263 modern machine learning. We implemented the scheme with the basic generalized linear  
264 model (GLM) and compared the differences between the two approaches. Both CDF and  
265 GLM are simple statistical models. The CDF is based on a non-parametric approach, and  
266 the GLM is a parametric distribution (i.e., Gaussian distribution).

267 Before diving into the specific correction methods, we define the notations and  
268 symbols used in this study. While ST and VS represent the storm tracker and the Vaisala  
269 RS41-SGP radiosonde device, respectively, they are used as subscripts to denote the  
270 sensor of measurements. For example,  $P_{ST}$  means the pressure measured by ST, and  $T_{VS}$   
271 is the temperature recorded by VS. The  $\Delta$ (delta) symbol is used to denote the difference of  
272 the same variable between two sensors. Finally, the ' (prime) represents the corrected  
273 measure.

#### 274 **3.1 CDF-based Probability Matching**

275 CDF-based Probability matching, also known as histogram matching or quantile

mapping, is a statistical technique used to adjust the distribution of a dataset (e.g., a forecast distribution) to match that of another dataset (e.g., an observed distribution). The primary objective of this method is not to directly correct individual data points but to ensure that the overall statistical properties, such as the frequency of occurrence of specific values, match between the two datasets. In radiosonde observation, CDF-based probability matching is commonly used as a quality control tool to ensure data quality consistency for field campaigns (Nuret et al., 2008; Ciesielski et al., 2009).

Based on the paired entries collected in co-launches, the two-step correction scheme starts with correcting temperature ( $\Delta T$ ) based on the ground-checked pressure ( $P'_{ST}$ ) and the measured temperature ( $T_{ST}$ ). Then, the adjusted temperature ( $T'_{ST}$ ) is used together with the measured relative humidity ( $RH_{ST}$ ) to estimate the correction ( $\Delta RH$ ).

We first discretize the pressure and temperature variables in temperature correction into bins. Pressure is divided into 50 hPa intervals from 975–1025 hPa to 175–225 hPa, denoted by their centers, 1000 hPa to 200 hPa. The cumulative distribution function of temperature measured by ST and VS for each pressure bin is calculated as follows. The observed temperature records are sorted in ascending order, and then the proportion of observations is derived for every 1-degree interval from -80 to 40 degrees Celsius as the probability density. Based on the assumption that two sensors have the same CDF within this specific range, we derived the correction values,  $\Delta T$ , as a function of measured temperature,  $T_{ST}$ . Figure 5 demonstrates the CDF-based temperature correction



296 of the pressure bin 475–525 hPa as an example. The upper panel shows the CDF of  $T_{VS}$   
297 and  $T_{ST}$ , and the lower panel illustrates the correction ( $\Delta T$ ) as a function of the observed  
298 temperature ( $T_{ST}$ ). We grouped the co-launches into daytime and night-time and performed  
299 the above procedure for each pressure bin. The results are shown in Figure 6, the complete  
300 temperature correction table used in this study.

301 As shown in Figure 6, the temperature sensor of ST consistently shows warm bias  
302 in all pressure bins, and the bias is stronger at high altitudes. The night-time warm bias  
303 exhibits similar patterns to the daytime but with a lower quantity.

304 The correction of relative humidity (RH) is derived in the same way as the  
305 temperature, except for the independent variables, which are the corrected temperature  
306 ( $T'_{ST}$ ) and the measured relative humidity ( $RH_{ST}$ ). The corrected temperature is discretized  
307 into 10-degree intervals from -65 to 35 degrees Celsius. The relative humidity values are  
308 then rounded to integers and form 1% intervals from 0 to 100. Like the temperature  
309 correction procedure, the correction value is derived based on the CDF probability matching  
310 as a function of RH within each temperature bin. Figure 7 illustrates the complete RH  
311 correction table used in this study. Figure 7 indicates that the ST shows dry-bias (wet-bias)  
312 in lower (higher) altitudes. ST is generally dryer during the daytime.

313 Using the correction tables shown in Figures 5 and 6, the temperature and relative  
314 humidity measured by ST are corrected and evaluated. Mathematically, this procedure can  
315 be expressed as:

$$\Delta T = T_{VS} - T_{ST} = f(P'_{ST}, T_{ST}, Day) \quad (1)$$

$$\Delta RH = RH_{VS} - RH_{ST} = f(T'_{ST}, RH_{ST}, Day) \quad (2)$$

, where Day is a binary variable representing the daytime or night-time, and f is the CDF-based probability matching. Because we first correct the temperature and then use the corrected temperature to correct the humidity, we call this approach a two-step CDF-base calibration.

### 3.2 Generalized Linear Model

Despite the robustness and ease of implementation of CDF-based probability matching, the discretization steps and the form of the look-up table limit its application. For example, the discretization of pressure and temperature is empirical. Though the resulting CDFs and correction tables look reasonable, it is hard to justify that this is the only way to split a continuous variable into bins. In other words, by focusing on matching the overall distribution, probability matching may overlook or alter some of the finer-scale details in the dataset. Furthermore, the look-up table makes adding extra independent variables more complicated. For example, we used daytime and night-time tables to simplify the influence of solar radiation so that we could use two tables for each correction. Another example is when we consider adding the effect of pressure in the correction of RH. In that case, we need to establish three-dimensional bins and justify whether the cut-off points are adequately selected. Therefore, we want to introduce the modeling of the multivariate probability distribution to our correction scheme.

336 In essence, modeling the joint probability distributions of multiple variables is  
337 fundamental in machine learning for capturing relationships and dependencies among  
338 numerous predictors. It forms the backbone for various algorithms and techniques to predict,  
339 generate, and understand multi-dimensional data. In equations (1) and (2), the mapping  
340 function,  $f$ , can be seen as a model of the joint probability distribution of the independent  
341 variables. While the CDF-based probability matching algorithm models this distribution by  
342 discretizing the independent variables, it can be replaced by different algorithms that keep  
343 the predictors in their continuous form.

344 The Generalized Linear Model (GLM, Nelder and Wedderburn, 1972) is a versatile  
345 statistical framework used for modeling the relationship between a dependent variable  
346 (response) and one or more independent variables (predictors) in a wide range of  
347 applications. GLMs extend the concept of linear regression to handle a broader array of data  
348 types and distributions. They are valuable for offering interpretable coefficients to  
349 understand the impact of predictors on the response. GLMs have become a fundamental  
350 tool in statistics and data analysis due to their flexibility and applicability across various fields.  
351 In this study, we used GLMs in three different settings: first, the same scheme as CDF-  
352 based probability matching (GLM1, as specified in equations (1) and (2)); second, using the  
353 same set of predictors for T and RH corrections (GLM2); and finally, replacing daytime with  
354 Julian-day and hour-of-day (GLM3).

355 To develop the GLM-based corrections, we used the paired entry dataset and the

least squared algorithm to fit linear regression models for the response variables ( $\Delta T$  and  $\Delta RH$ ) and the predictors ( $P'_{ST}$ ,  $T_{ST}$ ,  $RH_{ST}$ , and Day). This study used the Python algorithm implementation from scikit-learn (Pedregosa et al., 2011). The resulting regression equations are used to correct the storm tracker data.

In the second GLM configuration, we use the variables of  $P'_{ST}$ ,  $T_{ST}$ ,  $RH_{ST}$ , and Day to predict the corrections of temperature ( $\Delta T$ ) and relative humidity ( $\Delta RH$ ). The resulting models can be mathematically denoted as:

$$\Delta T = f(P'_{ST}, T_{ST}, RH_{ST}, Day) \quad (3)$$

$$\Delta RH = f(P'_{ST}, T_{ST}, RH_{ST}, Day) \quad (4)$$

Previous studies have suggested that solar radiation could be the leading cause of the warm bias in the radiosonde data. This is why we established correction tables for daytime and night-time separately. To simplify the correction process and limit the number of tables created, the solar radiation is represented by the binary variable of Day. However, with GLMs, we can easily use continuous variables in their original form. Hence, we used the “Julian day from the summer solstice” (Jday) and the “hour-of-day from noon” (Hour) to replace the Day variable. The resulting models are:

$$\Delta T = f(P'_{ST}, T_{ST}, RH_{ST}, Jday, Hour) \quad (5)$$

$$\Delta RH = f(P'_{ST}, T_{ST}, RH_{ST}, Jday, Hour) \quad (6)$$

These three settings are noted as GLM1, GLM2, and GLM3 in the later text.

Because all of our co-launches were conducted over the Taiwan area, the Julian

376 day and the hour of the day can properly approximate the value of the clear day radiation.  
377 Though the resulting correction formula can be applied to other regions, the differences in  
378 the pressure-altitude relationship might slightly interfere with other predictors. Therefore, we  
379 recommend adding the location information (i.e., longitude and latitude) or directly using the  
380 derived values of clear-sky radiation to develop the correction formula in other regions.

381

#### 382 **4. Results**

383 Figure 8 illustrates the patterns and deviations between ST and VS at various  
384 pressure levels. The panels (a), (b), and (c) demonstrate the temperature of VS and ST, and  
385 the differences between the two sensors. The relative humidity is shown in panels (d), (e),  
386 and (f). As shown in Figure 8, the ST exhibits warm and dry biases in general, and the biases  
387 increase as the altitude rises.

388 We applied the four correction methods described in the previous section, i.e.,  
389 CDF, GLM1, GLM2, and GLM3, to the 663 sounding profiles. Using the VS as the reference  
390 observations, we calculated the root-mean-squared errors (RMSEs) as the evaluation  
391 metrics. We did not use the correlation coefficients for evaluation because two sensors have  
392 correlation coefficients higher than 0.99, even without corrections. The reason for this lies in  
393 the co-launching strategy, which ensures that both instruments endure the same  
394 environmental conditions. The means and standard deviations of RMSEs for all correction  
395 methods are shown in Table 3 and Figure 9. As shown in Figure 9, we can see a significant

396 bias reduction for all correction methods. We performed t-tests on the raw and corrected  
397 values, and the improvement of all four methods is statistically significant (for p-values little  
398 than  $10e-29$ ). We also compared the CDF and GLM, and the results show that CDF  
399 correction is slightly better than GLMs for both temperature and relative humidity. The  
400 difference between CDF and GLMs is significant in the t-test, though the significant level is  
401 much lower than their bias reduction.

402 We also conducted t-tests on different GLM settings. The GLM1 and GLM2 did  
403 not show significant differences in temperature and relative humidity correction results.  
404 However, the GLM3 showed great improvement compared to GLM1 and GLM2. This  
405 suggested that solar radiation parameters can influence the correction more than a simple  
406 day/night indicator.

407 Table 3 and Figure 9 also show the evaluations for all records below 500- and  
408 700-hPa heights. As shown in the results, ST can approximate the VS measurements with a  
409 temperature error of less than 1 degree Kelvin and a relative humidity error of less than 10%.  
410 Suppose we focus on the observations below 700 hPa. In that case, the averaged RMSE  
411 can be as low as 0.66-degree Kelvin for temperature and 4.61% for relative humidity,  
412 comparable to the uncertainties of VS temperature and relative humidity measurements  
413 (Vaisala, 2017). Such results suggested that the ST is sufficiently accurate, especially when  
414 focusing on the boundary layer and lower atmosphere.

415 In addition to the overall performance of ST, we illustrated the RMSEs distribution

of the 663 soundings in Figure 10. The upper panel, (a), illustrates the distribution of RMSEs before correction, and the lower panel, (b), shows the results after the CDF-based correction. As shown in Figure 10, the proposed correction methods reduced both the biases and spreads. The reduction in the standard deviation of RMSE in Table 3 also represents this fact. Based on Figure 10, we selected three cases with low, middle, and high biases in RH before correction to discuss in the following section. The one-by-one comparison of the 633 profiles can be found in the supporting materials.

423

## 424 **5. Discussion**

### 425 *5.1 The Random Errors of ST*

426 The specifications of the temperature and humidity sensor used in the ST reported  
427 the accuracy range as  $\pm 0.3^{\circ}\text{C}$  and  $\pm 2\%$  (Huang et al., 2020). We examined the random  
428 errors with cloud chamber laboratory examination and field observation datasets with dual  
429 ST launching.

430 Six STs of the same batch used in the co-launches were measured in controlled  
431 chambers. Each sensor was repeatedly measured at  $10^{\circ}\text{C}$ ,  $20^{\circ}\text{C}$ ,  $30^{\circ}\text{C}$ , and  $40^{\circ}\text{C}$ , and at  
432 relative humidity of 30%, 50%, 70%, and 90%. The results are shown in Figure 11. The  
433 standard deviations of the measured differences are  $0.24^{\circ}\text{C}$  (temperature) and 2.21%  
434 (relative humidity), respectively. The results reasonably agreed with the random errors  
435 reported by the manufacturer.

436           To assess the random error in the field, we conducted 42 observations with dual-  
437 ST. We aligned the records of two instruments with timestamps and evaluated the  
438 differences in temperature and humidity. In the 42 launches, there were a total of 96,284  
439 aligned entries. We used statistical fences to exclude extreme situations such as frozen or  
440 malfunctioning sensors (Everitt and Skrondal, 2010; Tukey, 1977). After applying this simple  
441 outlier removal technique, we have 85,641 temperature measurements and 81,616 pairs of  
442 relative humidity. The paired measurements are shown in Figure 12. The derived standard  
443 deviation for temperature is 0.52°C, and for relative humidity is 2.25%. The random errors  
444 measured in the field are slightly higher than those measured in the laboratory and reported  
445 by the manufacturer. There were 42 dual-ST attached to VS co-launches in the field, all of  
446 them were conducted during the day. We realized that the sensor performance could have  
447 diurnal variations and we had performed the correction according to the day-night difference.  
448 By following the types of errors defined in Collins (2001), we attributed the day-night  
449 variability as systematic error, which our correction methods can remedy. Hence, we didn't  
450 further distinguish the random errors for day and night.

451           The results of the sounding-by-sounding evaluation presented in the earlier  
452 section, 0.66-degree for temperature and 4.61% for humidity, are slightly larger than the  
453 random errors measured in the field. This suggests that there is room to develop more  
454 sophisticated correction methods.

455           According to previous studies, the ST sensor has a about 5-second response time



456 (Huang et al., 2020). Several time-lag analyses were conducted to verify this and the impact  
457 to the measurement correction, the results suggest insignificant changes to the bias  
458 correction. However, given that Miloshevich et al. (2001, 2004) discussed the errors  
459 introduced by the sensor's time lag and proposed a correction algorithm, we plan to  
460 incorporate further sophisticated time-lag correction approaches in the future.

## 461 *5.2 General performance of ST*

462 Figure 13 illustrates the paired entries of VS and ST before and after corrections.  
463 As described in the previous section, the ST exhibits correlation coefficients higher than 0.99  
464 for temperature and RH even before any correction. Hence, the effect of corrections is  
465 represented by the narrower diagonals in the right panels in Figure 13.

466 Even though the statistical tests showed the significance of the correction results,  
467 they are not easily perceived. Hence, we selected a few sounding profiles to demonstrate  
468 the effectiveness of our correction methods. Figure 14 shows the T and RH profile of the  
469 sounding launched at 2021-08-03 12Z. This sounding was selected because of the overall  
470 low RH bias before and after correction. In Figure 14, the corrected temperature is  
471 adequately aligned to the reference ( $T_{VS}$ ), and the corrected relative humidity (RH) is entirely  
472 satisfactory, particularly below 350hPa, covering most tropospheric levels with water vapor  
473 and clouds. Consistent findings are prevalent within our dataset, indicating that the adjusted  
474 ST measurements are reliable across various observational scenarios.

475 However, the corrected results may perform less when encountering extreme wet

476 cases. Figure 15 is the sounding profile on 2018-08-27 06Z when the reference RH of VS is  
477 about 90% from ~850 to ~350-hPa heights. As shown in Figure 15, the temperature  
478 correction still works properly, except that the VS's temperature sensor showed much larger  
479 amplitude compared to VS. However, the RH measured by ST shows a dry bias of  
480 magnitude of 20% from ~850 to ~350-hPa heights while the patterns stay similar. The RH  
481 correction mechanisms adjust the RH toward the reference, but the deviations are still  
482 significant. Note that this observation occurred during a severe rainfall event caused by the  
483 convergence of the tropical depression and the southwest monsoon from August 23 to  
484 August 30, 2018. All fifteen co-launches conducted in this event exhibited high bias in RH,  
485 ranging from 10% to 24%, and five showed bias greater than 10% even after correction.  
486 This particularly biased case has RMSE ranked 99.93% in our dataset. Since such a large  
487 deviation rarely showed in the colaunches, we believe it could be caused by malfunction of  
488 this specific sensor.

489 In the left panel of Figure 15, we can also see a sudden change in GLM-corrected  
490 temperature around 310hPa. This should be caused by the missing values of ST in RH (see  
491 the missing yellow line section in the right panel). Because the GLM correction includes RH  
492 as an independent variable, when RH values are missing (treated as 0), the amount of  
493 correction can change accordingly.

494 Figure 16 illustrates the sounding profile on 2020-03-13 12Z. This is an average  
495 case with middle bias in RH before correction. Most of the 633 co-launches behave similarly

496 to this case.

497 From the cases shown above, we also notice the characteristics of different  
498 correction methods. The GLM adjustments look like horizontal shifts of the original values  
499 due to the linearity of the model.

500 Despite the simplicity of our correction methods, the temperature bias between  
501 ST and VS can be reduced from 3.0 K to 0.9 K, and the RH bias from 8.5% to 6.9%. Note  
502 that our correction methods also reduce the standard deviations from 1.8 K to 0.6 K and  
503 3.8% to 2.8%, respectively. Hence, we can expect 80% of ST observations to exhibit less  
504 than 1K bias in temperature and 8.8% bias in RH.

505 The corrected ST measurements aligned well with the VS data, especially when  
506 the sounding successfully reached an altitude higher than 300hPa. For those co-launches  
507 that ended early, though their bias is still low in statistics, their profiles usually looked  
508 problematic when visualized. We recommend further looking into the reasons that cause the  
509 sounding to end early.

### 510 *5.3 A ST observation in afternoon thunderstorm study*

511 The low cost of the ST can facilitate high spatial-temporal frequency of upper-air  
512 observations. While the ST provides reasonable measures after correction, its reliability in  
513 higher altitudes is still incompatible with the VS used in standard operation. Therefore, here,  
514 we demonstrate a use case to illustrate the strength of the ST. Figure 17 shows a set of  
515 continuous ST profiles on 2018-08-17 with one-hour intervals. This experiment used only

ST and was not included in the colaunch dataset. Figure 17 shows the evolution of a local convective system, which is not feasible in regular 12-hour interval radiosonde operation - the increase of atmospheric moisture at 1300 local time before the heavy rain occurrence is observed. Using the flexibility in deploying the ST during field campaigns allows us to capture vertical profiles in the lower troposphere at an hourly, or even a shorter time interval. This is notably advantageous for understanding the development of deep convection, which typically has a lifetime of 1 to 3 hours, and the surrounding environment, especially the lower boundary layer. A similar ST profile has been used in the study of the afternoon thunderstorm in Taipei compared to the results from CRESS cloud-resolving modeling (Tsujino et al. 2022). Note that the ST data here was corrected with the CDF-based method; better performance can be achieved with GLM-based methods.

527

## 528 **6. Concluding Remarks**

529 In this study, we assess the data quality control and calibration of the Storm  
530 Tracker (ST) with the co-launched Vaisala RS41-SGP (VS) in temperature, relative humidity,  
531 and winds for lower atmospheric observations. Although wind speed and direction are  
532 crucial information in radiosonde observation, we found from the co-launched data that the  
533 GPS-estimated ST wind differs from that of VS in insignificant magnitude. The GPS  
534 estimated ST wind error difference is about  $0.05 \text{ ms}^{-1}$ . To ensure the reliability of ST  
535 measurements in temperature and moisture, we conducted over a thousand co-launches of

the ST and the VS, evaluating and refining the performance of the ST through developed correction methods for temperature and humidity measurements. Based on the sounding-by-sounding comparison, the corrected ST soundings have a 1-K temperature and 7% relative humidity root mean square difference from the VS soundings. These error differences can be reduced to 0.66-K and 4.61% below the 700-hPa height. The biases of the corrected ST observations are slightly larger than the random errors, which were 0.24 K and 2.21% in the laboratory and 0.52 K and 2.23% in the field.

Derived from the co-launch dataset, two correction methods based on CDF and GLM algorithms were implemented to enhance the quality of temperature and humidity observations in the ST. Both methods work comparably well to reduce the biases of the ST. While the CDF-based correction is robust and reliable, the GLMs easily model and change the predictors. The ST observations closely aligned with the VS after corrections, particularly in the lower atmospheric layers below 700hPa. For synoptic weather, geostrophic adjustment dynamics suggest that spatial temperature variations in the free atmosphere may not be significant, reducing the need for high-frequency upper-air radiosonde observations. Consequently, most operational radiosonde observations worldwide are conducted daily at 00Z and 12Z, with 12–24 hours intervals. However, atmospheric phenomena originating from the boundary layer are often smaller in scale and closely related to local terrain. For example, a single convective cell typically lasts minutes, while thunderstorms persist for a few hours. To better understand these types of weather, a low-

556 cost and lightweight device capable of deploying multiple sensors simultaneously or at  
557 intervals of less than an hour can enhance field experiments. This approach provides  
558 valuable insights into the lower atmosphere's significant variations in temperature and  
559 moisture, especially for convective systems that may lead to disastrous rainfall or flash  
560 flooding. This positions the ST as a promising candidate for supplementing regular upper-  
561 air observations for high spatial and temporal resolution in the lower atmosphere. Our work  
562 also demonstrated that low-cost commercial sensor components can help high-frequency  
563 observations in specific targets with carefully developed correction methods.

564         Although we used the linear regression version of GLMs in this study, the concept  
565 of modeling the joint probability distribution can be extended to various statistical models  
566 such as decision trees, support vector machines (SVM), and artificial neural networks (ANN).  
567 The simple GLMs in this study assume the response is a Gaussian distribution of the linear  
568 combination of predictors. Other machine learning models can establish nonlinear mappings  
569 between the predictors and response without assuming any distributions. However,  
570 investigating more machine learning models is beyond the scope of this study.

571         In summary, while the VS remains the standard for upper-air observation, ST is  
572 suitable for Planetary Boundary Layer (PBL) or lower atmosphere studies in areas with  
573 complex terrain. The ST can complement the VS observation with high spatial and temporal  
574 resolution observation of the lower atmosphere. This may be useful for mesoscale storm  
575 observations in East Asia, where PBL conditions can vary significantly within short distances.

576 However, it is important to note that the correction results presented here are specific to  
577 Taiwan's observation. This is especially true for the CDF method, as the variability of the  
578 CDF method data is height-dependent, so the direct use of our CDF calibration should be  
579 cautious. On the other hand, the GLM method may provide a reasonable calibration to the  
580 ST sounding when longitude and latitude are used as predictors or local clear-sky radiation  
581 is directly used. To ensure broader applicability, we suggest conducting co-launches during  
582 field campaigns. This approach would allow users to derive in-situ correction formulas using  
583 the proposed methods. Our experiments indicate that ST between VS launches may  
584 enhance meteorological data collection and analysis in the lower atmosphere.

585

#### 586 **Data Availability Statement**

587 The data for this project is confidential but may be obtained with Data Use Agreements with  
588 the National Taiwan University. Researchers interested in access to the data may contact  
589 authors. It can take some months to negotiate data use agreements and gain access to the  
590 data. The author will assist with reasonable replication attempts for two years following  
591 publication.

592 Code for data cleaning and analysis is provided in a replication package. It is available at  
593 [https://www.dropbox.com/scl/fo/ah7i6z4f7u2yzijfh7ua3/h?rlkey=ar4g2hq7hwkop2eyzw83el](https://www.dropbox.com/scl/fo/ah7i6z4f7u2yzijfh7ua3/h?rlkey=ar4g2hq7hwkop2eyzw83el8ih&dl=0)  
594 [8ih&dl=0](https://www.dropbox.com/scl/fo/ah7i6z4f7u2yzijfh7ua3/h?rlkey=ar4g2hq7hwkop2eyzw83el8ih&dl=0) for review. It will be uploaded to GitHub once the paper has been conditionally  
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597

598

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615

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734

## List of Figures

735

736 Fig. 1 (a) The Vaisala RS41-SGP radiosonde (weighted 84g, body dimension: 155 x 63 x  
737 46 mm), (b) the storm tracker mini-radiosonde (weighted 20 g with battery, body  
738 dimension: 70 x 29 x 18 mm), and (c) an example of the co-launched soundings via the  
739 TASSE experiment. More ST hardware details are described in Hwang et al. (2020).

740 Fig. 2 The sounding of 2018-06-26 03:00 UTC (11:00 LST) by VS (solid lines) and ST  
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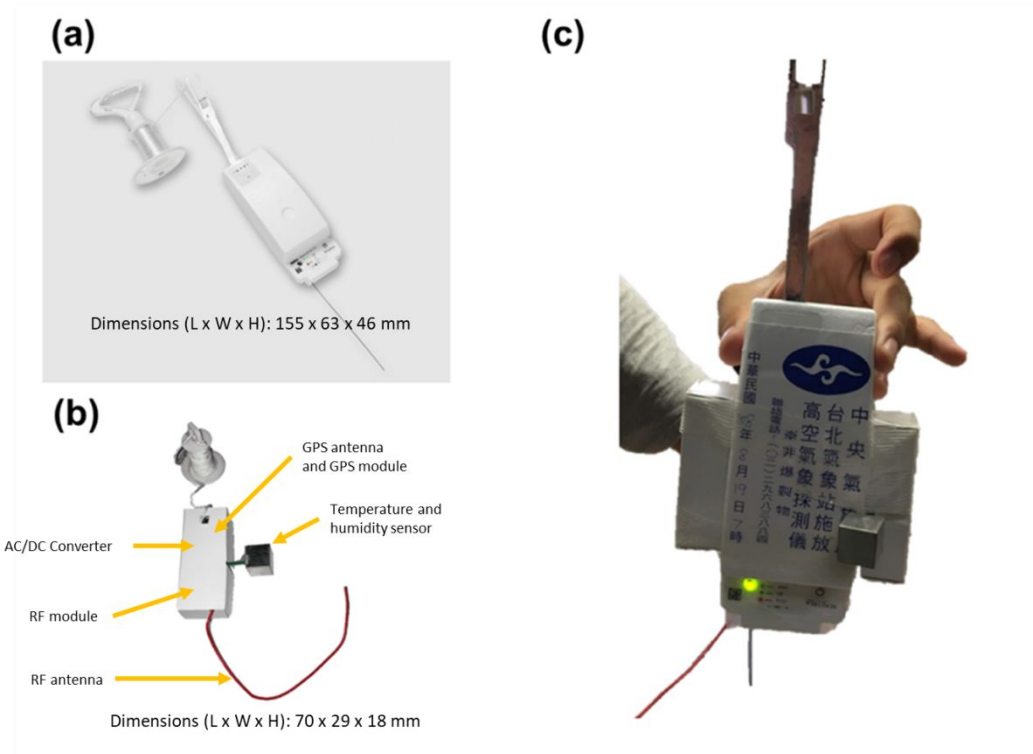
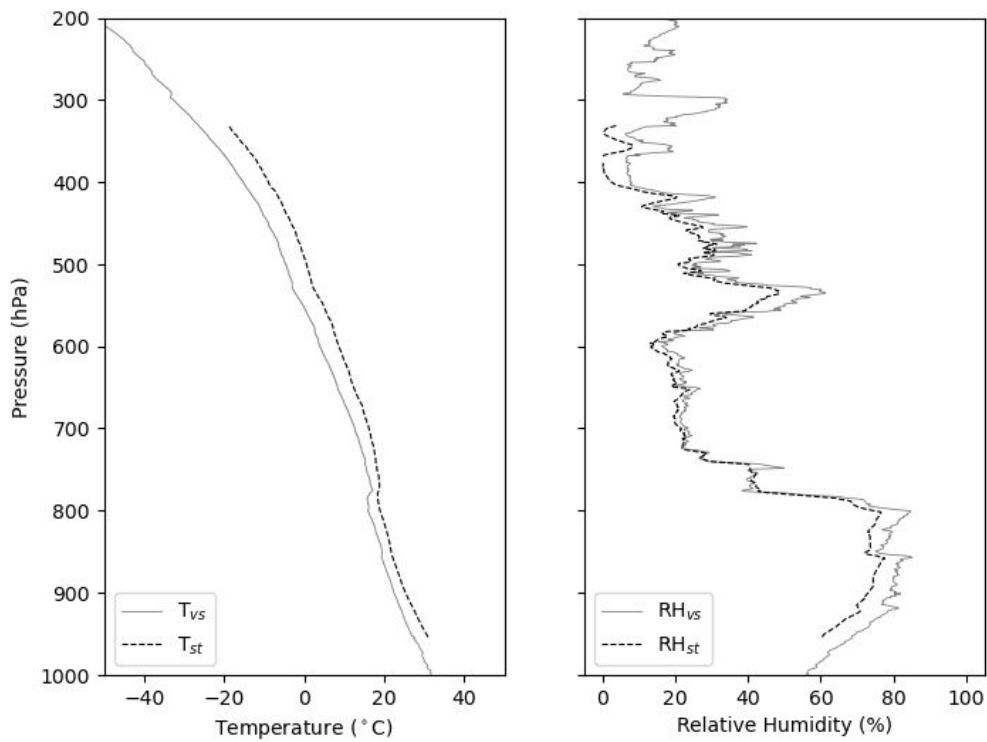


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2018-06-26 03Z(UTC)



795

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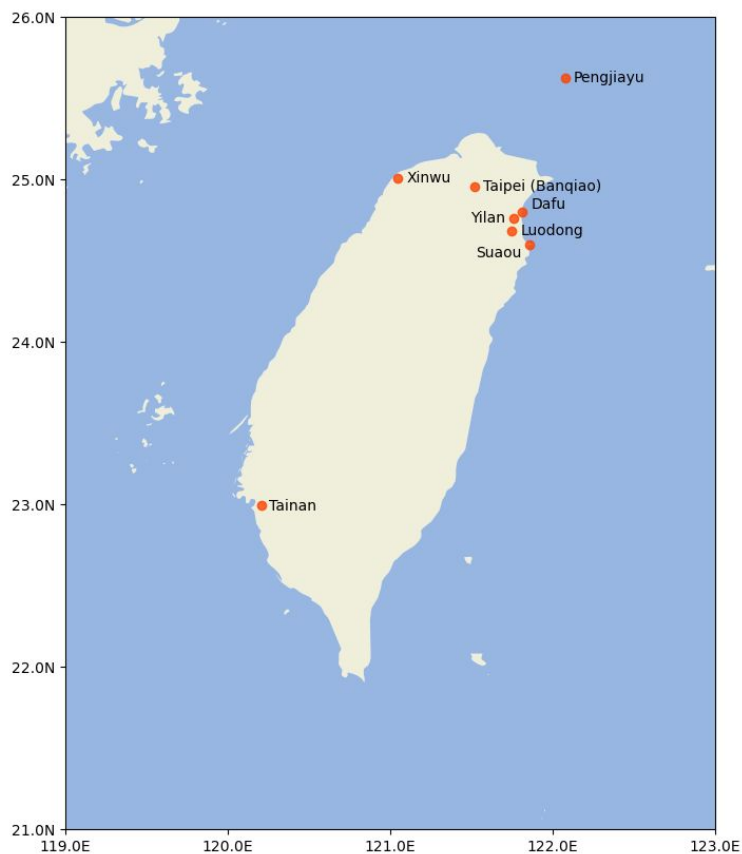
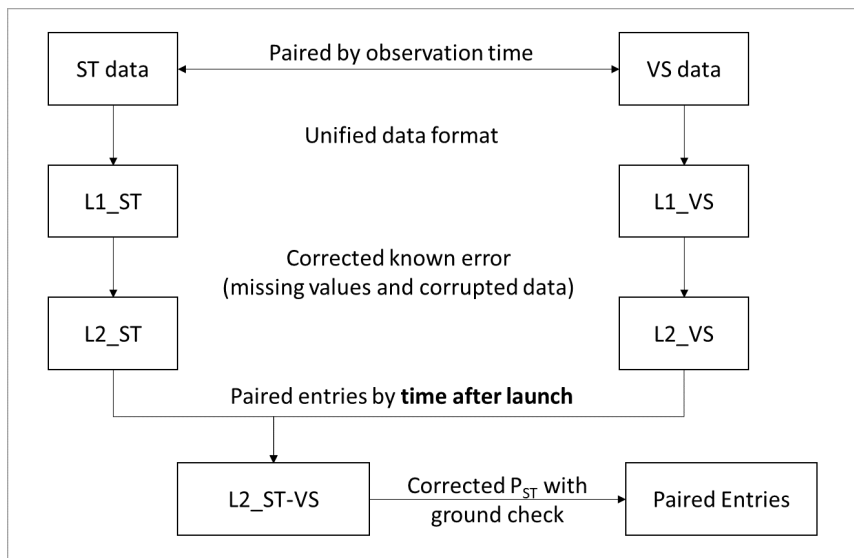


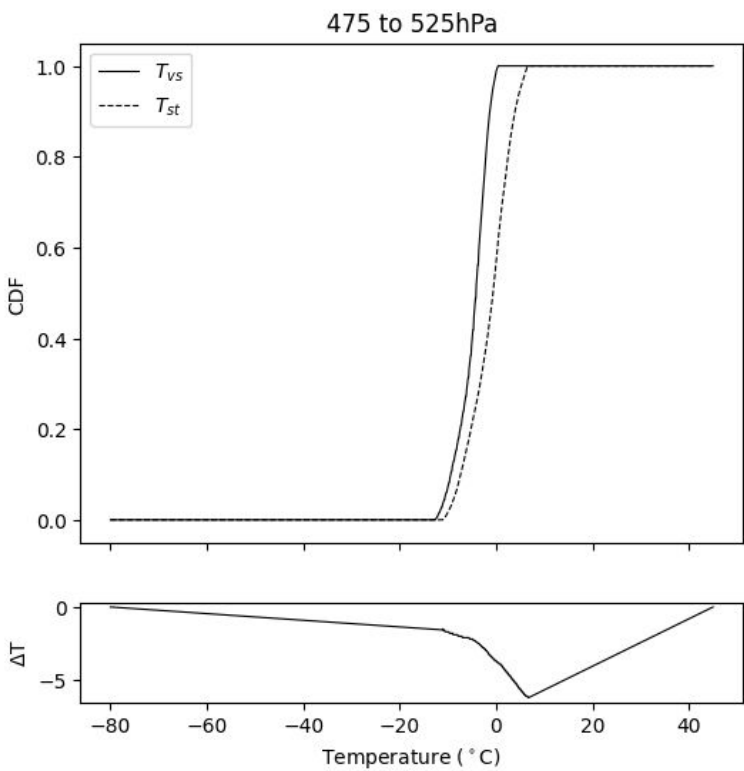
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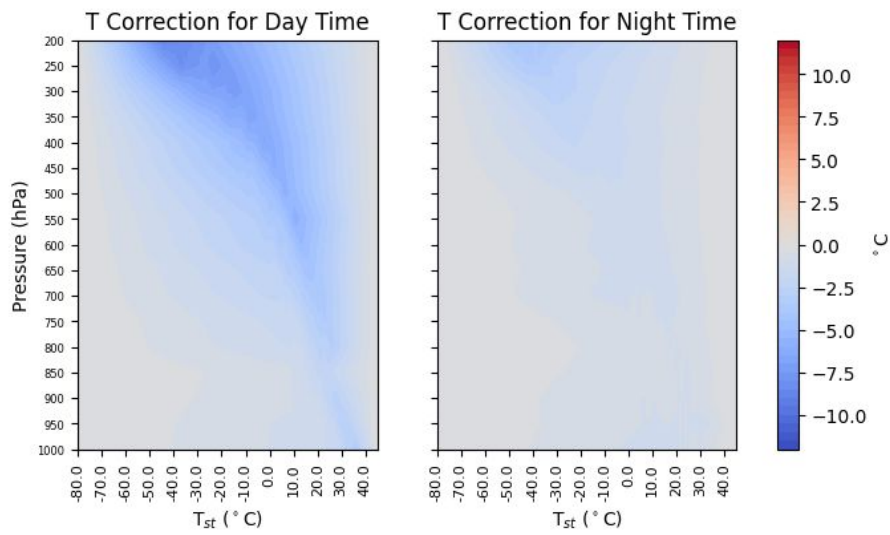
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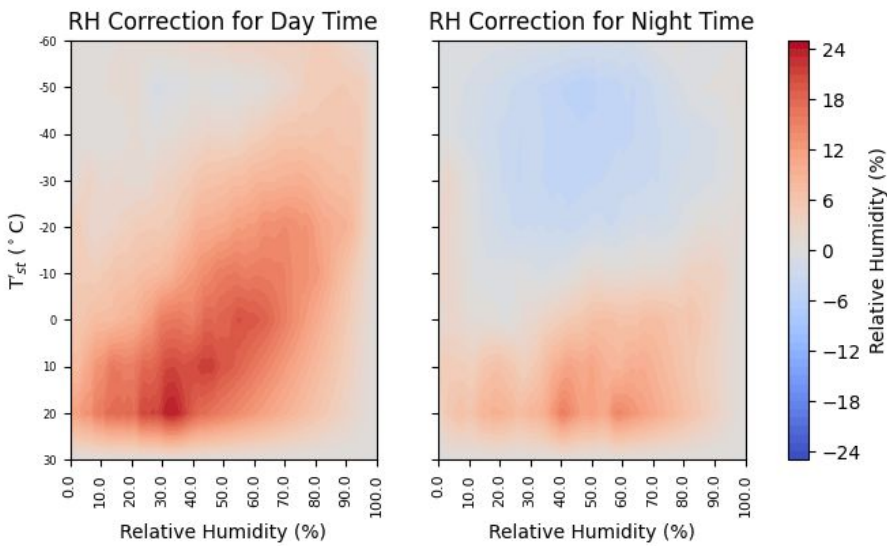


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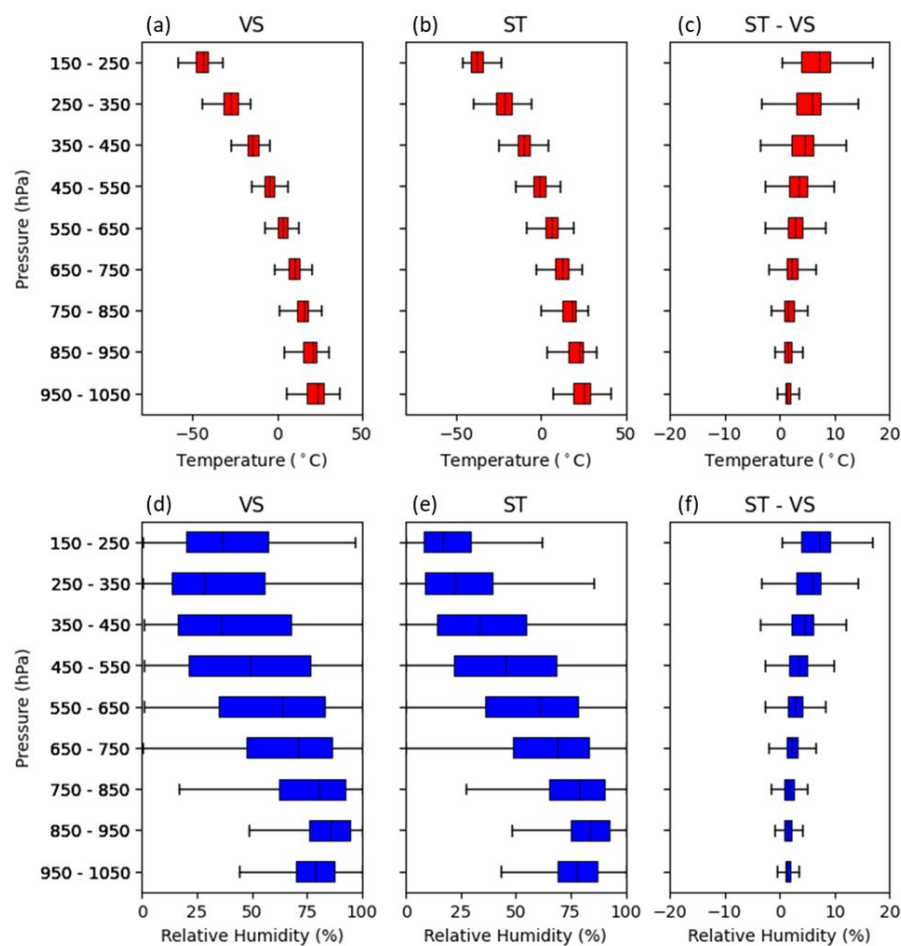


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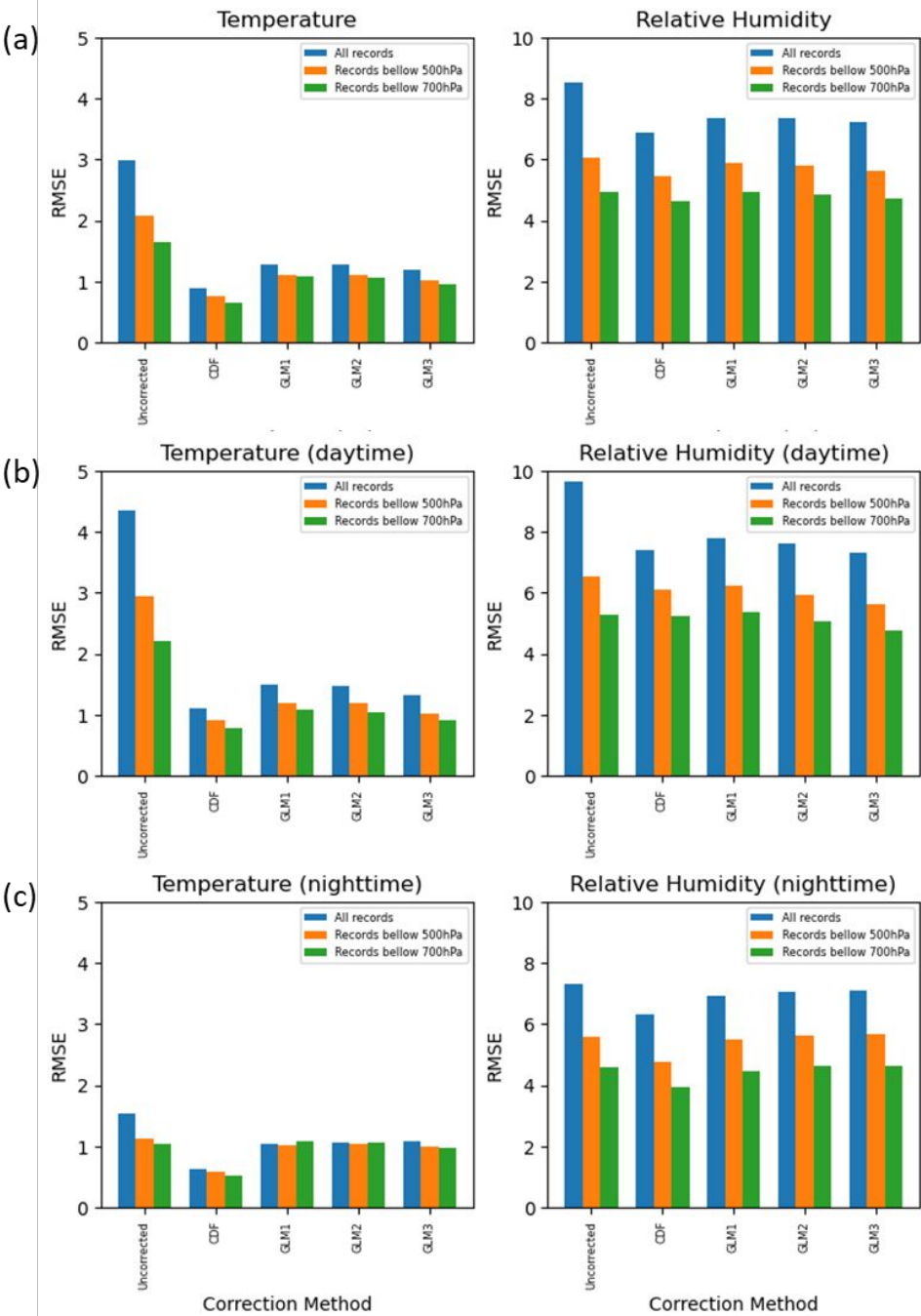
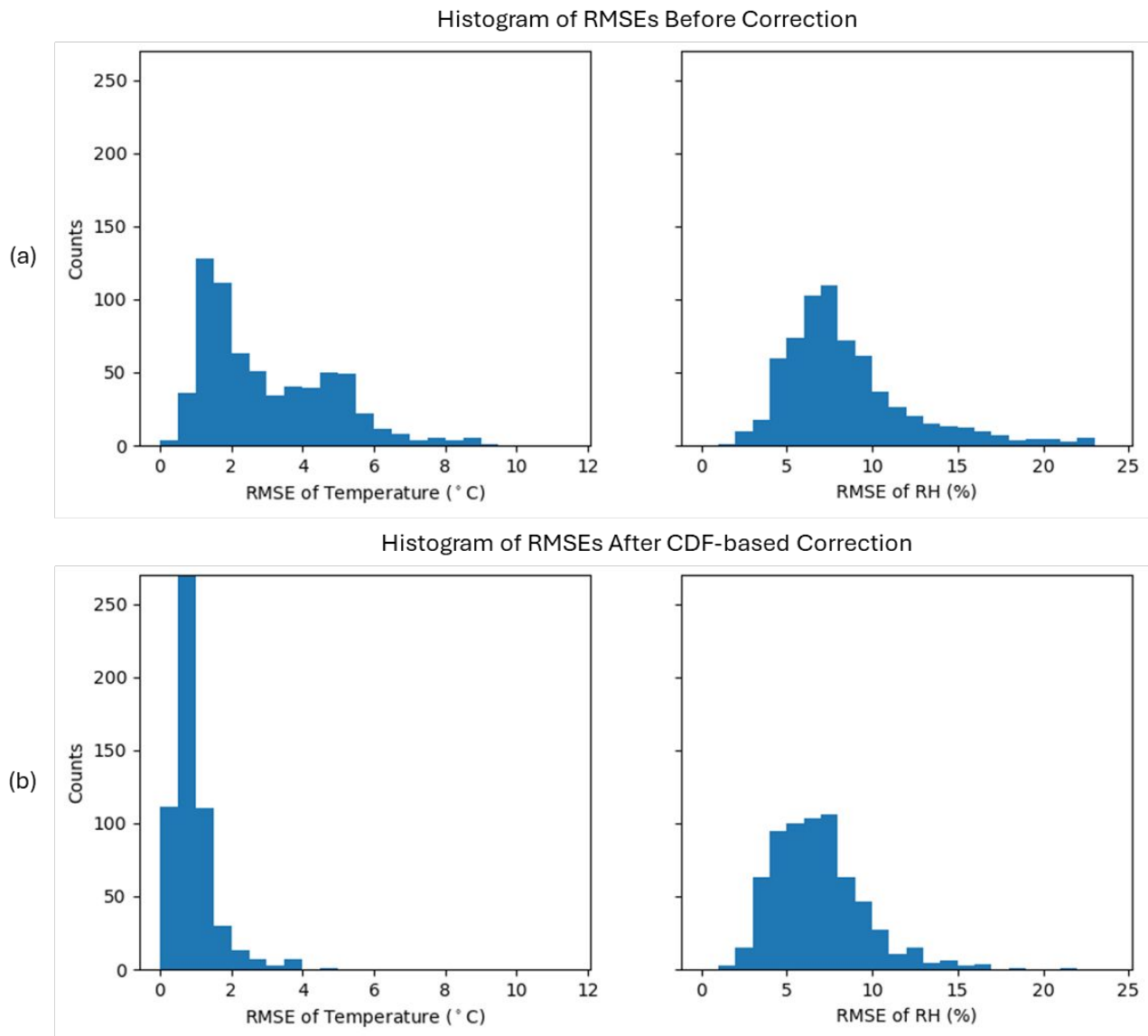
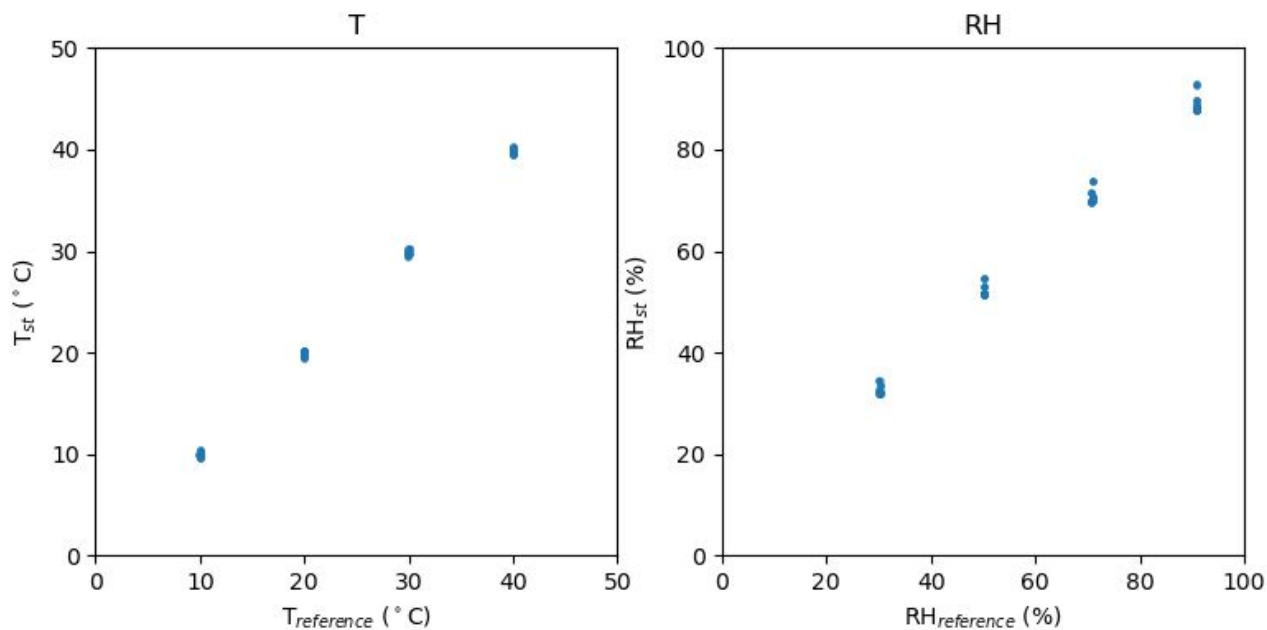


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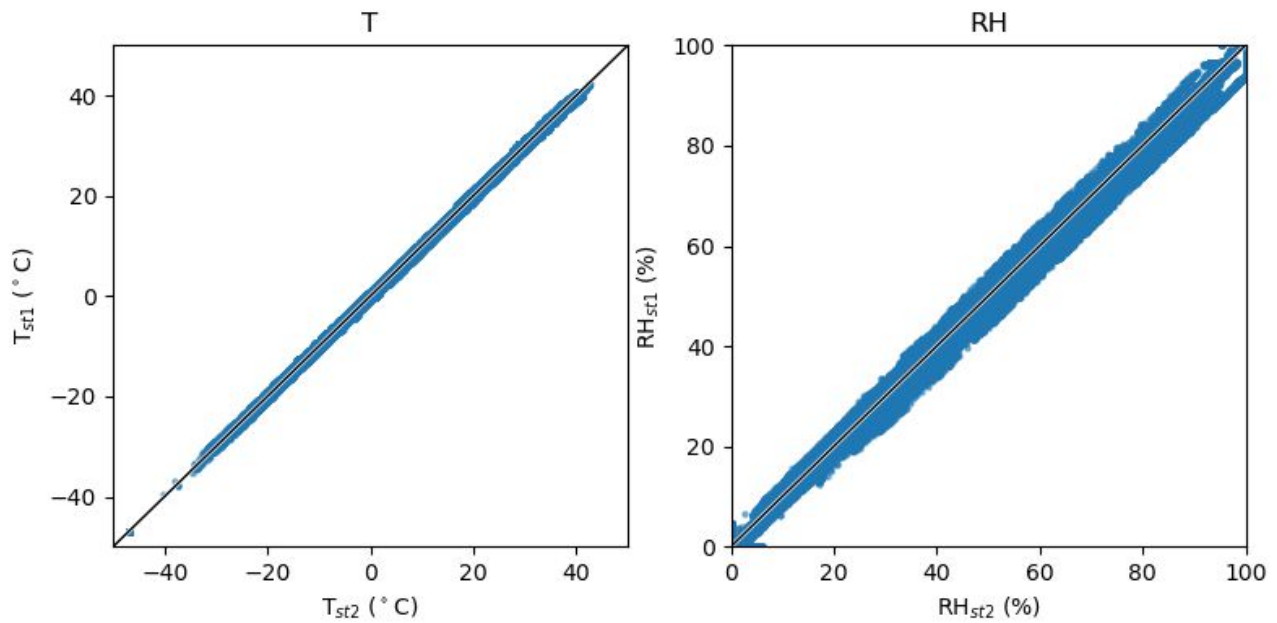
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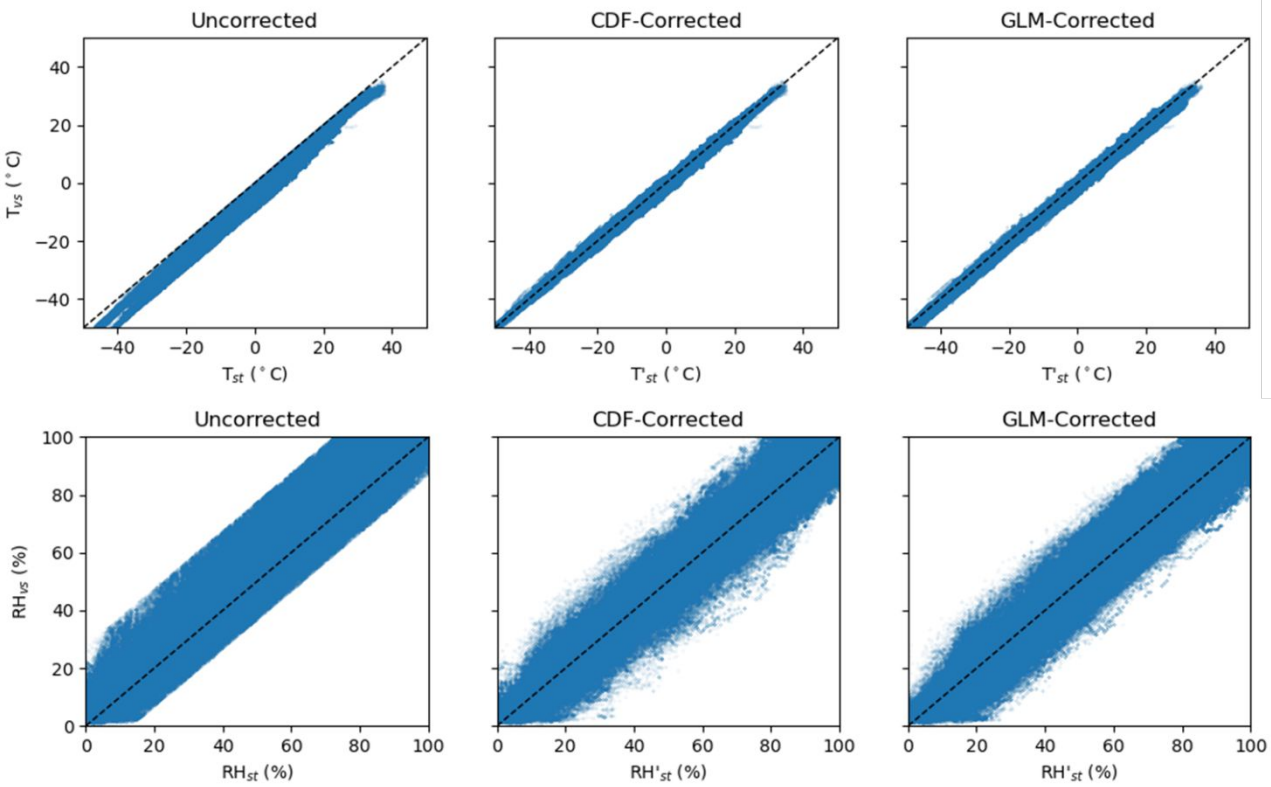


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2021-08-03 12Z

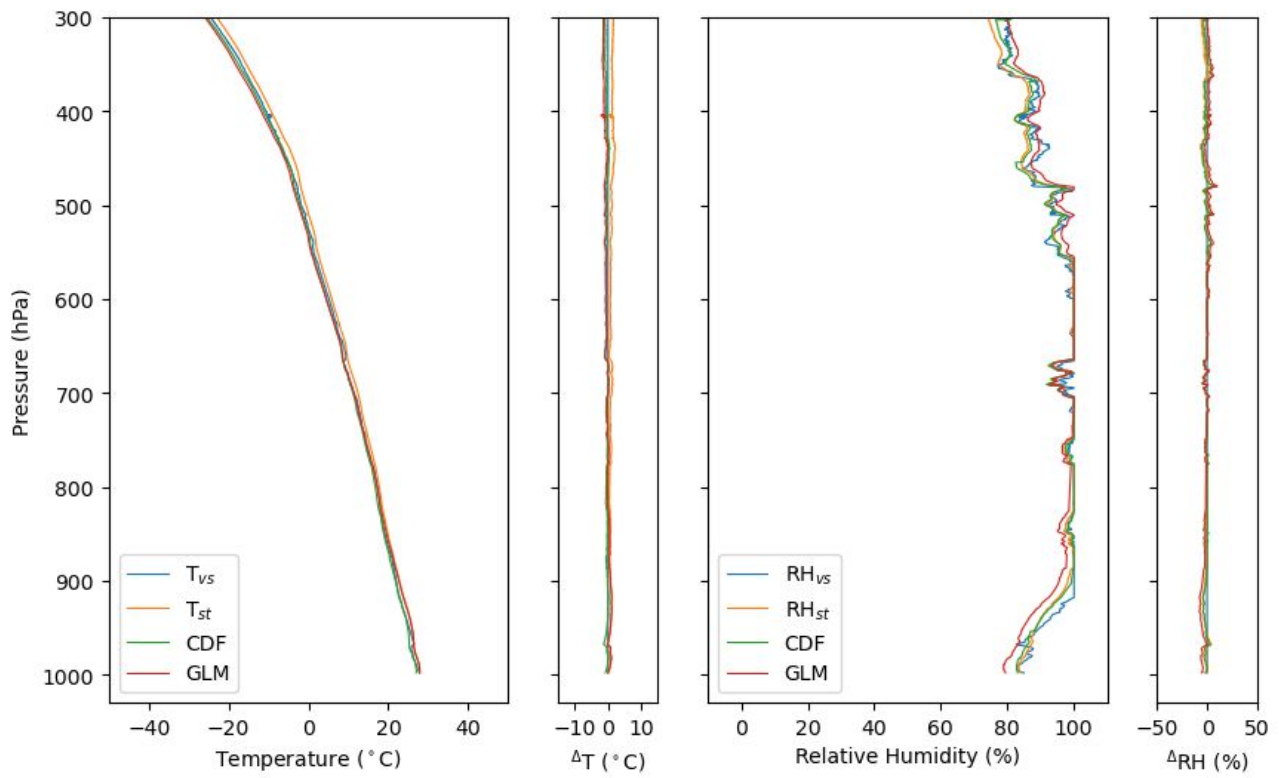
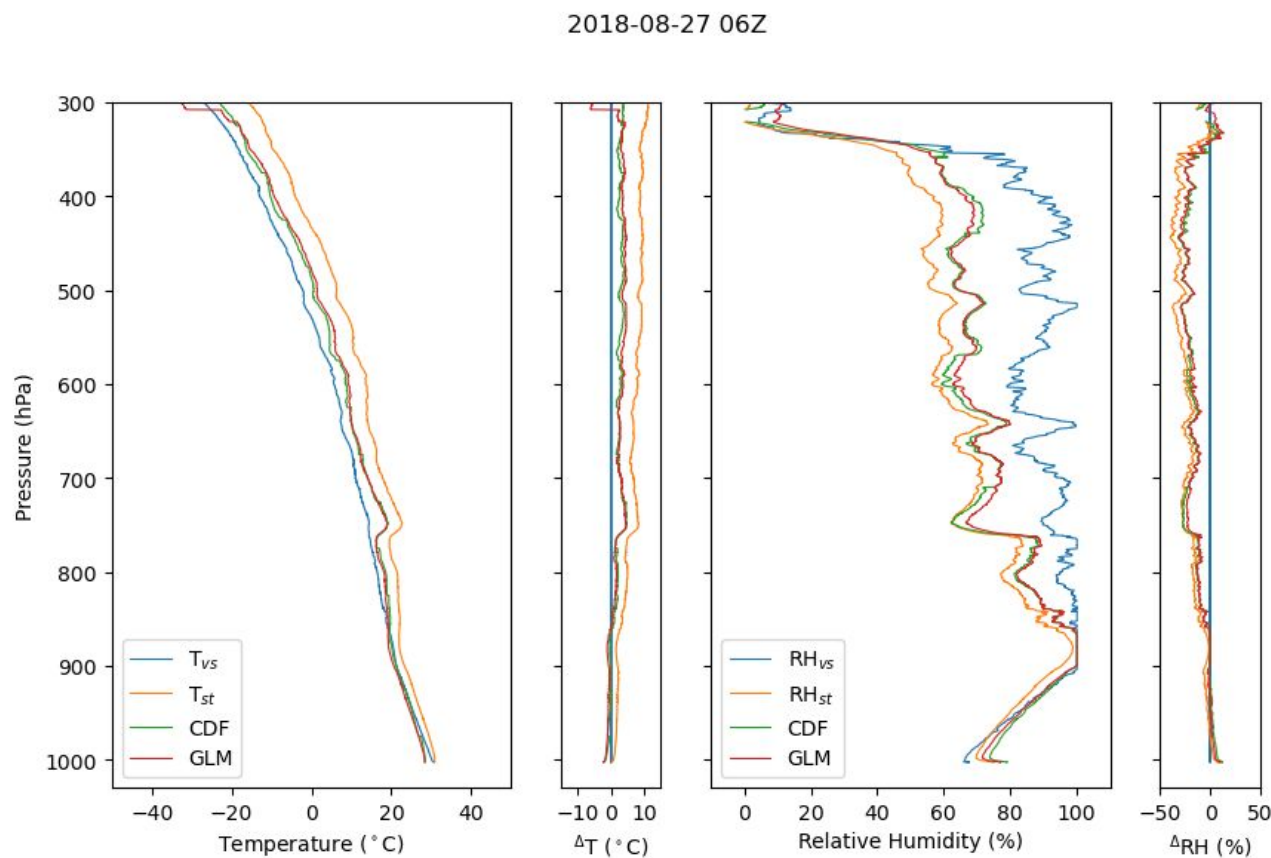


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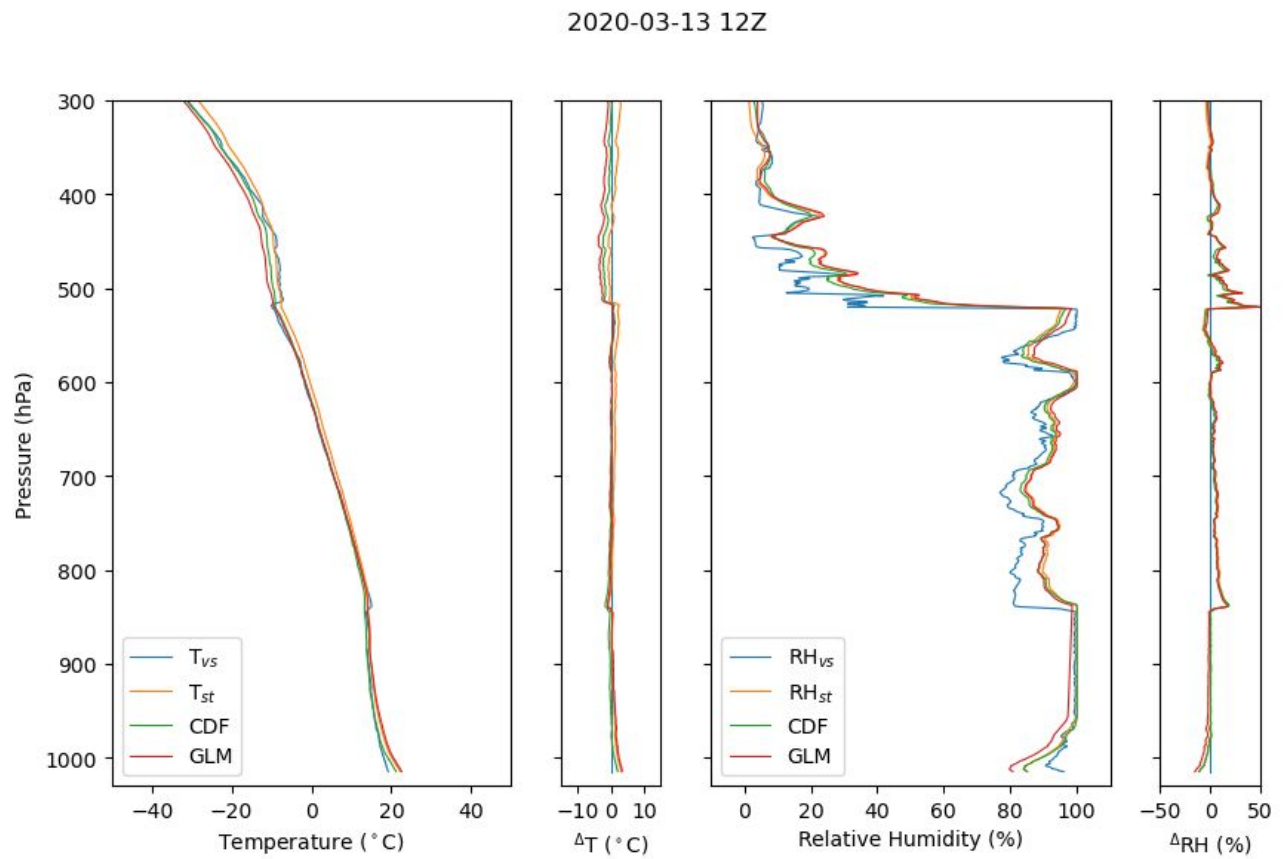
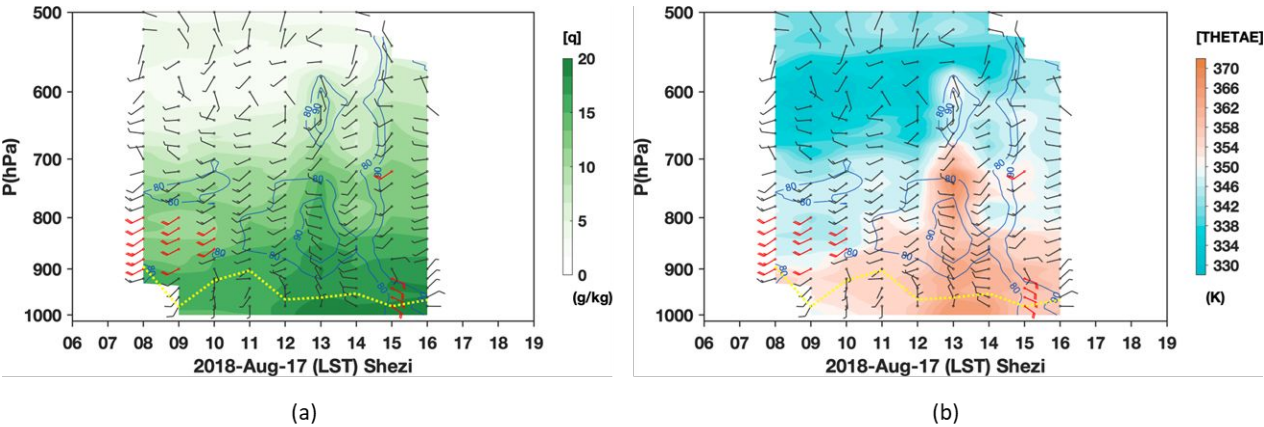


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874 (a) with the wind field. The derived equivalent potential temperature,  $\Theta_e$ , is shown in panel  
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877

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882 RH.

883

884    **Table 1    The summary of the field experiments conducting ST-VS co-launches.**

Experiment	Time	Location	Total Numbers of ST-VS Co-launch
Taipei Summer Storm Experiment ( <b>TASSE</b> )	2018-2020	Taipei (Banqiao)	478
Yilan Experiment of Severe Rainfall ( <b>YESR2020</b> )	2020. Nov	Yilan, Suaou, Luodong, Dafu	46
Taiwan-Area Heavy rain Observation and Prediction Experiment ( <b>TAHOPE</b> )	2019-2022	Taipei (Banqiao), Pengjiayu	382
Northern Coast Observation, Verification, and Investigation of Dynamics ( <b>NoCOVID21</b> )	2021. May-Jun	Taipei (Banqiao)	49
Mountain Cloud Climatology ( <b>MCC</b> )	2022. Oct-Nov	Suaou	23
Other		Tainan, Xinwu	51

885

886

887 Table 2 The summary of the 1,029 co-launches.

Month	2018		2019		2020		2021		2022		Total	
	Day	Night	Day	Night	Day	Night	Day	Night	Day	Night	Day	Night
1					21	24			25	26	46	50
2					29	29					29	29
3					27	31					27	31
4					30	30	15	13			45	43
5					6	5	6	4			12	9
6	14		20				30	32			64	32
7	14		60	12			22	23			96	35
8	41		85				23	22			149	22
9							25	26			25	26
10							29	28	7	3	36	31
11					20	17	40	41	6	7	66	65
12							30	31			30	31
Total	69	0	165	12	133	136	220	220	38	36	625	404

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889

890    Table 3    The RMSE of ST and VS with different correction methods for temperature and  
891    RH.

Variable	Correction Method	mean RMSE			stdev of RMSE		
		full	500hPa	700hPa	full	500hPa	700hPa
Temperature	Uncorrected	2.9969	2.0753	1.6446	1.8399	1.2291	0.8894
	CDF	0.8778	0.7568	0.6560	0.5579	0.4166	0.3367
	GLM1	1.2714	1.1126	1.0732	0.6612	0.4549	0.3682
	GLM2	1.2745	1.1128	1.0533	0.6625	0.4633	0.3693
	GLM3	1.1991	1.0105	0.9483	0.6284	0.4566	0.3579
Relative Humidity	Uncorrected	8.5265	6.0721	4.9336	3.8236	2.9284	2.3624
	CDF	6.8946	5.4707	4.6098	2.8107	2.7488	2.4442
	GLM1	7.4604	5.8673	4.9267	2.9158	2.6489	2.3084
	GLM2	7.4152	5.7997	4.8478	2.7785	2.4307	2.0590
	GLM3	7.2683	5.6355	4.7043	2.6668	2.3372	1.9878

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